

# Numerical Experiments on the Neutron Star Magnetosphere

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*In collaboration with:*

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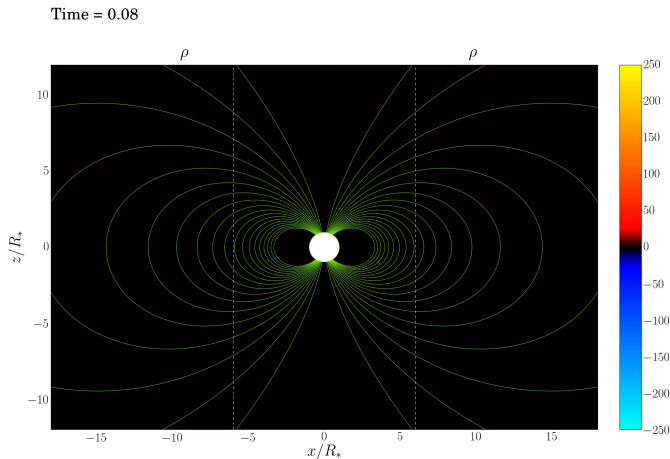
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## Aperture is Particles, Electrodynamics, Radiative Transfer at Ultra-Relativistic Energies

- A general PIC framework we designed and developed from scratch, in order to study the neutron star magnetospheres.
- Models radiative cooling, curvature radiation and its conversion into electron-positron pairs
- Can run on workstation GPUs or on CPU clusters. Tested on the Yeti cluster at Columbia and shows good scaling.

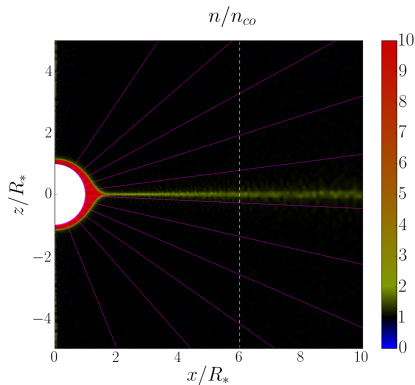
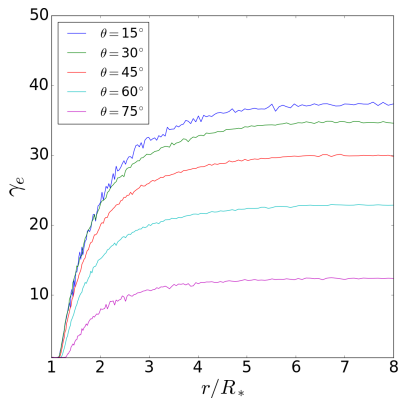
# NE1: Pulsar Simulations

One of the first self-consistent kinetic pulsar magnetosphere simulations (Chen & Beloborodov 2014).



## NE2: Monopole Solution

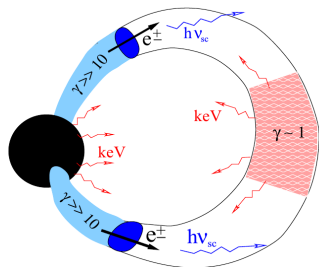
Mainly work done by Rui Hu



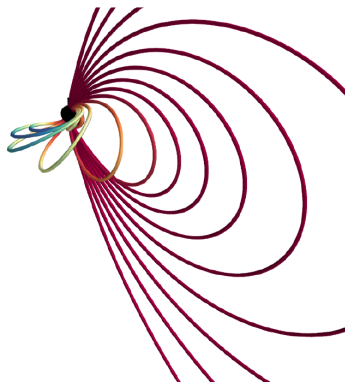
Michel monopole solution with a finite atmosphere

# NE3: Corona of Magnetars

Simulating the twist and untwist of current bundles along dipolar field lines



(Beloborodov 2013)



(Parfrey et al. 2013)

# Twisting the Magnetar Magnetosphere

The surface of the star rotates differentially, launching a twist on the magnetic field lines, traveling as Alfvén waves along the flux tube.

Equatorial twist profile:

$$\omega(\theta) = \omega_0 \frac{\Theta}{\sin \theta} \exp \left[ (1 - \Theta^4)/4 \right]$$

where  $\Theta = (\theta - \pi/2)/\Delta\theta_m$ . (Mikic & Linker 1994)

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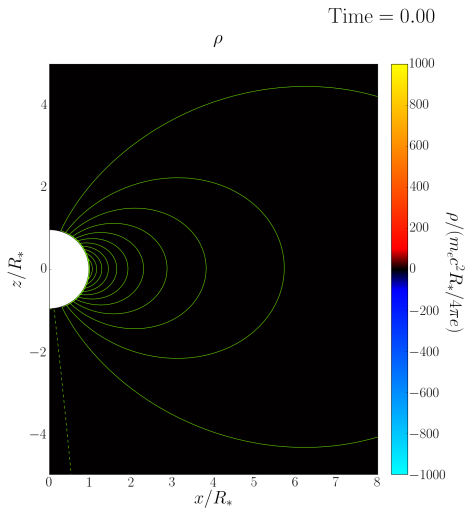
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Twist is implanted within a short time, but the resulting  $B_\phi$  survives on a much longer time scale.

# Twisting the Magnetar Magnetosphere





# Mechanism of Untwisting

The current bundle untwists through Ohmic dissipation. Voltage drop is limited by the pair creation threshold voltage. In this simulation

$$e\Phi_e/mc^2 \approx 30, \quad e\Phi_{max}/mc^2 \approx 200$$

# Mechanism of Untwisting

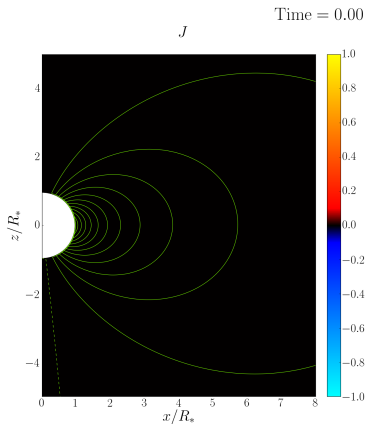
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$$\frac{d\psi}{dt} = 2\pi c \frac{\partial \Phi_e}{\partial f}$$

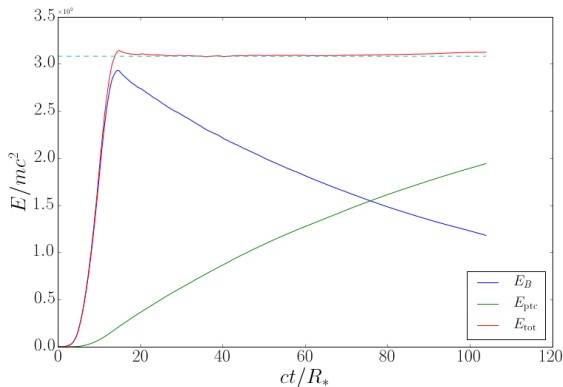
# Mechanism of Untwisting

The magnetosphere untwists with a growing cavity of zero current (Beloborodov 2009).



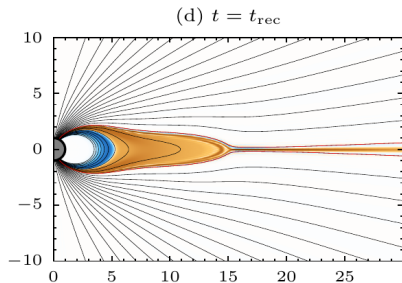
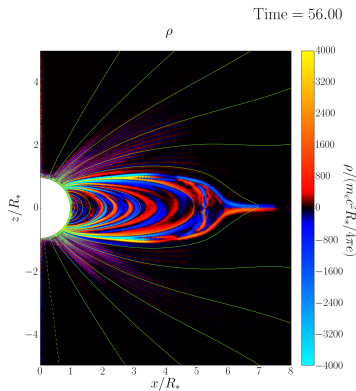
# The Energy Dump

Magnetic energy is slowly converted to particle kinetic energy, and eventually advected into the star.



# Over-twisted Magnetosphere

When the twist angle becomes larger than some critical angle  $\psi_{crit}$ , the behavior of the  $j$ -bundle changes qualitatively. Untwist happens violently via reconnection.



(Parfrey et al. 2013)

Would be interesting to study the reconnection!