Multiple choice (5 pts each) – circle the correct answers.

- 1. Suppose the 4 electrons in Be all have the same $m_s = 1/2$. Which is the lowest energy allowed configuration under this condition?
 - (a) $1s^4$
 - (b) $1s^2 2s^2$
 - (c) $1s2s2p^2$
 - (d) 1s2s2p3s
- 2. A particle moves in 1D. For x < 0, the potential energy is 0. For x > 0, the potential energy is 1 eV. If the particle has a total energy of 10 eV, which statement(s) are correct?
 - (a) the wave function is discontinuous at x = 0.
 - (b) the de Broglie wavelength for x < 0 is greater than that for x > 0.
 - (c) for x > 0, the wave function exponentially decreases towards ∞ : $\psi(x) = Ae^{-\kappa x}$.
 - (d) for x < 0, the wave function exponentially decreases towards $-\infty$: $\psi(x) = Ae^{\kappa x}$.
 - (e) none of the above are correct.
- 3. For this question, don't worry about the spin of the electron. Circle all of the *impossible* states of hydrogen (n, l, m_l) .

(a)
$$(5,0,0)$$
 (b) $(2,1,-1)$ (c) $(5,5,2)$
(d) $(3,2,3)$ (e) $(1,0,0)$ (f) $(50,3,0)$

- 4. The ${}_{1}^{3}H_{2}$ nucleus has a half life of 12 years. You make one ${}_{1}^{3}H_{2}$ when you are 25 years old. How old will you be when it radioactively decays?
 - (a) 37 years old
 - (b) 49 years old
 - (c) 61 years old
 - (d) When it decays can not be predicted.
 - (e) none of the above
- 5. A hydrogen atom in its ground state absorbs a photon and goes to a free electron and proton. The minimum photon energy for this process is 13.6 eV. For the same process in muonic hydrogen (a bound muon and proton in the ground state absorbs a photon and goes to a free muon and proton), what is the minimum photon energy required?
 - (a) more than 13.6 eV.
 - (b) less than 13.6 eV.
 - (c) 13.6 eV.

Problems: Show all work to receive full credit. Fundamental constants (e, h, masses, etc) are top of the equation page.

10 pts 1. ${}^{8}_{4}Be_{4}$ decays by emission of an alpha particle. a) What is the total energy released? b) What is the speed of the alpha particle?

10 pts 2. A hydrogen atom in the ground state starts at rest where the magnetic field points in the z-direction with magnitude 2.00 T. The magnetic field gives an interaction energy of $U = -\vec{\mu} \cdot \vec{B}$. It accelerates to a spot 1 cm away where the magnetic field points in the z-direction with a magnitude 2.60 T. a) What is the value of the spin magnetic quantum number? b) What is the kinetic energy of the hydrogen atom?

10 pts 3. In an atom, there are two electrons with $m_s = 1/2$. One is in the 2*p* state and the other is in the 3*d* state. a) What is the total number of two electron states with these conditions? b) Suppose the two angular momenta add to give a total orbital angular momentum *L*. What are the allowed values of *L*? c) How many states are there for each *L*?

10 pts 4. A proton is fired directly at a ${}^{235}_{92}U_{143}$ nucleus. What minimum energy does the proton need to reach the nucleus?

15 pts 5. A particle of mass m experiences a potential which is $U(x) = U_0 > 0$ for x < 0 and $U(x) = -U_0$ for x > 0. For $-U_0 < E < U_0$, determine the wave function for all x. (Hint: you should probably break it up into a solution for x < 0 and a solution for x > 0; then use the correct rules to connect the two pieces.)

$$g = 9.80 \frac{m}{s^2} \qquad h = 6.63 \times 10^{-34} \ J \ s = 4.14 \times 10^{-15} \ eV \ s \qquad \hbar = \frac{h}{2\pi} \qquad c = 3.00 \times 10^8 \ \frac{m}{s}$$
$$M_{elec} = 9.11 \times 10^{-31} \ kg \qquad M_{prot} = 1.67 \times 10^{-27} \ kg \qquad M_{muon} = 1.88 \times 10^{-28} \ kg$$
$$1.60 \times 10^{-19} \ J = 1 \ eV \qquad e = 1.60 \times 10^{-19} \ C \qquad 1/(4\pi\varepsilon_0) = 8.99 \times 10^9 \ N \ m^2/C^2$$

Galilean relativity	$x' = x - ut, v'_x = v_x - u$	2.1	Lorentz velocity	$v_x' = \frac{v_x - u}{1 - v_x u/c^2},$	2.5
Einstein's postulates	(1) The laws of physics are the same in all inertial frames. (2) The speed of light has the same value c in all inertial frames.	2.3	transformation	$v'_{y} = \frac{v_{y}\sqrt{1 - u^{2}/c^{2}}}{1 - v_{x}u/c^{2}},$ $v'_{z} = \frac{v_{z}\sqrt{1 - u^{2}/c^{2}}}{1 - v_{x}u/c^{2}}$	
Time dilation	$\Delta t = \frac{\Delta t_0}{\sqrt{1 - u^2/c^2}}$ (Δt_0 = proper time)	2.4	Clock synchronization	$\Delta t' = \frac{uL/c^2}{\sqrt{1 - u^2/c^2}}$	2.5
Length contraction	$L = L_0 \sqrt{1 - u^2/c^2}$ (L_0 = proper length)	2.4	Relativistic momentum	$\vec{\mathbf{p}} = \frac{m\vec{\mathbf{v}}}{\sqrt{1 - v^2/c^2}}$	2.7
Velocity addition	$v = \frac{v' + u}{1 + v' u/c^2}$	2.4	Relativistic kinetic energy	$K = \frac{mc^2}{\sqrt{1 - v^2/c^2}} - mc^2$	2.7
velocity addition	$v = \frac{1}{1 + v' u/c^2}$	2.7	Rest energy	$E_0 = mc^2$	2.7
Doppler effect (source and	$f' = f \sqrt{\frac{1 - u/c}{1 + u/c}}$	2.4	Relativistic total energy	$E = K + E_0 = \frac{mc^2}{\sqrt{1 - v^2/c^2}}$	2.7
observer separating)	$\int \sqrt{1+u/c}$		Momentum-energy relationship	$E = \sqrt{(pc)^2 + (mc^2)^2}$	2.7
Lorentz transformation	$x'=\frac{x-ut}{\sqrt{1-u^2/c^2}},$	2.5	Extreme relativistic approximation	$E \cong pc$	2.7
	y' = y, z' = z, $t' = \frac{t - (u/c^2)x}{\sqrt{1 - u^2/c^2}}$		Conservation laws	In an isolated system of particles, the total momentum and the relativistic total energy remain constant.	2.8
Hubble's law	$v = H_0 d$	15.1	Age of matter-		
Number density	N/V =	15.2	dominated	$t = 1/\sqrt{6\pi G \rho_{\rm m}}$	15.7

dominated universe

Temperature of universe at

Fraction of photons above

age t

 E_0

Number density of photons	$N/V = (2.03 \times 10^7 \text{photons/m}^3 \cdot \text{K}^3)T^3$	15.2
Energy density of photons	$U = (4.72 \times 10^3 \mathrm{eV/m^3 \cdot K^4})T^4$	15.2
Gravitational frequency change	$\Delta f/f = gH/c^2$	15.4
Deflection of starlight	$\theta = 2GM/Rc^2$	15.5
Perihelion precession	$\Delta\phi = \frac{6\pi GM}{c^2 r_{\min}(1+e)}$	15.5
Schwarzschild radius	$r_{\rm S} = 2GM/c^2$	15.6

dominated universe	$t = 1/\sqrt{6\pi G \rho_{\rm m}}$	15.7
Age of radiation-		

$$t = \sqrt{3/32\pi G\rho_{\rm r}}$$
 15.7

$$T = \frac{1.5 \times 10^{10} \,\mathrm{s}^{1/2} \cdot \mathrm{K}}{t^{1/2}}$$
 15.8

$$f = 0.42e^{-E_0/kT} \times \left[\left(\frac{E_0}{kT}\right)^2 + 2\left(\frac{E_0}{kT}\right) + 2 \right]$$

$$3H^2$$
15.9

Critical density
$$\rho_{cr} = \frac{3H}{8\pi G}$$

of universe $= 0.97 \times 10^{-26} \text{ kg/m}^3$ 15.10

Double-slit maxima	$y_n = n \frac{\lambda D}{d} n = 0, 1, 2, 3, \dots$	3.1	Rayleigh-Jeans $I(\lambda) = \frac{2\pi c}{\lambda^4} kT$ 3.5 formula	.3
Bragg's law for X-ray diffraction	$2d\sin\theta=n\lambda n=1,2,3,\cdots$	3.1	Planck's blackbody $I(\lambda) = \frac{2\pi hc^2}{\lambda^5} \frac{1}{e^{hc/\lambda kT} - 1}$ 3.2 distribution	.3
Energy of photon	$E = hf = hc/\lambda$	3.2	Compton $\frac{1}{\pi} - \frac{1}{\pi} = \frac{1}{-2}(1 - \cos\theta),$ 3.4	4
Maximum kinetic energy of photoelectrons	$K_{\rm max} = eV_{\rm s} = hf - \phi$	3.2	Compton scattering $\frac{1}{E'} - \frac{1}{E} = \frac{1}{m_e c^2} (1 - \cos \theta), \qquad 3.4$ $\lambda' - \lambda = \frac{h}{m_e c} (1 - \cos \theta)$.4
Cutoff	$\lambda_{ m c}=hc/\phi$	3.2	Bremsstrahlung $\lambda_{\min} = hc/K = hc/e\Delta V$ 3.5	.5
wavelength			Pair production $hf = E_+ + E = 3.1$.5
Stefan's law	$I = \sigma T^4$	3.3	$(m_{\rm e}c^2 + K_{+}) + (m_{\rm e}c^2 + K_{-})$	
Wien's displacement law	$\lambda_{\max}T = 2.8978 \times 10^{-3} \mathrm{m} \cdot \mathrm{K}$	3.3	Electron-positron $(m_ec^2 + K_+) + (m_ec^2 + K)$ 3.3 annihilation $= E_1 + E_2$.5

De Broglie wavelength Single slit diffraction	$\lambda = h/p$ $a\sin\theta = n\lambda \ n = 1, 2, 3, \dots$	4.1 4.2	Statistical momentum uncertainty	$\Delta p_x = \sqrt{(p_x^2)_{\rm av} - (p_{x,\rm av})^2}$	4.4
Classical position-wavelength	$\Delta x \Delta \lambda \sim \varepsilon \lambda^2$	4.3	Wave packet (discrete k)	$y(x) = \sum A_t \cos k_t x$	4.5
uncertainty			Wave packet	$y(x) = \int A(k) \cos kx dk$	4.5
Classical frequency- time uncertainty	$\Delta f \Delta t \sim \varepsilon$	4.3	(continuous k) Group speed of wave	$v_{\text{group}} = \frac{d\omega}{dk}$	4.6
Heisenberg position- momentum uncertainty	$\Delta x \Delta p_x \sim \hbar$	4.4	packet	^r group dk	
Heisenberg energy-time uncertainty	$\Delta E \Delta t \sim \hbar$	4.4			

Time-independent Schrödinger equation	$-\frac{\hbar^2}{2m}\frac{d^2\psi}{dx^2} + U(x)\psi(x) = E\psi(x)$		Infinite potential energy well	$\psi_n(x) = \sqrt{\frac{2}{L}} \sin \frac{n\pi x}{L},$ $E_n = \frac{h^2 n^2}{8mL^2} (n = 1, 2, 3,)$	5.4
Time-dependent Schrödinger equation	$\Psi(x,t)=\psi(x)e^{-i\omega t}$	5.3	Two-dimensional	$\psi(x,y) = \frac{2}{L} \sin \frac{n_x \pi x}{L} \sin \frac{n_y \pi y}{L}$	5.4
Probability density	$P(x) = \psi(x) ^2$	5.3	infinite well	$E = \frac{h^2}{8mL^2}(n_x^2 + n_y^2)$	
Normalization condition	$\int_{-\infty}^{+\infty} \psi(x) ^2 dx = 1$	5.3		omL	
Probability in interval x_1 to x_2	$P(x_1:x_2) = \int_{x_1}^{x_2} \psi(x) ^2 dx$	5.3	Simple harmonic oscillator ground state	$\psi(x) = (m\omega_0/\hbar\pi)^{1/4} e^{-(\sqrt{km}/2\hbar)x^2}$	5.5
Average or expectation value of $f(x)$	$[f(x)]_{av} = \int_{-\infty}^{+\infty} \psi(x) ^2 f(x) dx$	5.3	Simple harmonic oscillator energies	$E_n = (n + \frac{1}{2})\hbar\omega_0 \ (n = 0, 1, 2, \ldots)$	5.5
Constant potential energy, $E > U_0$	$\psi(x) = A \sin kx + B \cos kx,$ $k = \sqrt{2m(E - U_0)/\hbar^2}$	5.4	Potential energy step, $E > U_0$	$\psi_0(x < 0) = A \sin k_0 x + B \cos k_0 x \psi_1(x > 0) = C \sin k_1 x + D \cos k_1 x$	5.6
Constant potential energy, $E < U_0$	$\psi(x) = Ae^{k'x} + Be^{-k'x},$ $k' = \sqrt{2m(U_0 - E)/\hbar^2}$	5.4	Potential energy step, $E < U_0$	$\psi_0(x < 0) = A \sin k_0 x + B \cos k_0 x$ $\psi_1(x > 0) = C e^{k_1 x} + D e^{-k_1 x}$	5.6

Scattering impact
$$b = \frac{zZ}{2K} \frac{e^2}{4\pi\varepsilon_0} \cot \frac{1}{2}\theta$$
 6.3
parameter

Fraction scattered $f_{>\theta} = nt\pi b^2$ at angles $> \theta$

 $N(\theta) =$

 $d = \frac{1}{4\pi\varepsilon_0} \frac{zZe^2}{K}$

Rutherford scattering formula

Distance of closest approach

Balmer formula

$$\lambda = (364.5 \text{ nm}) \frac{n^2}{n^2 - 4}$$

(n = 3, 4, 5, ...)
$$r_{-} = \frac{4\pi\varepsilon_0 \hbar^2}{n^2} n^2 = a_0 n^2$$

 $\frac{nt}{4r^2} \left(\frac{zZ}{2K}\right)^2 \left(\frac{e^2}{4\pi\varepsilon_0}\right)^2 \frac{1}{\sin^4\frac{1}{2}\theta}$

Radii of Bohr orbits in hydro

Radii of Bohr
orbits in hydrogen
$$r_{n} = \frac{4\pi c_{0}n}{me^{2}}n^{2} = a_{0}n^{2} \qquad 6.5$$

$$(n = 1, 2, 3, ...)$$
Energies of Bohr
orbits in hydrogen
$$E_{n} = -\frac{me^{4}}{32\pi^{2}\varepsilon_{0}^{2}\hbar^{2}}\frac{1}{n^{2}} \qquad 6.5$$

$$= \frac{-13.60 \text{ eV}}{n^{2}} (n = 1, 2, 3, ...)$$

Prbital angular
$$|\vec{\mathbf{L}}| = \sqrt{l(l+1)}\hbar$$
 7.2

Orbital angular
$$|\mathbf{L}| = \sqrt{l(l+1)\hbar}$$
 7.2
momentum $(l = 0, 1, 2, ...)$

 $\Delta L_{\tau} \Delta \phi \ge \hbar$

Orbital magnetic
$$L_z = m_l \hbar$$

quantum number $(m_l = 0, \pm 1, \pm 2, \dots, \pm l)$

 $\cos\theta = \frac{L_z}{|\mathbf{L}|} = \frac{m_l}{\sqrt{l(l+1)}}$

Spatial quantization

Angular momentum uncertainty relationship

Hydrogen $n = 1, 2, 3, \ldots$ quantum numbers $l = 0, 1, 2, \ldots, n-1$ $m_l = 0, \pm 1, \pm 2, \ldots, \pm l$

Hydrogen energy
$$E_n = -\frac{me^4}{32\pi^2\varepsilon_0^2\hbar^2}\frac{1}{n^2}$$

levels

Hydrogen wave
functions
$$\psi_{n,l,m_l}(r,\theta,\phi) =$$

 $R_{n,l}(r)\Theta_{l,m_l}(\theta)\Phi_{m_l}(\phi)$ Radial probability $P(r) = r^2 |R_{n,l}(r)|^2$

density
$$P(r) = r^2 | R_j$$

Pauli exclusion
principleNo two electrons in a single
atom can have the same
set of quantum numbers
$$(n, l, m_l, m_s)$$
.8.1Filling order of
atomic subshells $1s, 2s, 2p, 3s, 3d, 4s, 3d,$
 $4p, 5s, 4d, 5p, 6s, 4f, 5d,$
 $6p, 7s, 5f, 6d$ 8.2 Capacity of
subshell nl $2(2l + 1)$ 8.2

Excitation energy $E_n - E_1$ of level n

 $|E_n|$

Binding (or ionization) energy of level n

6.3

6.3

6.3

6.4

Hydrogen wavelengths in

Bohr model

Single-electron atoms with Z > 1Reduced mass of proton-electron system

Angular probabi-

Orbital magnetic

dipole moment

Spin magnetic

dipole moment

Spin angular

Spin magnetic quantum number

Spectroscopic

Selection rules for

photon emission

Normal Zeeman

Fine-structure

Adding

notation

effect

estimate

momentum

lity density

7.2

7.2

7.2

7.3

7.3

7.3

7.4

$$\lambda = \frac{64\pi^3 \varepsilon_0^2 \hbar^3 c}{m e^4} \left(\frac{n_1^2 n_2^2}{n_1^2 - n_2^2} \right)$$

$$= \frac{1}{R_\infty} \left(\frac{n_1^2 n_2^2}{n_1^2 - n_2^2} \right)$$

$$r_n = \frac{a_0 n^2}{Z}, E_n = -(13.60 \text{ eV}) \frac{Z^2}{n^2}$$

$$m = \frac{m_e m_p}{m_e + m_p}$$

$$6.8$$

7.5 $P(\theta, \phi) = |\Theta_{l,m_l}(\theta) \Phi_{m_l}(\phi)|^2$

$$\vec{\mu}_{\rm L} = -(e/2m)\vec{\rm L}$$
 7.6

$$\vec{\mu}_{\rm S} = -(e/m)\vec{\rm S} \qquad 7.6$$

$$|\vec{\mathbf{S}}| = \sqrt{s(s+1)}\hbar = \sqrt{3/4}\hbar$$
 7.6
(for $s = 1/2$)

$$S_z = m_s \hbar (m_s = \pm 1/2)$$
 7.6

$$s (l = 0), p(l = 1), d(l = 2),$$
 7.7
 $f(l = 3), ...$

$$\Delta l = \pm 1$$
 $\Delta m_l = 0, \pm 1$ 7.7, 7.8

$$\Delta \lambda = \frac{\lambda^2}{hc} \Delta E = \frac{\lambda^2}{hc} \mu_{\rm B} B$$
 7.8

$$\Delta E = mc^2 \alpha^4 / n^5 \quad (\alpha \approx 1/137) \qquad 7.9$$

Energy of
$$E_n = (-13.6 \text{ eV}) \frac{Z_{\text{eff}}^2}{n^2}$$
 8.3
Screened electron $\Delta E = (10.2 \text{ eV})(Z-1)^2$ 8.5
for K_{α} X rays
Adding angular $L_{\text{max}} = l_1 + l_2$, 8.6
momenta l_1, m_{l_1} $L_{\text{min}} = |l_1 - l_2|$,

and
$$l_2, m_{l2}$$

Hund's rules for First $S = M_{S,max}$, then 8.6
ground state $L = M_{L,max}$

6.5

6.5

Nuclear radius Nuclear binding energy	$R = R_0 A^{1/3}, R_0 = 1.2 \text{ fm}$ $B = [Nm_n + Zm(^1_1 H_0) - m(^A_Z X_N)]c^2$	12.2 12.3	Q value of decay $X \rightarrow X' + x$	$Q = [m_{\mathrm{X}} - (m_{\mathrm{X}'} + m_{\mathrm{x}})]c^2$	12.6
Proton separation	$S_{\rm p} = [m(^{A-1}_{Z-1}{\rm X}'_N) + m(^{1}{\rm H}) - m(^{A}_{Z}{\rm X}_N)]c^2$	12.3	Q value of alpha decay	$Q = [m(\mathbf{X}) - m(\mathbf{X}') - m(^{4}\mathrm{He})]c^{2}$	12.7
energy Neutron	$S_{n} = [m({}^{A-1}_{Z}X_{N-1}) + m_{n} - m({}^{A}_{Z}X_{N})]c^{2}$	12.3	of alpha	$K_{\alpha} \cong Q(A-4)/A$	12.7
separation energy			particle Q values of	$Q_{\beta^-} = [m(^A \mathbf{X}) - m(^A \mathbf{X}')]c^2,$	12.8
Range of	$mc^2 = \hbar c/x$	12.4	beta decay	$Q_{\beta^+} = [m(^{\Lambda}X) - m(^{\Lambda}X') - 2m_{\rm e}]c^2$	
exchanged particle			Recoil in gamma decay	$K_{\rm R} = E_{\gamma}^2 / 2Mc^2$	12.9
Activity	$a = \lambda N, \lambda = \ln 2/t_{1/2} = 0.693/t_{1/2}$	12.5			
Radioactive decay law	$N = N_0 e^{-\lambda t}, a = a_0 e^{-\lambda t}$	12.5			

Forces	Strong, electromagnetic, weak, gravitational	14.1	Conservation of baryon number <i>B</i>	In any process, B remains constant.	14.3
Field particles	Gluon (g), photon (γ), weak boson (W [±] , Z ⁰), graviton	14.1	Conservation of strangeness S	In strong and electromagnetic processes, S remains constant; in	14.3
Leptons	$e^{-}, v_{e}, \mu^{-}, v_{\mu}, \tau^{-}, v_{\tau}$	14.2		weak processes, $\Delta S = 0$ or ± 1 .	
Mesons	$\pi^{\pm}, \pi^{0}, \mathbf{K}^{\pm}, \mathbf{K}^{0}, \overline{\mathbf{K}}^{0}, \eta, \rho^{\pm}, \eta',$ $\mathbf{D}^{\pm}, \psi, \mathbf{B}^{\pm}, \Upsilon, \dots$	14.2	Q value in decays or reactions	$Q = (m_{\rm i} - m_{\rm f})c^2$	14.5, 14.6
Baryons	p, n, Λ^0 , $\Sigma^{\pm,0}$, $\Xi^{-,0}$, Ω^- ,	14.2		$K_{\rm th} = -Q(m_1 + m_2 + m_3 + m_4 + m_5 + \cdots)/2m_2$	14.6
Conservation of lepton number L	In any process, L_e, L_μ , and L_τ remain constant.	14.3	Quarks	u, d, c, s, t, b	14.7

SOME PARTICLE MASSES

	kg	u	MeV/c ²		Ζ	A	Atomic mass (u)	Abundance or Half-life
Electron	$9.1093829 \times 10^{-31}$	$5.485799095 \times 10^{-4}$	0.51099893	Н	1	1	1.0078250	99.985%
Proton	$1.67262178 \times 10^{-27}$	1.0072764668	938.27205			2	2.014102	0.015%
FIOIDII			936.27203			3	3.016049	12.3 y
Neutron	$1.67492735 \times 10^{-27}$	1.0086649160	939.56538	He	2	3	3.016029	0.000137%
Deuteron	$3.3435835 \times 10^{-27}$	2.0135532127	1875.61286	110	2	4	4.002603	99.999863%
Alpha	$6.6446568 \times 10^{-27}$	4.001506179	3727.3792	Li	3	6	6.015123	7.59%
						7	7.016005	92.41%
						8	8.022487	0.84 s
				Be	4	7	7.016930	53.2 d
						8	8.005305	0.07 fs
CONVE	RSION FACTORS					9	9.012182	100%
GOINVE	NOION FACIONO					10	10.013534	1.5 My
	10 1	11 (1) (0) 28 2				11	11.021658	13.8 s
$1 \mathrm{eV} = 1.60$	$0217657 \times 10^{-19} \mathrm{J}$	1 barn (b) = 10^{-28} m^2		в	5	8	8.024607	0.77 s
$1 u = 931.49406 MeV/c^2$		$1 \text{ curie (Ci)} = 3.7 \times 10^{10}$	⁰ decays/s			9	9.013329	0.85 as
- 16	$6053892 \times 10^{-27} \mathrm{kg}$	1 light-year = 9.46×10^{10}	5 m			10	10.012937	19.8%
						11	11.009305	80.2%
1 y = 3.15	$56 \times 10^7 \mathrm{s} \cong \pi \times 10^7 \mathrm{s}$	1 parsec = 3.26 light	year			12	12.014352	20.2 ms