

## Chapter 2 - The Hydrogen Atom

In this chapter, we look at aspects of the H-atom wavefunction and how it determines the atomic behavior in different circumstances.

The Hamiltonian

$$H = \frac{p^2}{2m_e} - \frac{e^2}{4\pi\epsilon_0 r} \quad \text{Convert to spherical coordinates}$$

$$H\psi = -\frac{\hbar^2}{2m_e r^2} \frac{1}{r} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \psi}{\partial r} \right) + \frac{\hbar^2}{2m_e r^2} \psi - \frac{e^2}{4\pi\epsilon_0 r} \psi$$

The angular momentum operator  $L^2\psi = -\hbar^2 \left[ \frac{1}{\sin\theta} \frac{\partial}{\partial\theta} \left( \sin\theta \frac{\partial \psi}{\partial\theta} \right) + \frac{1}{\sin^2\theta} \frac{\partial^2 \psi}{\partial\phi^2} \right]$

The three operators,  $H$ ,  $L^2$ ,  $L_z$ , commute with each other. This implies can find simultaneous eigenstate of all three.

$$\Psi_{nlm}(r) = R_{nl}(r) Y_{lm}(\theta, \phi)$$

$$H\Psi_{nlm} = E_{nl} \Psi_{nlm}, \quad L^2 \Psi_{nlm} = \hbar^2 l(l+1) \Psi_{nlm}, \quad L_z \Psi_{nlm} = \hbar m_l \Psi_{nlm}$$

Don't blindly use formulas from one book in the equations of another. Can have different phases in definition.

The  $Y_{lm}(\theta, \phi)$  come whenever the potential is spherically symmetric. They are normalized over surface of unit sphere.

$$\iint_0^{2\pi} Y_{lm}^*(\theta, \phi) Y_{l'm'}(\theta, \phi) d\phi \sin\theta d\theta = \delta_{ll'} \delta_{mm'}$$

$$Y_{lm} = \begin{pmatrix} \text{const} \\ P_l^m(\cos\theta) e^{im\phi} \end{pmatrix} \quad \text{Associated Legendre function}$$

Look at the  $Y_{lm}$  and discuss generic properties.

As  $|ml|$  increases, less probability near  $\cos\theta \approx \pm 1$ !

Number of nodal planes

The  $Y_{lm}$  are eigenstates of the parity operator

$$P Y_{lm}(\theta, \phi) = Y_{lm}(\pi - \theta, \phi + \pi) = (-1)^l Y_{lm}(\theta, \phi) \quad \text{Show?}$$

The radial part of  $\psi$  satisfies

$$-\frac{\hbar^2}{2m} \left[ \frac{1}{r^2} \frac{d}{dr} \left( r^2 \frac{dR_{nl}(r)}{dr} \right) \right] + \underbrace{\left( \frac{\hbar^2 l(l+1)}{2mr^2} - \frac{e^2}{4\pi\epsilon_0 r} \right) R_{nl}(r)}_{V_{\text{eff}}} = E_{nl} R_{nl}(r)$$

$$\text{Use } R_{nl}(r) = P_{nl}(r)/r \Rightarrow -\frac{\hbar^2}{2m} P_{nl}''(r) + V_{\text{eff}}(r)P_{nl}(r) = E_{nl} P_{nl}(r)$$

Sketch  $V_{\text{eff}}(r)$  and implications for different  $l$

See book for definition in Table 2.2 (Beware  $\rho = \frac{Zr}{na_0}$ , different in each!)

Show plots

Useful expressions

$$|\Psi_{(0)}|^2 = \frac{1}{\pi} \left( \frac{Z}{na_0} \right)^3 \quad l=0$$

$$l>0 \quad \left\langle \frac{1}{r^3} \right\rangle = \frac{1}{l(l+\frac{1}{2})(l+1)} \left( \frac{Z}{na_0} \right)^3$$

$$\left\langle \frac{1}{r} \right\rangle = \frac{1}{n^2} \frac{Z}{a_0}$$

Estimate effect from finite nuclear radius

$$V_{\text{act}}(r) - V_{\text{pure}}(r) \approx \frac{e^2}{4\pi\epsilon_0 r} \quad \text{for } r < r_N \sim 10^{-15} \text{ m}$$

$$\langle \Delta V \rangle = \iiint_{\text{nuc}} \Delta V(r) |\Psi_{(0)}(\vec{r})|^2 dr \approx \iiint_{a_0}^{r_N} \frac{e^2}{4\pi\epsilon_0 r} |\Psi_{(0)}(\vec{r})|^2 r^2 dr \sin \theta d\theta d\phi$$

$$\approx 4\pi |\Psi_{(0)}|^2 \int_0^{r_N} \frac{e^2}{4\pi\epsilon_0 r} r^2 dr = 2\pi |\Psi_{(0)}|^2 \frac{r_N^2 e^2}{4\pi\epsilon_0}$$

$$= 2 \frac{Z^3}{n^3} \left( \frac{r_N}{a_0} \right)^2 \frac{e^2}{4\pi\epsilon_0 a_0}$$

$$= 2 \frac{Z^3}{n^3} \left( \frac{r_N}{a_0} \right)^2 27.21 \text{ eV} \approx \frac{Z^3}{n^3} 10^{-8} \text{ eV} \approx \frac{Z^3}{n^3} \text{ MHz}$$

In QM, you need to be able to find the "matrix elements" of operators

$$Q_{ab} = \langle \psi_a | Q_b | \psi_b \rangle$$

An important one in AMO is the effect from light

In length gauge, the operator part is

$$Q_b = \hat{\vec{E}}_{\text{rad}} - \frac{1}{r} \quad \vec{E}_x = \hat{x} \cdot \vec{E}_{\text{rad}} \quad \text{etc}$$

polarization

Substitute the form for  $\psi$  for H-atom

$$Q_{n_2 m_2, n_1 m_1} = D_{12} I_{\text{ang}}$$

$$D_{12} = \int_0^\infty R_{n_2 l_2}(r) r R_{n_1 l_1}(r) r^2 dr$$

$$I_{\text{ang}} = \sum_{\substack{\text{all} \\ \text{ang}}} \sum_{l_2 m_2} Y^*(l_2, \phi) (\hat{E}_x \sin \theta \cos \phi + \hat{E}_y \sin \theta \sin \phi + \hat{E}_z \cos \theta) Y_{l_2 m_2}(\theta, \phi) d\phi \sin \theta d\theta$$

$$\text{Use } \cos \phi = \frac{1}{2} (e^{i\phi} + e^{-i\phi}) \quad \sin \phi = \frac{-i}{2} (e^{i\phi} - e^{-i\phi})$$

$$I_{\text{ang}} = \langle l_2 m_2 | \frac{\hat{E}_x - i \hat{E}_y}{\sqrt{2}} \frac{\sin \theta e^{i\phi}}{\sqrt{2}} + \frac{\hat{E}_x + i \hat{E}_y}{\sqrt{2}} \frac{\sin \theta e^{-i\phi}}{\sqrt{2}} + \hat{E}_z \cos \theta | l_1 m_1 \rangle$$

Depending on polarization, different states can be connected.

$\pi$ -Polarization

$$\begin{aligned} \langle l_2 m_2 | \cos \theta | l_1 m_1 \rangle &= \delta_{m_2 m_1} \left[ \delta_{l_2, l_1+1} \sqrt{\frac{l_2^2 - m_1^2}{(2l_2+1)(2l_1+1)}} + \delta_{l_2+1, l_1} \sqrt{\frac{l_2^2 - m_1^2}{(2l_2-1)(2l_1+1)}} \right] \\ &= \delta_{m_2 m_1} \left[ \delta_{l_2, l_1+1} + \delta_{l_2+1, l_1} \right] \sqrt{\frac{l_2^2 - m_1^2}{4l_2^2 - 1}} \quad l_2 = \max(l_1, l_2) \end{aligned}$$

$\sigma$ -Polarization

$$\langle l_2 m_2 | \sin \theta e^{+i\phi} | l_1 m_1 \rangle = \delta_{m_2, m_1 \pm 1} \left[ \pm \delta_{l_2, l_1+1} \sqrt{\frac{(l_2 \pm m_1)(l_2+1 \mp m_1)}{(2l_2+1)(2l_2-1)}} \pm \delta_{l_2+1, l_1} \sqrt{\frac{(l_1 \mp m_1)(l_1-1 \pm m_1)}{(2l_1+1)(2l_1-1)}} \right]$$

Note the selection rules!

The spin orbit interaction comes from a relativistic correction to the Hamiltonian. One way to think about: In the frame of the electron, the proton going around makes a B-field that interacts with the electron spin.

$$H_{\text{so}} \approx \frac{1}{2m_e^2 c^2} \frac{e^2}{4\pi\epsilon_0} \frac{1}{r^3} \vec{L} \cdot \vec{S}$$

(book eq. 2.51) uses  $\vec{l} = \frac{\vec{L}}{\hbar}$ ,  $\vec{s} = \frac{\vec{S}}{\hbar}$ )

Instead of using states  $|l, m_l\rangle |s, m_s\rangle$  combine them into total angular momentum states. This uses Clebsch-Gordan coefficients

$$\frac{\vec{J}}{\hbar} = \frac{1}{2} [\vec{L} + \vec{S}] \rightarrow \vec{L} \cdot \vec{S} = \frac{1}{2} [\vec{J}^2 - \vec{L}^2 - \vec{S}^2]$$

Look at the general case  $\vec{J} = \vec{J}_1 + \vec{J}_2$  with eigenstates  $|J_1, J_2, J, M_J\rangle$  and eigenvalues

$$\frac{\vec{J}^2}{\hbar^2} |J_1, J_2, J, M_J\rangle = \hbar^2 J(J+1) |J_1, J_2, J, M_J\rangle \quad J_z |J_1, J_2, J, M_J\rangle = \hbar M_J |J_1, J_2, J, M_J\rangle$$

The allowed values are  $J = J_1 + J_2, J_1 + J_2 - 1, \dots, |J_1 - J_2|$ ;  $-J \leq M_J \leq J$

Often need the connection between  $|J_1, J_2, J, M_J\rangle$  and the  $|J_1, J_2, M_1, M_2\rangle = |J_1, M_1\rangle |J_2, M_2\rangle$ . These are the Clebsch-Gordan coefficients

$$|J_1, J_2, J, M_J\rangle = \sum_{M_1, M_2} |J_1, J_2, M_1, M_2\rangle \langle J_1, J_2, M_1, M_2 | J_1, J_2, J, M_J\rangle$$

Example  $\ell=1, s=\frac{1}{2}$

$$|1, \frac{1}{2}, \frac{3}{2}, \frac{3}{2}\rangle = |1, 1\rangle |1, \frac{1}{2}, \frac{1}{2}\rangle$$

$$|1, \frac{1}{2}, \frac{3}{2}, \frac{1}{2}\rangle = \frac{1}{\sqrt{3}} |1, 1\rangle |1, \frac{1}{2}, -\frac{1}{2}\rangle + \sqrt{\frac{2}{3}} |1, 0\rangle |1, \frac{1}{2}, \frac{1}{2}\rangle$$

$$|1, \frac{1}{2}, \frac{3}{2}, -\frac{1}{2}\rangle = \sqrt{\frac{2}{3}} |1, 0\rangle |1, \frac{1}{2}, -\frac{1}{2}\rangle + \sqrt{\frac{1}{3}} |1, -1\rangle |1, \frac{1}{2}, \frac{1}{2}\rangle$$

$$|1, \frac{1}{2}, \frac{3}{2}, -\frac{3}{2}\rangle = |1, -1\rangle |1, \frac{1}{2}, -\frac{1}{2}\rangle$$

$$|1, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}\rangle = \sqrt{\frac{2}{3}} |1, 1\rangle |1, \frac{1}{2}, -\frac{1}{2}\rangle - \sqrt{\frac{1}{3}} |1, 0\rangle |1, \frac{1}{2}, \frac{1}{2}\rangle$$

$$|1, \frac{1}{2}, \frac{1}{2}, -\frac{1}{2}\rangle = \sqrt{\frac{1}{3}} |1, 0\rangle |1, \frac{1}{2}, -\frac{1}{2}\rangle - \sqrt{\frac{2}{3}} |1, -1\rangle |1, \frac{1}{2}, \frac{1}{2}\rangle$$

The main effect is to split the energy levels. The total energies are

$$E_{\text{tot}} = E_{n\ell} + \frac{1}{2} \beta_{n\ell} [j(j+1) - \ell(\ell+1) - s(s+1)]$$

$$\beta_{n\ell} = \frac{\hbar^2}{2m_e c^2} \frac{e^2}{4\pi\epsilon_0} \frac{z^3}{n^3 a_0^3} \frac{1}{\ell(\ell+1)s(s+1)} = \frac{\hbar^2}{2m_e a_0^2} \alpha^2 \frac{z^3}{n^3} \frac{1}{\ell(\ell+1)s(s+1)}$$

Get the form for some states

$$E_{n0\frac{1}{2}} = E_{n0}$$

$$E_{n1\frac{3}{2}} = E_{n1} + \frac{1}{2} \beta_{n1}$$

$$E_{n2\frac{5}{2}} = E_{n2} + \beta_{n2}$$

$$E_{n1\frac{1}{2}} = E_{n1} - \beta_{n1}$$

$$E_{n2\frac{3}{2}} = E_{n2} - \frac{3}{2} \beta_{n2}$$

Note that the weighted average doesn't change. (The trace of  $\vec{L} \cdot \vec{S}$  is zero.)

$$(6E_{n2\frac{5}{2}} + 4E_{n2\frac{3}{2}})/10 = E_{n2} + \frac{6}{10}\beta_{n2} - \frac{6}{10}\beta_{n2} = E_{n2}$$

Compare to experiment

$$(E_{2\frac{3}{2}} - E_{2\frac{1}{2}})/hc = R_\infty \alpha^2 \frac{1}{8} \frac{3/2}{1/2} = 109737.31568160 \text{ cm}^{-1} (7.2973525693 \times 10^{-3})/16$$

$$= 0.365229 \text{ cm}^{-1}$$

CODATA

CODATA

$$\text{From NIST} = 82259.2850014 \text{ cm}^{-1} - 82258.9191133 \text{ cm}^{-1} = 0.3658881 \text{ cm}^{-1}$$

Why the difference?

What is the frequency?  $f = 0.366 \text{ cm}^{-1} 3 \times 10^{10} \frac{\text{cm}}{\text{s}} = 11 \text{ GHz}$

Note the spin-orbit splitting is proportional to  $\frac{z^3}{n^3} \frac{1}{\ell(\ell+1)}$

Other relativistic effects? Mass (that is relativistic KE), magnetic dipole of proton (hyperfine), QED (Lamb shift)

The spin-orbit effect changes the selection rules for transitions

$$\langle n_1 l_1 j_1 m_1 | \neq | n_2 l_2 j_2 m_2 \rangle = D_{l_2} \langle l_1 s j_1 m_1 | \cos\theta | l_2 s j_2 m_2 \rangle$$

Still must have  $|l_1 - l_2| = 1$  and  $m_1 = m_2$ . Can have  $|j_1 - j_2| = 1$  or 0

Example  $n_1 s m_1 = \frac{1}{2} \rightarrow n_2 p j m_2$

$$\langle l_1 s j_1 m_1 \rangle = |0 \downarrow \downarrow \downarrow \downarrow \downarrow \downarrow \rangle = |0, 0\rangle | \frac{1}{2}, \frac{1}{2}\rangle \rightarrow 0$$

$$\langle 0 \downarrow \downarrow \downarrow \downarrow \downarrow \downarrow | \cos\theta | 1 \downarrow \downarrow \downarrow \downarrow \downarrow \downarrow \rangle = \langle \frac{1}{2}, \frac{1}{2} \uparrow \downarrow \downarrow \downarrow \downarrow \downarrow \rangle \langle 0, 0 | \cos\theta | 1, 1 \rangle$$

$$|1 \downarrow \downarrow \downarrow \downarrow \downarrow \downarrow \rangle = \sqrt{\frac{2}{3}} \langle \frac{1}{2}, \frac{1}{2} | \downarrow \downarrow \downarrow \downarrow \downarrow \downarrow \rangle \langle 0, 0 | \cos\theta | 1, 0 \rangle \neq 0$$

$$|1 \downarrow \downarrow \downarrow \downarrow \downarrow \downarrow \rangle = \sqrt{\frac{2}{3}} \langle \frac{1}{2} \frac{1}{2} \uparrow \downarrow \downarrow \downarrow \downarrow \downarrow \rangle \langle 0, 0 | \cos\theta | 1, 0 \rangle = 0$$

$$|1 \downarrow \downarrow \downarrow \downarrow \downarrow \downarrow \rangle = \langle \frac{1}{2}, \frac{1}{2} \downarrow \downarrow \downarrow \downarrow \downarrow \downarrow \rangle \langle 0, 0 | \cos\theta | 1, -1 \rangle = 0$$

$$|1 \downarrow \downarrow \downarrow \downarrow \downarrow \downarrow \rangle = -\frac{1}{\sqrt{3}} \langle \frac{1}{2}, \frac{1}{2} \downarrow \downarrow \downarrow \downarrow \downarrow \downarrow \rangle \langle 0, 0 | \cos\theta | 1, 0 \rangle \neq 0$$

$$|1 \downarrow \downarrow \downarrow \downarrow \downarrow \downarrow \rangle = \frac{1}{\sqrt{3}} \langle \frac{1}{2}, \frac{1}{2} \downarrow \downarrow \downarrow \downarrow \downarrow \downarrow \rangle \langle 0, 0 | \cos\theta | 1, 0 \rangle = 0$$

Note the matrix element is 2x bigger to  $j=\frac{3}{2}$  than to  $\frac{1}{2}$ . Also note sum of squares is same as without spin or bit.