

## Chap 4 - Quantum Mechanics in 3 Dimensions

We will start by first doing two dimensions. This will let us sneak up on 3-D! The most important question is "How to write Schrodinger Equation for 2D?"

Always  $i\hbar \frac{\partial}{\partial t} |\Psi(t)\rangle = \hat{H} |\Psi(t)\rangle$

The question is "what is  $\hat{H}$  in 2D?"  $\hat{H} = \frac{\hat{p}^2}{2m} + V(\hat{x}, \hat{y}, t)$

$$\hat{H} = \frac{1}{2m} (\hat{P}_x^2 + \hat{P}_y^2) + V(\hat{x}, \hat{y}, t)$$

A very common representation  $\hat{P}_x^2 + \hat{P}_y^2 = \left(\frac{\hbar}{i} \frac{\partial}{\partial x}\right)^2 + \left(\frac{\hbar}{i} \frac{\partial}{\partial y}\right)^2 = -\hbar^2 \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2}\right)$

A Schr. Eq. representation

$$i\hbar \frac{\partial}{\partial t} \Psi(x, y, t) = -\frac{\hbar^2}{2m} \left( \frac{\partial^2 \Psi(x, y, t)}{\partial x^2} + \frac{\partial^2 \Psi(x, y, t)}{\partial y^2} \right) + V(x, y, t) \Psi(x, y, t)$$

Just like the 1D case you can find eigenstates with the form

$$\Psi(x, y, t) = \psi_n(x, y) e^{-i E_n t / \hbar}$$

The  $\psi_n$  are solutions of  $H \psi_n = E_n \psi_n$

$$-\frac{\hbar^2}{2m} \left( \frac{\partial^2 \psi_n(x, y)}{\partial x^2} + \frac{\partial^2 \psi_n(x, y)}{\partial y^2} \right) + V(x, y) \psi_n(x, y) = E_n \psi_n(x, y)$$

For a general  $V(x, y)$ , the  $\psi_n(x, y)$  does not have special symmetry and the eigenfunctions can be numbered by a single index. If the potential is separable, it is often more useful to have 2 indices.

Special Case:  $V(x, y) = V^{(1)}(x) + V^{(2)}(y)$

Separation of variables in spatial coordinates:

$$\Psi_n(x, y) = f_{n_x}(x) g_{n_y}(y)$$

Substitute into Sch. Eq.

$$g(y) \left( -\frac{\hbar^2}{2m} \frac{d^2 f_{n_x}(x)}{dx^2} + V^{(1)}(x) f_{n_x}(x) \right) + f_{n_x}(x) \left( -\frac{\hbar^2}{2m} \frac{d^2 g_{n_y}(y)}{dy^2} + V^{(2)}(y) g_{n_y}(y) \right) = E_n f_{n_x}(x) g_{n_y}(y)$$

Set  $E_n = E_{n_x}^{(1)} + E_{n_y}^{(2)}$

$$-\frac{\hbar^2}{2m} \frac{d^2 f_{n_x}(x)}{dx^2} + V^{(1)}(x) f_{n_x}(x) = E_{n_x}^{(1)} f_{n_x}(x) \quad -\frac{\hbar^2}{2m} \frac{d^2 g_{n_y}(y)}{dy^2} + V^{(2)}(y) g_{n_y}(y) = E_{n_y}^{(2)} g_{n_y}(y)$$

If  $V^{(1)} = V^{(2)}$ , then the Hamiltonian has the property  $H(x, y) = H(y, x)$ . This means the  $\Psi_n(x, y)$  can have the property  $\Psi_n(x, y) = \pm \Psi_n(y, x)$ . Does our solution have this property? No! Unless  $n_x = n_y$ .

Example of square well  $V(x, y) = 0$  for  $0 \leq x \leq a$  and  $0 \leq y \leq a$   
 $\Psi_{n_x n_y}(x, y) = \sqrt{\frac{2}{a}} \sin\left(\frac{n_x \pi x}{a}\right) \sqrt{\frac{2}{a}} \sin\left(\frac{n_y \pi y}{a}\right)$

$$\Psi_{3,4}(x, y) = \frac{2}{a} \sin\left(\frac{3\pi x}{a}\right) \sin\left(\frac{4\pi y}{a}\right) \neq \Psi_{3,4}(y, x) = \frac{2}{a} \sin\left(\frac{3\pi y}{a}\right) \sin\left(\frac{4\pi x}{a}\right)$$

The combinations  $\left. \begin{aligned} \Psi_{3,4}^{\pm}(x, y) &= \frac{1}{\sqrt{2}} (\Psi_{3,4}(x, y) \pm \Psi_{3,4}(y, x)) \\ &= \frac{1}{\sqrt{2}} (\Psi_{3,4}(x, y) \pm \Psi_{4,3}(x, y)) \end{aligned} \right\}$  have the symmetry property  
 $\Psi_{3,4}^{\pm}(x, y) = \pm \Psi_{3,4}(y, x)$

When the eigenstates have the form  $f_{n_x}(x) g_{n_y}(y)$ , the nodal lines are a grid in x and y.

Cylindrical coordinates :  $x = \rho \cos\phi$      $y = \rho \sin\phi$   
 $\rho = \sqrt{x^2 + y^2}$      $\tan\phi = y/x$

Useful when  $V(x, y) = V(\rho)$

When  $V$  is only a function of  $\rho$ , the classical angular momentum

$$L_z = (\vec{r} \times \vec{p}) \cdot \hat{z} = xP_y - yP_x$$

is a conserved quantity. To see this first get the form of the force

$$F_x = -\frac{\partial V(\rho)}{\partial x} = -\frac{\partial \rho}{\partial x} \frac{dV}{d\rho} = -\frac{x}{\rho} \frac{dV}{d\rho} \quad F_y = -\frac{\partial V(\rho)}{\partial y} = -\frac{y}{\rho} \frac{dV}{d\rho}$$

$$\frac{dL_z}{dt} = \frac{dx}{dt} P_y + x \frac{dP_y}{dt} - \frac{dy}{dt} P_x - y \frac{dP_x}{dt}$$

$$= V_x P_y - V_y P_x + x F_y - y F_x = 0 + -\frac{xy}{\rho} \frac{dV}{d\rho} - \frac{yx}{\rho} \frac{dV}{d\rho} = 0$$

In Q.M. we need to show  $[\hat{L}_z, \hat{H}] = 0$

$$[L_z, V] = [xP_y, V] - [yP_x, V] = x[P_y, V] - y[P_x, V]$$

$$= \frac{\hbar}{i} (x \frac{\partial V}{\partial y} - y \frac{\partial V}{\partial x}) = \frac{\hbar}{i} (x \frac{y}{\rho} \frac{dV}{d\rho} - y \frac{x}{\rho} \frac{dV}{d\rho}) = 0$$

$$[L_z, P_x^2 + P_y^2] = [xP_y, P_x^2 + P_y^2] - [yP_x, P_x^2 + P_y^2] = [x, P_x^2] P_y - [y, P_y^2] P_x$$

$$= 2i\hbar P_x P_y - 2i\hbar P_y P_x = 0 \quad \underline{\text{Yes!}}$$

Since  $[\hat{L}_z, \hat{H}] = 0$ , every energy eigenstate can also be an eigenstate of  $L_z$ . Very important!!

Can we find an expression for  $\hat{L}_z$  in terms of  $\rho, \phi$ ?

$$L_z = \frac{\hbar}{i} (x \frac{\partial}{\partial y} - y \frac{\partial}{\partial x}) = \frac{\hbar}{i} (x \frac{\partial \rho}{\partial y} \frac{\partial}{\partial \rho} + x \frac{\partial \phi}{\partial y} \frac{\partial}{\partial \phi} - y \frac{\partial \rho}{\partial x} \frac{\partial}{\partial \rho} - y \frac{\partial \phi}{\partial x} \frac{\partial}{\partial \phi})$$

$$= \frac{\hbar}{i} \left( x \frac{y}{\rho} \frac{\partial}{\partial \rho} + x \frac{x}{x^2 + y^2} \frac{\partial}{\partial \phi} - y \frac{x}{\rho} \frac{\partial}{\partial \rho} - y \left( -\frac{y}{x^2 + y^2} \right) \frac{\partial}{\partial \phi} \right) = \frac{\hbar}{i} \frac{\partial}{\partial \phi}$$

Find the eigenstates of  $L_z$ .

$$\frac{\hbar}{i} \frac{\partial}{\partial \phi} F_m(\phi) = \lambda_m F_m(\phi) \Rightarrow F_m(\phi) = e^{i \lambda_m \phi / \hbar}$$

$m$  is an unfortunate index, but standard. Don't confuse w/ mass!

How to figure out which  $\lambda_m$  allowed?

$$F_m(\phi + 2\pi) = F_m(\phi) \Rightarrow e^{i \lambda_m 2\pi / \hbar} = 1 \quad \lambda_m = m \hbar \\ m = \dots -2, -1, 0, 1, 2, \dots$$

Stop and recognize what this means. The allowed eigenvalues of  $L_z$  are integer times  $\hbar$ ! (Reminder of Bohr quantization cond.)

From many different sources  $\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} = \frac{\partial^2}{\partial \rho^2} + \frac{1}{\rho} \frac{\partial}{\partial \rho} + \frac{1}{\rho^2} \frac{\partial^2}{\partial \phi^2}$

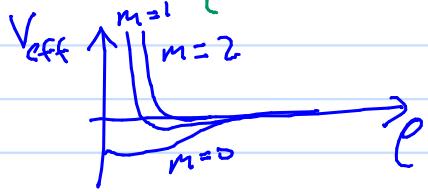
$$\hat{H} = -\frac{\hbar^2}{2M} \left( \frac{\partial^2}{\partial \rho^2} + \frac{1}{\rho} \frac{\partial}{\partial \rho} \right) - \frac{\hbar^2}{2M\rho^2} \frac{\partial^2}{\partial \phi^2} + V(\rho)$$

$$= -\frac{\hbar^2}{2M} \left( \frac{\partial^2}{\partial \rho^2} + \frac{1}{\rho} \frac{\partial}{\partial \rho} \right) + \frac{1}{2M\rho^2} L_z^2 + V(\rho)$$

The time independent Schr. Eq. can have solutions of the form.

$$Y_n(\rho, \phi) = F_{n_\rho, m}(\rho) e^{im\phi} \\ -\frac{\hbar^2}{2M} \left( \frac{d^2 F_{n_\rho, m}}{d\rho^2} + \frac{1}{\rho} \frac{d F_{n_\rho, m}}{d\rho} \right) e^{im\phi} + \frac{\hbar^2 m^2}{2M\rho^2} F_{n_\rho, m} e^{im\phi} + V(\rho) F_{n_\rho, m} e^{im\phi} = E_{n_\rho, m} F_{n_\rho, m} e^{im\phi} \\ -\frac{\hbar^2}{2M} \left( \frac{d^2 F_{n_\rho, m}}{d\rho^2} + \frac{1}{\rho} \frac{d F_{n_\rho, m}}{d\rho} \right) + V_{\text{eff}}(\rho) F_{n_\rho, m} = E_{n_\rho, m} F_{n_\rho, m} \quad V_{\text{eff}} = V + \frac{\hbar^2 m^2}{2M\rho^2}$$

It is very important to realize what is going on. For each  $m$  (eigenvl of  $L_z$ ), there are an  $\infty$  number of solutions that have  $n_\rho - 1$  nodes in the function  $F_{n_\rho, m}$



Useful energy relations:  $E_{n_\rho, m} = E_{n_\rho, -m}$   
 $E_{n_\rho+1, m} > E_{n_\rho, m}$   
 $E_{n_\rho, |m|+1} > E_{n_\rho, |m|}$

The orthonormality relation for the two special cases have a special form.

$$\iint \Psi_{n_x n_y}^*(x, y) \Psi_{n'_x n'_y}^*(x, y) dx dy = \delta_{n_x n'_x} \delta_{n_y n'_y}$$

$$\iint \Psi_{n_\rho m}^*(\rho, \phi) \Psi_{n'_\rho m'}^*(\rho, \phi) d\phi \rho d\rho = \delta_{n_\rho n'_\rho} \delta_{m m'}$$

Now let's look at the 3D case  $H \Psi_n(x, y, z) = E_n \Psi_n(x, y, z)$

As with the 2D case, the  $H$  for the 3D case is extended in the obvious way.

$$H = \frac{1}{2M} (P_x^2 + P_y^2 + P_z^2) + V(x, y, z) = -\frac{\hbar^2}{2M} \left( \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) + V(x, y, z)$$

As with the 2D case, the generic  $V(x, y, z)$  does not give any special symmetry to  $\Psi_n(x, y, z)$  so can just number them with  $n$ . To have bound state,  $V(x, y, z) > E$  as  $x, y, z \rightarrow \infty$  in every direction is obvious restriction but is too strong.

Example  $V = \frac{1}{2} m \omega^2 (x^2 + y^2) - \alpha \delta(z)$  could give bound state  $E > 0$ .

We will focus on special cases where  $E_n$  and  $\Psi_n$  can be found analytically.

Special Case:  $V = V^{(1)}(x) + V^{(2)}(y) + V^{(3)}(z)$

Use separation of variables  $\Psi(x, y, z) = \Psi_{n_x}^{(1)}(x) \Psi_{n_y}^{(2)}(y) \Psi_{n_z}^{(3)}(z)$

Plug into Sch. Eq.

$$\begin{aligned} & \left( H^{(1)} \Psi_{n_x}^{(1)}(x) \right) \Psi_{n_y}^{(2)}(y) \Psi_{n_z}^{(3)}(z) + \Psi_{n_x}^{(1)}(x) \left( H^{(2)} \Psi_{n_y}^{(2)}(y) \right) \Psi_{n_z}^{(3)}(z) + \Psi_{n_x}^{(1)}(x) \Psi_{n_y}^{(2)}(y) \left( H^{(3)} \Psi_{n_z}^{(3)}(z) \right) \\ &= E_{n_x}^{(1)} \Psi_{n_x}^{(1)} \Psi_{n_y}^{(2)} \Psi_{n_z}^{(3)} + E_{n_y}^{(2)} \Psi_{n_x}^{(1)} \Psi_{n_y}^{(2)} \Psi_{n_z}^{(3)} + E_{n_z}^{(3)} \Psi_{n_x}^{(1)} \Psi_{n_y}^{(2)} \Psi_{n_z}^{(3)} \end{aligned}$$

$$\Rightarrow E_{n_x n_y n_z} = E_{n_x}^{(1)} + E_{n_y}^{(2)} + E_{n_z}^{(3)}$$

Example of "cube" well  $V(x, y, z) = 0$  if  $0 \leq x \leq a_x$  and  $0 \leq y \leq a_y$  and  $0 \leq z \leq a_z$   
 $= \infty$  otherwise

$$\Psi_{n_x n_y n_z} = \sqrt{\frac{2}{a_x}} \sin\left(\frac{n_x \pi x}{a_x}\right) \sqrt{\frac{2}{a_y}} \sin\left(\frac{n_y \pi y}{a_y}\right) \sqrt{\frac{2}{a_z}} \sin\left(\frac{n_z \pi z}{a_z}\right)$$

$$E_{n_x n_y n_z} = \frac{\hbar^2 \pi^2}{2M} \left( \frac{n_x^2}{a_x^2} + \frac{n_y^2}{a_y^2} + \frac{n_z^2}{a_z^2} \right)$$

Special - Special case  $V^{(1)} = V^{(2)} = V^{(3)} = V$

For this case most states are degenerate.

$n_x = n_y = n_z$  degeneracy 1

2 n's same, other different degeneracy 3 (example 2,2,5 ; 2,5,2 ; 5,2,2)  
 all n's different degeneracy 6

Special case  $V = V^{(1)}(r) + V^{(2)}(z) \Rightarrow H = H^{(1)}(r, \phi) + H^{(2)}(z)$

Use separation of variables  $\Psi_{n_r m n_z} = F_{n_r m}^{(1)}(r) e^{im\phi} \Psi_{n_z}^{(2)}$

$$\begin{aligned} H \Psi_{n_r m n_z} &= \left( H^{(1)}(r, \phi) F_{n_r m}^{(1)}(r) e^{im\phi} \right) \Psi_{n_z}^{(2)} + F_{n_r m}^{(1)}(r) e^{im\phi} \left( H^{(2)}(z) \Psi_{n_z}^{(2)} \right) \\ &= E_{n_r m}^{(1)} F_{n_r m}^{(1)} e^{im\phi} \Psi_{n_z}^{(2)} + E_{n_z}^{(2)} F_{n_r m}^{(1)} e^{im\phi} \Psi_{n_z}^{(2)} \end{aligned}$$

$$E_{n_r m n_z} = E_{n_r m}^{(1)} + E_{n_z}^{(2)}$$

These eigenstates are also eigenstates of  $L_z$   $L_z \Psi_{n_r m n_z} = \hbar m \Psi_{n_r m n_z}$

One of the most common form for  $V$  is that it only depends on distance from the origin:  $V(x, y, z) = V(r)$

Classically this gives a force with magnitude that only depends on  $r$  and direction  $\hat{r}$

Classically  $\frac{dL_z}{dt} = 0$  and  $\frac{dL_x}{dt} = 0$  and  $\frac{dL_y}{dt} = 0$

In Q.M. can show  $[\hat{L}_i, \hat{H}] = 0$  which means  $\langle \hat{L}_i \rangle = \text{constant}$

Very important: this does not mean can find eigenstates of  $\hat{H}$ ,  $\hat{L}_x$ ,  $\hat{L}_y$ , and  $\hat{L}_z$ . Reason: the different  $L$ 's do not commute!

To see how this works, work out the different commutators:

$$L_x = y P_z - z P_y \quad L_y = z P_x - x P_z \quad L_z = x P_y - y P_x$$

$$\begin{aligned} [L_x, x] &= 0 & [L_x, y] &= i\hbar z & [L_x, z] &= -i\hbar y \\ [L_y, x] &= -i\hbar z & [L_y, y] &= 0 & [L_y, z] &= i\hbar x \\ [L_z, x] &= i\hbar y & [L_z, y] &= -i\hbar x & [L_z, z] &= 0 \end{aligned}$$

$$\begin{aligned} [L_x, P_x] &= 0 & [L_x, P_y] &= i\hbar P_z & [L_x, P_z] &= -i\hbar P_y \\ [L_y, P_x] &= -i\hbar P_z & [L_y, P_y] &= 0 & [L_y, P_z] &= i\hbar P_x \\ [L_z, P_x] &= i\hbar P_y & [L_z, P_y] &= -i\hbar P_x & [L_z, P_z] &= 0 \end{aligned}$$

$$\begin{aligned} [L_x, L_x] &= 0 & [L_x, L_y] &= i\hbar L_z & [L_x, L_z] &= -i\hbar L_y \\ [L_y, L_x] &= -i\hbar L_z & [L_y, L_y] &= 0 & [L_y, L_z] &= i\hbar L_x \\ [L_z, L_x] &= i\hbar L_y & [L_z, L_y] &= -i\hbar L_x & [L_z, L_z] &= 0 \end{aligned}$$

$$[L_a, Q_b] = i\hbar \epsilon_{abc} Q_c \quad Q = P \text{ or } x \text{ or } L$$

$$\epsilon_{123} = \epsilon_{312} = \epsilon_{231} = 1 \quad \epsilon_{213} = \epsilon_{132} = \epsilon_{321} = -1 \quad \text{all else} = 0$$

Can show that  $L_i$  commutes with any scalar combination of  $\vec{P}$ ,  $\vec{r}$ ,  $\vec{L}$

$$\begin{aligned} [L_x, x P_x + y P_y + z P_z] &= [L_x, y P_y] + [L_x, z P_z] \\ &= y [L_x, P_y] + [L_x, y] P_y + z [L_x, P_z] + [L_x, z] P_z \\ &= i\hbar y P_z + i\hbar z P_y + z (-i\hbar P_y) + -i\hbar y P_z \\ &= 0 \end{aligned}$$

$$[L_z, L_x^2 + L_y^2 + L_z^2] = [L_z, L^2] = 0 \quad \text{Very important.}$$

$$[L_i, H] = 0 \quad \text{means} \quad [L^2, H] = 0$$

We can find states that are simultaneously eigenstates of  $H$ ,  $L^2$ , and  $L_z$

$$\boxed{\Psi_{n_r, l, m}}$$

$n$  - quantum number for  $L_z$

$l$  - " " " "  $L^2$

$n_r$  - " " " "  $H_r$

From many places  $\nabla^2 = \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2}$

Spherical coordinates

$$x = r \sin \theta \cos \phi$$

$$y = r \sin \theta \sin \phi$$

$$z = r \cos \theta$$

many math books

interchange

$\theta$  and  $\phi$

$$\begin{aligned} \nabla^2 &= \frac{\partial^2}{\partial r^2} + \frac{2}{r} \frac{\partial}{\partial r} + \frac{1}{r^2} \frac{\partial^2}{\partial \theta^2} + \frac{1}{r^2} \cot \theta \frac{\partial}{\partial \theta} + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2}{\partial \phi^2} \\ &= \frac{\partial^2}{\partial r^2} + \frac{2}{r} \frac{\partial}{\partial r} + \frac{1}{r^2} \left( \frac{\partial^2}{\partial \theta^2} + \cot \theta \frac{\partial}{\partial \theta} + \frac{1}{\sin^2 \theta} \left[ \frac{\partial^2}{\partial \phi^2} \right] \right) \end{aligned}$$

Try a separation of variables

$$\Psi_{n_r, l, m}(r, \theta, \phi) = R_{n_r, l} (r) \Theta_{l, m} (\theta) \Phi_m (\phi)$$

$$\hat{H} \Psi = -\frac{\hbar^2}{2M} \left( \frac{d^2 R_{n_r}}{dr^2} + \frac{2}{r} \frac{d R_{n_r}}{dr} \right) \Theta_{l, m} \Phi_m - \frac{\hbar^2}{2Mr^2} \left( \frac{d^2 \Theta_{l, m}}{d\theta^2} + \cot \theta \frac{d \Theta_{l, m}}{d\theta} - \frac{m^2}{\sin^2 \theta} \Theta_{l, m} \right) R_{n_r, l} \Phi_m$$

$$+ V(r) R_{n_r, l} \Theta_{l, m} \Phi_m = E_{n_r, l, m} R_{n_r, l} \Theta_{l, m} \Phi_m$$

$$\frac{d^2 \Theta_{l, m}}{d\phi^2} = -m^2 \Theta_{l, m}$$

The equation in  $\theta$  gives

$$\frac{d^2 \Theta_{l, m}}{d\theta^2} + \cot \theta \frac{d \Theta_{l, m}}{d\theta} - \frac{m^2}{\sin^2 \theta} \Theta_{l, m} = -l(l+1) \Theta_{l, m}$$

with  $l = |m|, |m|+1, |m|+2, \dots$

The equation in  $r$  gives

$$-\frac{\hbar^2}{2m} \left( \frac{d^2 R_{n_r, l}}{dr^2} + \frac{2}{r} \frac{d R_{n_r, l}}{dr} \right) + \frac{\hbar^2 l(l+1)}{2Mr^2} R_{n_r, l} + V(r) R_{n_r, l} = E_{n_r, l} R_{n_r, l}$$

$$V_{\text{eff}}(r) = V(r) + \frac{\hbar^2 l(l+1)}{2Mr^2}$$

Important: Notice  $E$  does not depend on the  $m$ -quantum number.  
 Why??

Since the equation in  $\Theta$  is the same for all  $V(r)$ , look at its properties first.

$$\frac{1}{\sin \theta} \frac{d}{d\theta} \left( \sin \theta \frac{d(\Theta_{lm})}{d\theta} \right) - \frac{m^2}{\sin^2 \theta} \Theta_{lm} + l(l+1) \Theta_{lm} = 0$$

$$\frac{d}{d(\cos \theta)} \left( \sin^2 \theta \frac{d(\Theta_{lm})}{d(\cos \theta)} \right) - \frac{m^2}{\sin^2 \theta} \Theta_{lm} + l(l+1) \Theta_{lm} = 0 \quad \text{set } x = \cos \theta$$

$$\frac{d}{dx} \left( (1-x^2) \frac{d(\Theta_{lm})}{dx} \right) - \frac{m^2}{1-x^2} \Theta_{lm} + l(l+1) \Theta_{lm} = 0 \quad -1 \leq x \leq 1$$

The solutions of this equation are well known: associated Legendre function:  $P_l^m(x)$

Important to note that there are different sign conventions for  $m < 0$ . Griffiths uses  $P_l^{-m}(x) = P_l^m(x)$

The Legendre polynomial  $P_l(x) = P_l^0(x)$   $l=0, 1, 2, 3, \dots$

$$\text{Rodrigues formula: } P_l(x) \equiv \frac{1}{2^l l!} \frac{d^l}{dx^l} (x^2 - 1)^l$$

The number of nodes of  $P_l(x)$  is  $l$ ,  $P_l^m(x)$  is  $l - |m|$

$$P_0(x) = 1 \quad P_1(x) = x \quad P_2(x) = \frac{3}{2}x^2 - \frac{1}{2} \quad P_3(x) = \frac{5}{2}x^3 - \frac{3}{2}x \dots$$

Some properties of the  $P_l(x)$

$$\textcircled{1} \quad P_l(x) = (-1)^l P_l(-x)$$

$$\textcircled{2} \quad P_l(1) = 1 \quad \text{and} \quad P_l(-1) = (-1)^l$$

$$\textcircled{3} \quad |P_l(x)| < 1 \quad \text{for } l > 0 \quad \text{and} \quad |X| < 1$$

$$\textcircled{4} \quad \int_{-1}^1 P_l(x) P_{l'}(x) dx = 2 \delta_{ll'} / (2l+1)$$

The associated Legendre function  $P_l^m(x) = (1-x^2)^{\frac{|m|}{2}} \frac{d^{|m|}}{dx^{|m|}} P_l(x)$

Since the highest power of  $P_l(x)$  is  $x^l$ ,  $P_l^m(x) = 0$  if  $|m| > l$

The normalized combination  $\Theta_{\ell m}(\theta) \Phi_m(\phi)$  is called the spherical harmonics and is the symbol  $Y_{\ell m}$

$$Y_{\ell m}(\theta, \phi) = \sqrt{\frac{2\ell+1}{4\pi}} \frac{(\ell-|m|)!}{(\ell+|m|)!} P_{\ell}^{|m|}(\cos\theta) e^{im\phi}$$
$$\epsilon = 1 \quad \text{if } m < 0$$
$$= -1 \quad \text{if } m \geq 0$$

The  $\epsilon$  is the result of convention and must be followed!

Another common notation is  $Y_{\ell m}(\theta, \phi)$

The volume element in spherical coordinates is  $\sin\theta d\theta d\phi r^2 dr$

The normalization is  $\int_0^{\pi} \int_0^{2\pi} Y_{\ell m}^*(\theta, \phi) Y_{\ell m}(\theta, \phi) d\phi \sin\theta d\theta = \delta_{\ell\ell'} \delta_{mm'}$

Look at the  $Y_{\ell m}$  and discuss generic properties.

As  $|m|$  increases, less probability near  $\cos\theta \approx \pm 1$

Now look at the radial part of the Schr. Eq.

$$-\frac{\hbar^2}{2m} \left[ \frac{d^2 R_{n,l}}{dr^2} + \frac{2}{r} \frac{dR_{n,l}}{dr} \right] + \frac{\hbar^2 l(l+1)}{2Mr^2} R_{n,l} + V(r) R_{n,l} = E_{n,l} R_{n,l}$$

You get a somewhat simpler equation if you substitute  $R(r) = \frac{u(r)}{r}$

$$R' = \frac{u'}{r} - \frac{u}{r^2} \quad R'' = \frac{u''}{r} - \frac{2u'}{r^2} + \frac{2u}{r^3}$$

$$R'' + \frac{2}{r} R = \frac{u''}{r} - \frac{2u'}{r^2} + \frac{2u}{r^3} + \frac{2u'}{r^2} - \frac{2u}{r^3} = \frac{u''}{r}$$

$$-\frac{\hbar^2}{2M} \frac{d^2 u}{dr^2} + \left( V(r) + \frac{\hbar^2 l(l+1)}{2Mr^2} \right) u = E u \quad \begin{matrix} \text{looks like the 1D} \\ \text{Schr. Eq.} \end{matrix}$$

The normalization condition  $\int_0^\infty |R(r)|^2 r^2 dr = \int_0^\infty |u(r)|^2 dr$

Must have  $u(0) = 0$  Why?

Not obvious but the  $u(r)$  has to have a specific power law as  $r \rightarrow 0$  (not necessarily linear) when  $r^2 V(r) \rightarrow 0$  as  $r \rightarrow 0$   
At small  $r$ , need to satisfy

$$-\frac{\hbar^2}{2M} \frac{d^2 u}{dr^2} + \frac{\hbar^2 l(l+1)}{2Mr^2} u(r) = 0 \rightarrow \frac{d^2 u}{dr^2} = \frac{l(l+1)}{r^2} u$$

Try  $u(r) = A r^b$   $A b(b-1) r^{b-2} = A l(l+1) r^{b-2}$

$$\Rightarrow b(b-1) = l(l+1)$$

$$b = l+1 \quad \text{or} \quad b = -l \quad \text{why?} \\ \text{C not allowed?}$$

There are not a lot of potentials where you can find the  $u$  analytically: infinite spherical well, harmonic oscillator,  $-4/r$ , ...

Example: Infinite spherical well

$$V(r) = 0 \quad \text{if } r < a \\ = \infty \quad \text{if } r > a$$

$$-\frac{\hbar^2}{2M} \frac{d^2 u}{dr^2} + \frac{\hbar^2 l(l+1)}{2Mr^2} u = E u$$

Change variables  $r = x/k$   $E = \frac{\hbar^2 k^2}{2m}$

$$\frac{d^2 u}{dx^2} + \left(1 - \frac{\ell(\ell+1)}{x^2}\right) u = 0$$

The two linearly independent solutions are  $x J_\ell(x)$  and  $x N_\ell(x)$   
 These are spherical Bessel functions (the  $N_\ell$  is often called the Neumann function).

$$J_\ell(x) = (-x)^\ell \left(\frac{1}{x} \frac{d}{dx}\right)^\ell \left(\frac{\sin x}{x}\right) \quad N_\ell(x) = (-x)^\ell \left(\frac{1}{x} \frac{d}{dx}\right)^\ell \left(\frac{\cos x}{x}\right)$$

$$J_{\ell+1}(x) = \frac{2\ell+1}{x} J_\ell(x) - J_{\ell-1}(x) \text{ and similar } N_\ell$$

$$\begin{aligned} J_0(x) &= \frac{\sin x}{x} & J_1(x) &= \frac{\sin x}{x^2} - \frac{\cos x}{x} & J_2(x) &= \left(\frac{3}{x^3} - \frac{1}{x}\right) \sin x - \frac{3 \cos x}{x^2} \\ N_0(x) &= -\frac{\cos x}{x^2} & N_1(x) &= -\frac{\cos x}{x^3} - \frac{\sin x}{x} & N_2(x) &= -\left(\frac{3}{x^3} - \frac{1}{x}\right) \cos x - \frac{3 \sin x}{x^2} \end{aligned}$$

It is not obvious from the functions but

$$x J_\ell(x) \rightarrow \text{const} \cdot x^{\ell+1} \quad x N_\ell(x) \rightarrow \text{const}/x^\ell$$

Undo the change of variable  $u(r) = A r J_\ell(kr) + B r N_\ell(kr)$

Apply the boundary condition at  $r=0$  What is it?  $\Rightarrow B=0$

Apply the boundary condition at  $r=a$  What is it?  $\Rightarrow J_\ell(ka)=0$

Define  $\beta_{n\ell}$  so that  $J_\ell(\beta_{n\ell})=0$   $\beta_{n\ell}$  is the  $n^{\text{th}}$  0 of  $J_\ell(x)$ .

$$k_{n\ell} = \beta_{n\ell}/a \rightarrow E_{n\ell} = \frac{\hbar^2}{2ma^2} \beta_{n\ell}^2$$

The  $\beta_{n0} = n\pi$  All others are found numerically (see online ab.)

$$\beta_{11} = 4.493 \dots \quad \beta_{21} = 7.725 \dots \quad \beta_{31} = 10.90 \dots \quad \beta_{12} = 5.763 \quad \beta_{22} = 9.095 \dots \quad \beta_{32} = 12.32$$

The full, normalized wave function is

$$\Psi_{n\ell m}^{(r, \theta, \phi)} = A_{n\ell} J_\ell(k_{n\ell} r) Y_\ell^m(\theta, \phi)$$

$$A_{n\ell} \int_0^a J_\ell^2(k_{n\ell} r) r^2 dr = 1$$

normalization

Special case: Harmonic oscillator

$$V(x) = \frac{1}{2} M \omega^2 x^2 + \frac{1}{2} M \omega^2 y^2 + \frac{1}{2} M \omega^2 z^2$$

$$= \frac{1}{2} M \omega^2 r^2$$

You can solve using cartesian coords

$$E_{n_x n_y n_z} = (n_x + n_y + n_z + \frac{3}{2}) \hbar \omega$$

Allowed energies:  $\frac{3}{2} \hbar \omega, \frac{5}{2} \hbar \omega, \frac{7}{2} \hbar \omega, \dots$

Degeneracy      1            3            6

You can solve using spherical coords

$$-\frac{\hbar^2}{2M} \frac{d^2 u}{dr^2} + \left[ \frac{1}{2} M \omega^2 r^2 + \frac{\hbar^2 l(l+1)}{2Mr^2} \right] u(r) = E_{n_r l} u(r)$$

Use transformation variable  $\xi = \sqrt{\frac{M\omega}{\hbar}} r$  and factor out the small  $r$ ,  $\xi^{l+1}$ , and large  $r$ ,  $e^{-\xi^2/2}$ , behaviors

$$u = h(\xi) e^{-\xi^2/2}$$

Follow the same steps as Sec 2.3.2. The  $h(\xi)$  turn out to be associated Laguerre polynomials times  $\xi^{l+1}$

Substitute  $\xi$

$$-\frac{1}{2} \hbar \omega \frac{d^2 u}{d\xi^2} + \frac{1}{2} \hbar \omega \left( \xi^2 + \frac{l(l+1)}{\xi^2} \right) u = E u$$

$K = 2E/\hbar\omega$

$$\frac{d^2 u}{d\xi^2} + \left( K - \xi^2 - \frac{l(l+1)}{\xi^2} \right) u = 0$$

Now take out the large  $\xi$  behavior  $u(\xi) = h(\xi) e^{-\xi^2/2}$

$$u' = h' e^{-\xi^2/2} - \xi h e^{-\xi^2/2}$$

$$u'' = h'' e^{-\xi^2/2} - \xi h' e^{-\xi^2/2} - h e^{-\xi^2/2} - \xi h e^{-\xi^2/2} + \xi^2 h e^{-\xi^2/2}$$

$$= (h'' - 2\xi h' + (\xi^2 - 1)h) e^{-\xi^2/2}$$

$$h'' - 2\xi h' + (K-1 - \frac{\ell(\ell+1)}{\xi^2})h = 0$$

Write  $h(\xi) = \xi^{\ell+1} (a_0 + a_1 \xi + a_2 \xi^2 + \dots)$   
 From bdry condition at  $r=0$

Terms with  $\xi^{\ell+1}$

$$(\ell+1)\ell a_0 - 0 - \ell(\ell+1)a_0 = 0$$

$$\text{'' '' } \xi^\ell \quad (\ell+2)(\ell+1)a_1 - 0 - \ell(\ell+1)a_1 = 0 \Rightarrow a_1 = 0$$

$$\text{'' '' } \xi^{\ell+1} \quad (\ell+3)(\ell+2)a_2 - 2(\ell+1)a_0 + (K-1)a_0 - \ell(\ell+1)a_2 = 0$$

$$\text{'' '' } \xi^{\ell+2} \quad (\ell+4)(\ell+3)a_3 - 2(\ell+2)a_1 + (K-1)a_1 - \ell(\ell+1)a_3 = 0 \Rightarrow a_3 = 0$$

Only the even  $a_j \neq 0$ . All odd  $a_j = 0$ .

$$[(\ell+3+j)(\ell+2+j) - \ell(\ell+1)] a_{j+2} = [2(\ell+j+1) - K+1] a_j$$

Just as with the cartesian derivation, there is a solution for any value of  $K$  (which is proportional to  $E$ ). Must have the series truncate so  $\psi \rightarrow 0$  as  $r \rightarrow \infty$

$$K_{n_r, l} = \frac{2E_{n_r, l}}{\tau \omega} = 2n_r + 2l + 3$$

$$E_{n_r, l} = (n_r + l + \frac{3}{2}) \tau \omega$$

Remember that for each  $n_r, l$  there are  $2l+1$   $m$ -states  
 $m = -l, -l+1, \dots 0, \dots l-1, l$

$$\Psi_{n_r, l, m} = A_{n_r, l} h_{n_r, l}(\sqrt{\frac{M\omega}{\hbar}} r) e^{-\frac{M\omega r^2}{2\hbar}} Y_l^m(\theta, \phi) / r$$

The  $n_r$  is the number of radial nodes

The  $h_{n,l}(\sqrt{\frac{M\omega}{\hbar}} r)/r = C r^l L_{n,l}^{l+1} \left( \frac{M\omega}{\hbar} r^2 \right)$   
 ~ Associated Laguerre function

Special Case: The "non-relativistic" Hydrogen atom  $V(r) = -\frac{e^2}{4\pi\epsilon_0 r}$

You can generalize to  $\text{He}^+$ ,  $\text{Li}^{2+}$ , ... by changing  $e^2 \rightarrow Ze^2$  in all subsequent equations.

The radial equation  $-\frac{\hbar^2}{2M} \frac{d^2 u}{dr^2} + \left( -\frac{e^2}{4\pi\epsilon_0 r} + \frac{\hbar^2 l(l+1)}{2Mr^2} \right) u(r) = E u$

See the book for the derivation. Similar to the Harmonic oscillator case.

What range of energies correspond to bound states?

The small  $r$  and large  $r$  turning points defined by the quadratic eq.  $-\frac{e^2}{4\pi\epsilon_0 r} + \frac{\hbar^2 l(l+1)}{2Mr^2} = E$

For a quick and dirty estimate of the "size" of an atom, ignore the  $Y^2$  term  
 $r \approx \frac{e^2}{4\pi\epsilon_0 E}$  (size scales like  $1/E$ )

For a given  $l$ , the smallest  $E$  must be larger than the value of  $V_{\text{eff}}$  at its minimum

$$V_{\text{eff}}' = 0 = \frac{e^2}{4\pi\epsilon_0 r_m^2} - \frac{\hbar^2 l(l+1)}{Mr_m^3} \quad r_m = \frac{\hbar^2 l(l+1) 4\pi\epsilon_0}{Me^2}$$

$$\begin{aligned} V_{\text{eff}}(r_m) &= -M \left( \frac{e^2}{4\pi\epsilon_0 \hbar c} \right)^2 \frac{1}{2l(l+1)} = -Mc^2 \underbrace{\left( \frac{e^2}{4\pi\epsilon_0 \hbar c} \right)^2}_{\alpha^2} \frac{1}{2l(l+1)} \\ &= -\frac{e^2}{4\pi\epsilon_0} \underbrace{\frac{1}{\left( \frac{4\pi\epsilon_0 \hbar c}{m e^2} \right)}}_{a_0} \frac{1}{2l(l+1)} \end{aligned}$$

$$\alpha = \text{fine structure constant} = \frac{e^2}{4\pi\epsilon_0 \hbar c} = 1/137.035\ 999 \dots$$

$$a_0 = \text{Bohr radius} = \frac{4\pi\epsilon_0 \hbar^2}{Me^2} = 0.529\ 177\ 21 \dots \times 10^{-10} \text{ m}$$

$$\alpha_C = \text{velocity unit} = 2.187691 \times 10^6 \text{ m/s}$$

$$a_0/\alpha_C = \text{time unit} = 2.418884 \times 10^{-17} \text{ s} = 24.19 \text{ attosec.}$$

See book for derivation of energies and wave functions.

$$E_{n_r, l} = -\frac{1}{2} M(\alpha_C)^2 / r^2 \quad \text{with } n = n_r + l + 1 \quad (n_r \text{ starts from 0})$$

$$\Psi_{n_r, l, m}^{(r, \theta, \phi)} = R_{n_r}(r) Y_l^m(\theta, \phi)$$

notice change in notation (standard due to history: Bohr)

$$R_{n_r}(r) = \frac{1}{r} \left( \frac{r}{a_0} \right)^{l+1} e^{-r/a_0} U\left(\frac{r}{a_0}\right)$$

$$U\left(\frac{r}{a_0}\right) = C_0 + C_1 \left(\frac{r}{a_0}\right) + C_2 \left(\frac{r}{a_0}\right)^2 + \dots$$

$$C_{j+1} = 2 \frac{j+l+1-n}{(j+1)(j+2l+2)} C_j$$

As you might expect, these functions are well known. They are also Associated Laguerre polynomials.

$$L_g(x) \equiv e^x \left( \frac{d}{dx} \right)^g (x^g e^{-x}) \quad L_{g-p}^{(p)}(x) = (-1)^p \frac{d^p}{dx^p} L_g(x)$$

$$\Psi_{n_r, l, m} = \left( \frac{2}{na_0} \right)^{3/2} \left( \frac{(n-l-1)!}{2n[(n+l)!]^3} \right)^{1/2} \left( \frac{2r}{na_0} \right)^l L_{n-l-1}^{2l+1} \left( \frac{2r}{na_0} \right) e^{-r/a_0} Y_l^m(\theta, \phi)$$

These are the normalized  $\Psi_{n_r, l, m}$   $\int \Psi_{n_r, l, m}^* \Psi_{n_r', l', m'} r^2 \sin\theta dr d\theta d\phi = \delta_{n_r, n_r'} \delta_{l, l'} \delta_{m, m'}$

Visualizing the  $\Psi_{n_r, l, m}$ .

Important notation: States with  $l=0$  called "s"-states,  $l=1$  are "p" states,  $l=2$  are "d" states,  $l=3$  are "f" states,  $l=4$  are "g" states..

## Important multiplicity and naming.

$n=1$	can only be $n_r=0, l=0$	1s state	(1 state)
$n=2$	can be $n_r=1, l=0$ or $n_r=0, l=1$	(2s and 2p)	(1+3=4 states)
$n=3$	can be $n_r=2, l=0$ ; $n_r=1, l=1$ ; $n_r=0, l=2$	(3s, 3p, 3d)	(1+3+5=9 states)
$n=4$	can be (3,0); (2,1); (1,2); (0,3)	(4s, 4p, 4d, 4f)	(1+3+5+7=16 states)

The degeneracy of level  $n$  is  $n^2$  spatial states (spin coming!)

Transitions caused by photons have selection rules. If light is linearly polarized along  $\hat{z}$   $\Delta m_{\text{fin}} = \Delta m_{\text{init}} \pm 1$ ,  $M_{\text{fin}} = M_{\text{init}}$ ,  $\Delta N \neq 0$   
 circularly polarized along  $\hat{z}$  " "  $m_{\text{fin}} = m_{\text{init}} + 1$  or  $m_{\text{init}} - 1$   
 electric depending on direction,  $\Delta N \neq 0$

These are called dipole transitions because the transition operator is proportional to  $e\hat{z}$  or  $e(x \pm iy)$  where  $e$  is charge.

The energy of the photon emitted or absorbed is

$$hf = |E_f - E_i| = E_{1,0} \left| \frac{1}{n_f^2} - \frac{1}{n_i^2} \right| = hc/\lambda$$

$$\frac{1}{\lambda} = R \left| \frac{1}{n_f^2} - \frac{1}{n_i^2} \right| \quad R = \frac{E_{1,0}}{hc} = 1.097 \times 10^7 \frac{1}{m} \quad \text{Rydberg constant}$$

Is  $n_f > n_i$  or  $n_f < n_i$  for photon emitted?  
 " " " " " " absorbed?

Can a photon be absorbed where  $hf > E_{1,0} \left( \frac{1}{n_i^2} - \frac{1}{\infty^2} \right)$ ?  
 What does it mean?

Show plots of different  $R_{n,a}$  and discuss trends.

Now do a formal approach to angular momenta.

$$\hat{L}_x = \hat{y}\hat{P}_z - \hat{z}\hat{P}_y \quad \hat{L}_y = \hat{z}\hat{P}_x - \hat{x}\hat{P}_z \quad \hat{L}_z = \hat{x}\hat{P}_y - \hat{y}\hat{P}_x$$

$$[\hat{L}_x, \hat{L}_y] = i\hbar \hat{L}_z \quad [\hat{L}_y, \hat{L}_z] = i\hbar \hat{L}_x \quad [\hat{L}_z, \hat{L}_x] = i\hbar \hat{L}_y$$

$$[\hat{L}^2, \hat{L}_x] = [\hat{L}^2, \hat{L}_y] = [\hat{L}^2, \hat{L}_z] = 0$$

These relations mean the eigenstates of  $\hat{L}^2$  can also be eigenstates of  $\hat{L}_z$  (Going to use slightly different notation from book.)

$$\hat{L}^2 |\Psi_{\lambda, \mu}\rangle = \hbar^2 \lambda(\lambda+1) |\Psi_{\lambda, \mu}\rangle \quad \hat{L}_z |\Psi_{\lambda, \mu}\rangle = \hbar^2 \mu^2 |\Psi_{\lambda, \mu}\rangle$$

$$\text{Why must } \langle \Psi_{\lambda, \mu} | \hat{L}^2 - \hat{L}_z^2 | \Psi_{\lambda, \mu} \rangle = \hbar^2 (\lambda(\lambda+1) - \mu^2) \geq 0 ?$$

Define the raising and lowering operators  $\hat{L}_{\pm} = \hat{L}_x \pm i\hat{L}_y$

$$[\hat{L}_z, \hat{L}_{\pm}] = [\hat{L}_z, \hat{L}_x] \pm i[\hat{L}_z, \hat{L}_y] = i\hbar \hat{L}_y \pm i\hbar \hat{L}_x = \pm \hbar (\hat{L}_x \pm i\hat{L}_y) = \pm \hbar \hat{L}_{\pm}$$

Use this to see how the raising/lowering works

Show  $|\Psi_{\lambda, \mu}\rangle$  is also eigenstate of  $\hat{L}^2$  and  $\hat{L}_z$

$$\begin{aligned} \hat{L}^2 \hat{L}_{\pm} |\Psi_{\lambda, \mu}\rangle &= \hat{L}_{\pm} \hat{L}^2 |\Psi_{\lambda, \mu}\rangle + [\hat{L}^2, \hat{L}_{\pm}] |\Psi_{\lambda, \mu}\rangle \\ &= \hbar^2 \lambda(\lambda+1) \hat{L}_{\pm} |\Psi_{\lambda, \mu}\rangle \xrightarrow{\text{The } \lambda \text{ doesn't change}} \end{aligned}$$

$$\begin{aligned} \hat{L}_z \hat{L}_{\pm} |\Psi_{\lambda, \mu}\rangle &= \hat{L}_{\pm} \hat{L}_z |\Psi_{\lambda, \mu}\rangle + [\hat{L}_z, \hat{L}_{\pm}] |\Psi_{\lambda, \mu}\rangle \\ &= \hbar \mu \hat{L}_{\pm} |\Psi_{\lambda, \mu}\rangle \pm \hbar \hat{L}_{\pm} |\Psi_{\lambda, \mu}\rangle = \hbar(\mu \pm 1) \hat{L}_{\pm} |\Psi_{\lambda, \mu}\rangle \end{aligned}$$

$$\text{This means } \hat{L}_{\pm} |\Psi_{\lambda, \mu}\rangle = \hbar A_{\lambda, \mu}^{\pm} |\Psi_{\lambda, \mu \pm 1}\rangle$$

The eigenvalues of  $\hat{L}_z$  are a series of values separated by  $\hbar$ .

We can find the  $A_{\lambda, \mu}^{\pm}$  using a similar procedure to when we found  $\hat{a}_n |\Psi_n\rangle = C_n |\Psi_{n+1}\rangle$  for the Harmonic oscillator.

$$\hat{L}_+ |\Psi_{\lambda, \mu}\rangle = \hbar A_{\lambda, \mu}^+ |\Psi_{\lambda, \mu+1}\rangle \quad A_{\lambda, \mu}^+ = \frac{1}{\hbar} \langle \Psi_{\lambda, \mu+1} | \hat{L}_+ | \Psi_{\lambda, \mu} \rangle$$

Write as Hermitian Conj.  $A_{\lambda, \mu}^+ = \frac{1}{\hbar} \langle \hat{L}_+^\dagger \Psi_{\lambda, \mu+1} | \Psi_{\lambda, \mu} \rangle$

Do Hermitian Conj.  $(\hat{L}_x + i\hat{L}_y)^+ = \hat{L}_x^+ - i\hat{L}_y^+ = \hat{L}_x - i\hat{L}_y = \hat{L}_-$

$$\text{Now use } L_- \quad A_{\lambda, \mu}^- = \frac{1}{\hbar} \langle \hat{L}_- \Psi_{\lambda, \mu+1} | \Psi_{\lambda, \mu} \rangle = A_{\lambda, \mu+1}^+$$

The "trick" is to do  $L_+$  then  $L_-$  so you get back

$$\hat{L}_- \hat{L}_+ |\Psi_{\lambda, \mu}\rangle = \hbar A_{\lambda, \mu}^+ \hat{L}_- |\Psi_{\lambda, \mu+1}\rangle = \hbar^2 A_{\lambda, \mu}^+ A_{\lambda, \mu+1}^- |\Psi_{\lambda, \mu}\rangle = \hbar^2 (A_{\lambda, \mu}^+)^2 |\Psi_{\lambda, \mu}\rangle$$

The  $|\Psi_{\lambda, \mu}\rangle$  is an eigenstate of  $L_- L_+$ . What is  $L_- L_+ ???$

$$\begin{aligned} \hat{L}_- \hat{L}_+ &= (\hat{L}_x - i\hat{L}_y)(\hat{L}_x + i\hat{L}_y) = \hat{L}_x^2 + \hat{L}_y^2 + i(\hat{L}_x \hat{L}_y - \hat{L}_y \hat{L}_x) \\ &= \hat{L}_x^2 + \hat{L}_y^2 - \hbar \hat{L}_z = \hat{L}^2 - (\hat{L}_z^2 + \hbar^2 \hat{L}_z) \end{aligned}$$

Since the  $|\Psi_{\lambda, \mu}\rangle$  is an eigenstate of  $\hat{L}^2$  and  $\hat{L}_z$ , we're done

$$\hat{L}_- \hat{L}_+ |\Psi_{\lambda, \mu}\rangle = [\hbar^2 \lambda(\lambda+1) - \hbar^2 \mu(\mu+1)] |\Psi_{\lambda, \mu}\rangle$$

$$A_{\lambda, \mu}^\pm = \sqrt{\lambda(\lambda+1) - \mu(\mu \pm 1)}$$

Remember the point! This lets us get the different eigenstates.

$$|\Psi_{\lambda, \mu \pm 1}\rangle = \frac{1}{\hbar A_{\lambda, \mu}^\pm} \hat{L}_\pm |\Psi_{\lambda, \mu}\rangle$$

It also lets us find the allowed values of  $\lambda, \mu$ . The state that corresponds to the maximum  $\mu$  must have

$$\hat{L}_+ |\Psi_{\lambda, \mu_{\max}}\rangle = 0 = \hbar A_{\lambda, \mu_{\max}}^+ |\Psi_{\lambda, \mu+1}\rangle$$

This gives  $\mu_{\max} = \lambda$ . Similar treatment of  $\mu_{\min}$  gives  $\mu_{\min} = -\lambda$ .

Since  $\lambda_{\min} = -\lambda_{\max}$  and  $\lambda_{\max} - \lambda_{\min} = \text{integer}$ , the

$$\lambda = \lambda_{\max} = \text{integer}/2$$

Allowed values  $(\lambda=0, m=0); (\lambda=1, m=-1, 0, 1); (\lambda=2, m=-2, -1, 0, 1, 2) \dots$

These are just the values we found earlier, but we use  $\lambda, m$

Representation: For  $\lambda, m = \text{integer}$   $\langle \theta, \phi | \Psi_{\ell, m} \rangle \equiv Y_m^{\ell}(\theta, \phi)$

But we've missed some. The condition was  $\lambda = \frac{\text{integer}}{2}$   
What do we do with  $\lambda = \frac{1}{2}, \frac{3}{2}, \frac{5}{2}, \dots$ .

Allowed values:  $(\lambda = \frac{1}{2}, m = -\frac{1}{2}, \frac{1}{2}), (\lambda = \frac{3}{2}, m = -\frac{3}{2}, -\frac{1}{2}, \frac{1}{2}, \frac{3}{2}) \dots$

We did not find these cases when we did the differential equation.  
They are not allowed for orbital angular momentum. All  
the different  $\lambda$  are allowed for spins.

$[\hat{S}_a, \hat{S}_b] = i\hbar \epsilon_{abc} \hat{S}_c \Rightarrow$  Same logic will give allowed values

$$\hat{S}^2 |S, m_s\rangle = \hbar^2 s(s+1) |S, m_s\rangle \quad \text{and} \quad \hat{S}_z |S, m_s\rangle = \hbar m_s |S, m_s\rangle$$

$$\hat{S}_{\pm} |S, m_s\rangle = \hbar \sqrt{s(s+1) - m_s(m_s \pm 1)} |S, m_s \pm 1\rangle$$

For spins, the representation of  $\hat{S}_a$  and states  $|S, m_s\rangle$  are taken to be  $(2s+1) \times (2s+1)$  matrices and  $2s+1$  length vector.

Every particle has a well defined spin. Electrons, quarks  $S = \frac{1}{2}$   
photons, gluons  $S = 1$

Composite particles can have several  $S$  depending on state.

H-atom ground state (proton  $[S = \frac{1}{2}]$  and electron  $[S = \frac{1}{2}]$ )  $S_{\text{tot}} = 0 \text{ or } 1$   
" "  $n=2$  states  $S_{\text{tot}} = 0 \text{ or } 1 \text{ or } 2$

Representation of spin  $\frac{1}{2}$   $|1\rangle = |\frac{1}{2}, \frac{1}{2}\rangle$  and  $|2\rangle = |\frac{1}{2}, -\frac{1}{2}\rangle$

$$\langle 1 | \hat{S}^2 | 1 \rangle = \hbar^2 \frac{1}{2} \frac{3}{2} \langle 1 | 1 \rangle = \frac{3}{4} \hbar^2 \quad \langle 1 | \hat{S}^2 | 2 \rangle = \frac{3}{4} \hbar^2 \langle 1 | 2 \rangle = 0$$

$$\langle 2 | \hat{S}^2 | 1 \rangle = \frac{3}{4} \hbar^2 \langle 2 | 1 \rangle = 0 \quad \langle 2 | \hat{S}^2 | 2 \rangle = \frac{3}{4} \hbar^2 \langle 2 | 2 \rangle = \frac{3}{4} \hbar^2$$

$$\hat{S}^2 = \frac{3}{4} \hbar^2 \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \quad \text{Representation of } \hat{S}^2 \text{ is } \hbar^2 S(S+1) \underline{\underline{I}}$$

$$\langle 1 | \hat{S}_z | 1 \rangle = \frac{1}{2} \hbar \langle 1 | 1 \rangle = \frac{\hbar}{2} \quad \langle 1 | \hat{S}_z | 2 \rangle = -\frac{\hbar}{2} \langle 1 | 2 \rangle = 0$$

$$\langle 2 | \hat{S}_z | 1 \rangle = \frac{\hbar}{2} \langle 2 | 1 \rangle = 0 \quad \langle 2 | \hat{S}_z | 2 \rangle = -\frac{\hbar}{2} \langle 2 | 2 \rangle = -\frac{\hbar}{2}$$

$$\hat{S}_z = \frac{\hbar}{2} \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \quad \text{Representation of } \hat{S}_z \text{ is diagonal matrix}$$

$$\langle 1 | \hat{S}_+ | 1 \rangle = \langle 1 | 0 | 1 \rangle = 0 \quad \langle 1 | \hat{S}_+ | 2 \rangle = \hbar \sqrt{\frac{3}{4} - (-\frac{1}{2})\frac{1}{2}} \langle 1 | 1 \rangle = \hbar$$

$$\langle 2 | \hat{S}_+ | 1 \rangle = \langle 2 | 0 | 1 \rangle = 0 \quad \langle 2 | \hat{S}_+ | 2 \rangle = \hbar \langle 2 | 1 \rangle = 0$$

$$\hat{S}_+ = \hbar \begin{pmatrix} 0 & 1 \\ 0 & 0 \end{pmatrix} \quad \hat{S}_+ = \hat{S}_- \Rightarrow \hat{S}_- = \hbar \begin{pmatrix} 0 & 0 \\ 1 & 0 \end{pmatrix}$$

$$\hat{S}_x = \frac{1}{2} (\hat{S}_+ + \hat{S}_-) \quad \hat{S}_x = \frac{\hbar}{2} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$$

$$\hat{S}_y = -\frac{i}{2} (\hat{S}_+ - \hat{S}_-) \quad \hat{S}_y = \frac{\hbar}{2} \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}$$

The  $\hat{S}_x$ ,  $\hat{S}_y$ ,  $\hat{S}_z$  are written in terms of the Pauli spin matrices when  $S = \frac{1}{2}$

$$\hat{S}_i = \frac{\hbar}{2} \underline{\underline{\sigma}}_i \quad \sigma_x = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \quad \sigma_y = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix} \quad \sigma_z = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

Every Pauli Spin matrix has the property  $\underline{\underline{\sigma}}^2 = \underline{\underline{I}}$

$$\hat{S}^2 = \hat{S}_x^2 + \hat{S}_y^2 + \hat{S}_z^2 = \left(\frac{\hbar}{2}\right)^2 \underline{\underline{\sigma}}_x^2 + \left(\frac{\hbar}{2}\right)^2 \underline{\underline{\sigma}}_y^2 + \left(\frac{\hbar}{2}\right)^2 \underline{\underline{\sigma}}_z^2 = \frac{3}{4} \hbar^2 \underline{\underline{I}}$$

The eigenvalues of  $\underline{\underline{\sigma}}_z$  are  $\pm 1$ . By symmetry the eigenvectors of  $\underline{\underline{\sigma}}_x$  and  $\underline{\underline{\sigma}}_y$  must also be  $\pm 1$ .

Example: Find the eigenvalues of  $\hat{S}_y$

$$\hat{S}_y |\Psi_\alpha\rangle = g_\alpha |\Psi_\alpha\rangle \quad \frac{\hbar}{2} \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix} \begin{pmatrix} \Psi_{1\alpha} \\ \Psi_{2\alpha} \end{pmatrix} = g_\alpha \begin{pmatrix} \Psi_{1\alpha} \\ \Psi_{2\alpha} \end{pmatrix}$$

$$-\frac{i\hbar}{2} \Psi_{2\alpha} = g_\alpha \Psi_{1\alpha} \quad \text{and} \quad i\frac{\hbar}{2} \Psi_{1\alpha} = g_\alpha \Psi_{2\alpha} = \left(\frac{g_\alpha^2}{g_\alpha^2 + \frac{\hbar^2}{4}}\right) \Psi_{1\alpha}$$

$$g_\alpha^2 = \frac{i\hbar}{2} \left(-\frac{i\hbar}{2}\right) = \frac{\hbar^2}{4} \quad \Rightarrow \quad g_+ = \frac{\hbar}{2} \quad g_- = -\frac{\hbar}{2}$$

$$-\frac{i\hbar}{2} \Psi_{2+} = \frac{\hbar}{2} \Psi_{1+} \quad \Rightarrow \quad \Psi_+ = \begin{pmatrix} \sqrt{2} \\ i/\sqrt{2} \end{pmatrix} \quad \langle \Psi_- | \Psi_+ \rangle = \left(\frac{1}{\sqrt{2}} \frac{i}{\sqrt{2}}\right) \begin{pmatrix} \sqrt{2} \\ i/\sqrt{2} \end{pmatrix}$$

$$-\frac{i\hbar}{2} \Psi_{2-} = -\frac{\hbar}{2} \Psi_{1-} \quad \Rightarrow \quad \Psi_- = \begin{pmatrix} \sqrt{2} \\ -i/\sqrt{2} \end{pmatrix} \quad = \frac{1}{2} - \frac{1}{2} = 0$$

Make sure look at eigenvalues and eigenvectors of  $\hat{S}_x$  and  $\hat{S}_z$

Example: Spin precession in a magnetic field

Remember that the energy for a dipole moment in a magnetic field is  $U = -\vec{\mu} \cdot \vec{B}$  and  $\vec{\mu} \propto \vec{s}/m$

$$H = -\gamma \vec{B} \cdot \vec{s}$$

We're free to pick the axis how we want so pick them where  
 $\vec{B} = B_0 \hat{z}$

$$H = -\gamma B_0 S_z \quad \Rightarrow \quad i\hbar \frac{\partial}{\partial t} \begin{pmatrix} \chi_1 \\ \chi_2 \end{pmatrix} = \begin{pmatrix} -\gamma B_0 \frac{\hbar}{2} & 0 \\ 0 & \gamma B_0 \frac{\hbar}{2} \end{pmatrix} \begin{pmatrix} \chi_1 \\ \chi_2 \end{pmatrix}$$

$$i\hbar \frac{d\chi_1}{dt} = -\gamma B_0 \frac{\hbar}{2} \chi_1 \quad \Rightarrow \quad \chi_1(t) = e^{i\gamma B_0 \frac{\hbar}{2} t} \chi_1(0)$$

$$i\hbar \frac{d\chi_2}{dt} = \gamma B_0 \frac{\hbar}{2} \chi_2 \quad \Rightarrow \quad \chi_2(t) = e^{-i\gamma B_0 \frac{\hbar}{2} t} \chi_2(0)$$

The eigenvalues are  $-\frac{\gamma B_0 \hbar}{2}$  and  $\frac{\gamma B_0 \hbar}{2}$

To understand what this is trying to tell us, let's examine the most general state

$$\chi(t) = e^{i\varphi} \begin{pmatrix} \cos(\alpha_z) \\ \sin(\alpha_z) e^{i\beta} \end{pmatrix}$$

Does the  $e^{i\varphi}$  change anything?

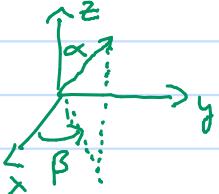
$$\langle \hat{S}_z \rangle = \chi^+ \underline{\underline{S}_z} \chi = (\cos(\frac{\alpha}{2}) \sin(\frac{\alpha}{2}) e^{-i\beta}) \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \begin{pmatrix} \cos(\frac{\alpha}{2}) \\ \sin(\frac{\alpha}{2}) e^{-i\beta} \end{pmatrix}$$

$$= \cos^2(\frac{\alpha}{2}) - \sin^2(\frac{\alpha}{2}) = \cos(\alpha)$$

$$\langle \hat{S}_x \rangle = (\cos(\frac{\alpha}{2}) \sin(\frac{\alpha}{2}) e^{-i\beta}) \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} \cos(\frac{\alpha}{2}) \\ \sin(\frac{\alpha}{2}) e^{i\beta} \end{pmatrix}$$

$$= \cos(\frac{\alpha}{2}) \sin(\frac{\alpha}{2}) (e^{i\beta} + e^{-i\beta}) = \sin(\alpha) \cos(\beta)$$

$$\langle \hat{S}_y \rangle = \cos(\frac{\alpha}{2}) \sin(\frac{\alpha}{2}) (-i e^{i\beta} + i e^{-i\beta}) = \sin(\alpha) \sin(\beta)$$



Now add the time dependence

$$\chi(t) = e^{i\gamma} \begin{pmatrix} \cos(\frac{\alpha}{2}) e^{i\gamma B_0 t/2} \\ \sin(\frac{\alpha}{2}) e^{i\beta} e^{-i\gamma B_0 t/2} \end{pmatrix} = e^{i(\gamma + \gamma B_0 t/2)} \begin{pmatrix} \cos(\frac{\alpha}{2}) \\ \sin(\frac{\alpha}{2}) e^{i(\beta - \gamma B_0 t)} \end{pmatrix}$$

$$\langle S_z \rangle = \cos(\alpha) \quad \langle S_x \rangle = \sin(\alpha) \cos(\beta - \gamma B_0 t) \quad \langle S_y \rangle = \sin(\alpha) \sin(\beta - \gamma B_0 t)$$

Why doesn't  $\langle S_z \rangle$  change? Why doesn't tilt change?  
What does this motion look like?

Look at the book discussion of Stern Gerlach experiment.

How to write the state when there are 2 or more particles with spin?

Two particles  $S_1, S_2$  will have joint states

$$|S_1 m_{S_1}, S_2 m_{S_2}\rangle = |S_1 m_{S_1}\rangle |S_2 m_{S_2}\rangle$$

These are rarely eigenstates of  $\hat{H}$

One of the more common forms of interaction is proportional to  $\vec{S}_1 \cdot \vec{S}_2$

This can be written as  $\vec{S}_1 \cdot \vec{S}_2 = \frac{1}{2} [(\vec{S}_1 + \vec{S}_2)^2 - \vec{S}_1^2 - \vec{S}_2^2]$

Instead of using eigenstates of  $\vec{S}_1^2, S_{1z}, \vec{S}_2^2, S_{2z}$  use the eigenstates of  $\vec{S}_1^2, \vec{S}_2^2, \vec{S}^2, S_z$

$$\vec{S} \equiv \vec{S}_1 + \vec{S}_2 \quad S_z = S_{1z} + S_{2z}$$

In classical mechanics, the  $\vec{S}_1$  and  $\vec{S}_2$  can be in any direction. The  $\max |\vec{S}_1 + \vec{S}_2| = |\vec{S}_1| + |\vec{S}_2|$  and the  $\min |\vec{S}_1 + \vec{S}_2| = ||\vec{S}_1| - |\vec{S}_2||$

In Q.M. the allowed eigenvalues of  $\vec{S}^2$  are  $\hbar^2 s(s+1)$  with  $s_1 + s_2, s_1 + s_2 - 1, \dots, |s_1 - s_2|$

Example: Spin  $\frac{1}{2}$  and spin  $\frac{3}{2}$  Count states

Total number of $\frac{5}{2}$ states	$2 \cdot \frac{5}{2} + 1 = 6$
" " "	$2 \cdot \frac{3}{2} + 1 = 4$
" " "	states $= 6 \cdot 4 = 24$

Total number of $\frac{5}{2} + \frac{3}{2} = 4$ states	$2 \cdot 4 + 1 = 9$	sum = 24 states
" " "	$2 \cdot \frac{3}{2} + 1 = 7$	
" " "	$2 \cdot \frac{1}{2} + 1 = 5$	
" " "	$2 \cdot (-\frac{1}{2}) + 1 = 3$	

Strategy for finding all of the eigenstates of  $\vec{S}^2$  and  $S_z$

- ① The state with  $M_s = s_1$  and  $M_s = s_2$  must be  $S = S_1 + S_2, M_s = s$
- ② Act on that state with  $\hat{S}_- / (\hbar A_{sm_s})$ . This must be  $S = S_1 + S_2, M_s = s-1$
- ③ Do  $\hat{S}_- / (\hbar A_{sm_{s-1}})$  This must be  $S = S_1 + S_2, M_s = s-2$   
etc until you've got  $S = S_1 + S_2, M_s = -s$
- ④ The  $S = S_1 + S_2 - 1, M_s = s_1 + s_2 - 1$  state must be the state orthogonal to the  $S = S_1 + S_2, M_s = s_1 + s_2 - 1$  state
- ⑤ Apply the  $\hat{S}_- / (\hbar A_{sm_s})$  to this state to get  $S = S_1 + S_2 - 1, M_s = s_1 + s_2 - 2$   
etc
- ⑥ The  $S = S_1 + S_2 - 2, M_s = s_1 + s_2 - 2$  must be orthogonal to the states  $S = S_1 + S_2, M_s = s_1 + s_2 - 2$  and  $S = S_1 + S_2 - 1, M_s = s_1 + s_2 - 2$  states.  
etc

Example: Spin 1 and spin  $\frac{1}{2}$

$$|\frac{3}{2}, \frac{3}{2}\rangle = |1,1\rangle |\frac{1}{2}, \frac{1}{2}\rangle$$

$$|\frac{3}{2}, \frac{1}{2}\rangle = \frac{1}{\sqrt{2}}(S_{1z} + S_{2z}) |1,1\rangle |\frac{1}{2}, \frac{1}{2}\rangle / \sqrt{(\frac{1}{4} - \frac{3}{4})} = \sqrt{\frac{1}{3}} |1,0\rangle |\frac{1}{2}, \frac{1}{2}\rangle + \sqrt{\frac{1}{3}} |1,1\rangle |\frac{1}{2}, \frac{1}{2}\rangle$$

:

$$|\frac{1}{2}, \frac{1}{2}\rangle = -\sqrt{\frac{1}{3}} |1,0\rangle |\frac{1}{2}, \frac{1}{2}\rangle + \sqrt{\frac{2}{3}} |1,1\rangle |\frac{1}{2}, \frac{1}{2}\rangle$$

You can look up these coefficients

$$|S, M_S\rangle = \sum_{M_1 M_2 S} |S_1 M_{1s}\rangle |S_2 M_{2s}\rangle \langle S_1 M_{1s} S_2 M_{2s} | S M_S \rangle \quad M_{1s} + M_{2s} = M_S$$

The  $\langle j_1 M_1 j_2 M_2 | j M \rangle$  are Clebsch Gordon coefficients

$$\text{Book notation } C_{m_1 m_2 m}^{s_1 s_2 s} = \langle S_1 M_1 S_2 M_2 | S M \rangle$$

Show how to read the C.G. table

Example:  $S_1 = \frac{3}{2} \quad S_2 = 1$

$$|S_1, S_2\rangle = |\frac{3}{2}, \frac{3}{2}\rangle |1, 1\rangle$$

$$|\frac{5}{2}, \frac{3}{2}\rangle = \sqrt{\frac{1}{3}} |\frac{3}{2}, \frac{3}{2}\rangle |1, 0\rangle + \sqrt{\frac{3}{5}} |\frac{3}{2}, \frac{1}{2}\rangle |1, 1\rangle$$

$$|\frac{3}{2}, \frac{3}{2}\rangle = \sqrt{\frac{3}{5}} |\frac{3}{2}, \frac{3}{2}\rangle |1, 0\rangle - \sqrt{\frac{2}{5}} |\frac{3}{2}, \frac{1}{2}\rangle |1, 1\rangle$$

$$|\frac{5}{2}, \frac{1}{2}\rangle = \sqrt{\frac{1}{10}} |\frac{3}{2}, \frac{3}{2}\rangle |1, -1\rangle + \sqrt{\frac{6}{10}} |\frac{3}{2}, \frac{1}{2}\rangle |1, 0\rangle + \sqrt{\frac{3}{10}} |\frac{3}{2}, -\frac{1}{2}\rangle |1, 1\rangle$$

$$|\frac{3}{2}, \frac{1}{2}\rangle = \sqrt{\frac{6}{15}} |\frac{3}{2}, \frac{3}{2}\rangle |1, -1\rangle + \sqrt{\frac{1}{15}} |\frac{3}{2}, \frac{1}{2}\rangle |1, 0\rangle - \sqrt{\frac{8}{15}} |\frac{3}{2}, -\frac{1}{2}\rangle |1, 1\rangle$$

$$|\frac{1}{2}, \frac{1}{2}\rangle = \sqrt{\frac{3}{6}} |\frac{3}{2}, \frac{3}{2}\rangle |1, -1\rangle - \sqrt{\frac{2}{6}} |\frac{3}{2}, \frac{1}{2}\rangle |1, 0\rangle + \sqrt{\frac{1}{6}} |\frac{3}{2}, -\frac{1}{2}\rangle |1, 1\rangle$$

Schematic of fine structure and hyperfine structure in H

$$\begin{array}{c} 2s \\ \hline \end{array} \xrightarrow{F=1} \begin{array}{c} 2p \\ \hline F=0 \end{array} \quad \begin{array}{c} \hline I=\frac{3}{2} \\ \hline \end{array} \quad \begin{array}{c} \hline I=\frac{1}{2} \\ \hline \end{array} \quad \begin{array}{c} \hline F=2 \\ F=1 \\ \hline \end{array} \quad \begin{array}{c} \hline F=1 \\ F=0 \\ \hline \end{array}$$

$$\vec{J} = \vec{L} + \vec{S}_e \quad \text{Fine}$$

$$\vec{F} = \vec{J} + \vec{S}_p \quad \text{Hyperfine}$$

$$1s \xrightarrow{-F=1} \boxed{1420 \text{ MHz}} \rightarrow 21 \text{ cm line}$$

astrophysics

Example : Energy splittings when  $\hat{H} = \varepsilon_0 \vec{S}_1 \cdot \vec{S}_2 / \hbar^2$

$$\vec{S}_1 \cdot \vec{S}_2 = \frac{1}{2} (\vec{S}^2 - \vec{S}_1^2 - \vec{S}_2^2)$$

$$E_s = \frac{\varepsilon_0}{2} [s(s+1) - s_1(s_1+1) - s_2(s_2+1)]$$

$$s = |s_1 - s_2|, |s_1 - s_2| + 1, \dots \\ \dots, s_1 + s_2 - 1, s_1 + s_2$$

For H ground state  $s=1$  and  $0$

$$E_1 - E_0 = (\varepsilon_0/2) [1 \cdot 2 - \frac{3}{4} - \frac{3}{4}] - (\varepsilon_0/2) [0 \cdot 1 - \frac{3}{4} - \frac{3}{4}] = \varepsilon_0 = h \cdot 1420 \text{ MHz}$$