

Black Hole Magnetospheres

Ioannis Contopoulos, RCAAM, Academy of Athens

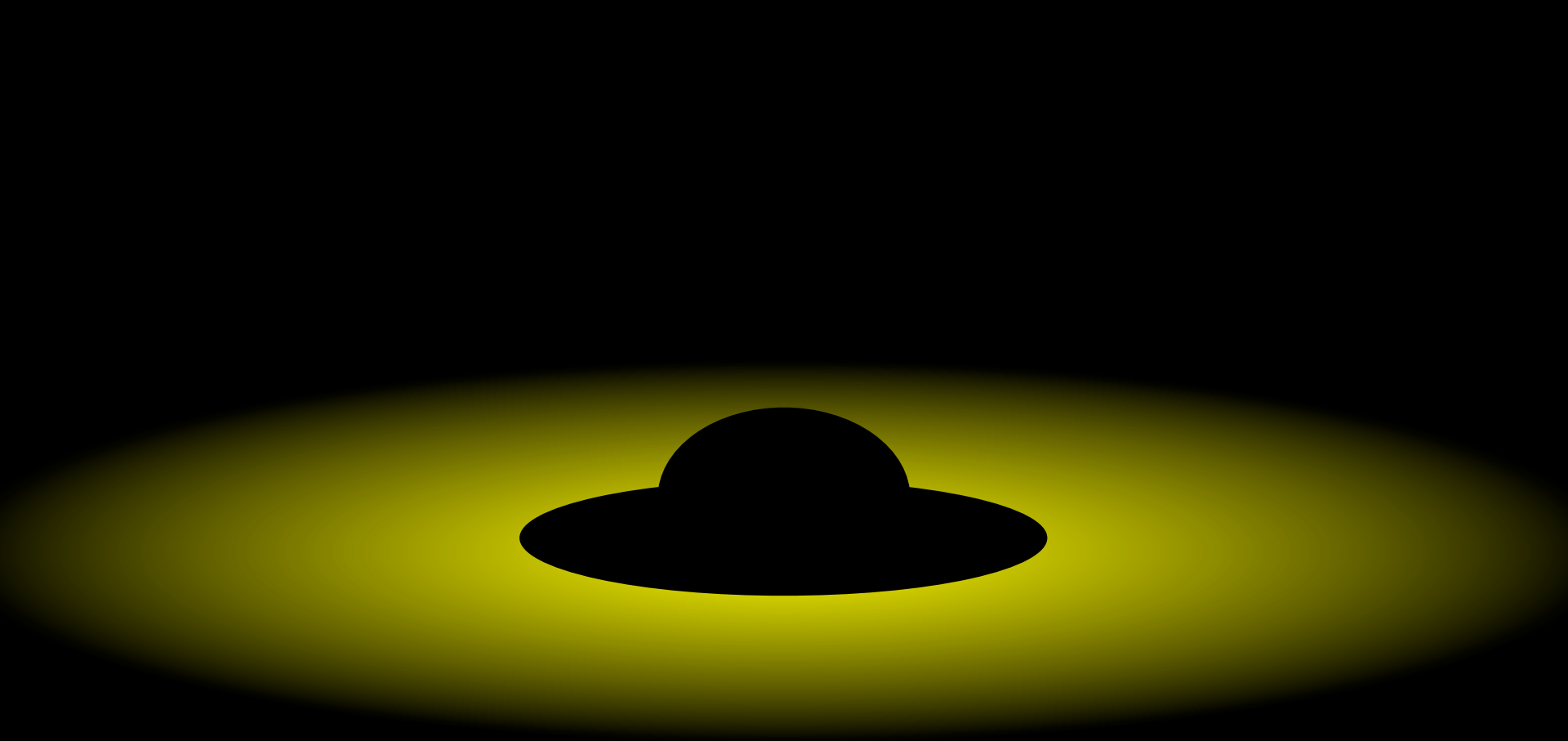


European Union
European Social Fund



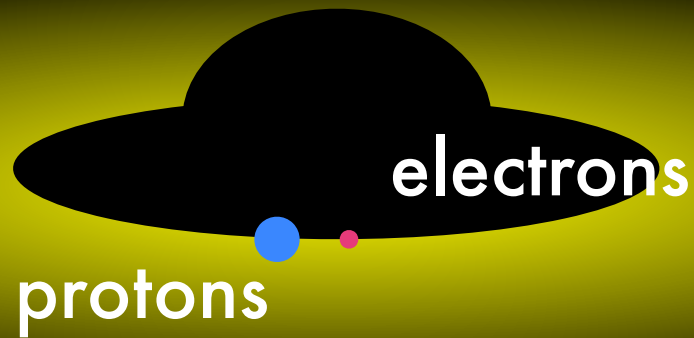
Co-financed by Greece and the European Union





The Cosmic Battery

Contopoulos & Kazanas 1998



The Cosmic Battery

Contopoulos & Kazanas 1998



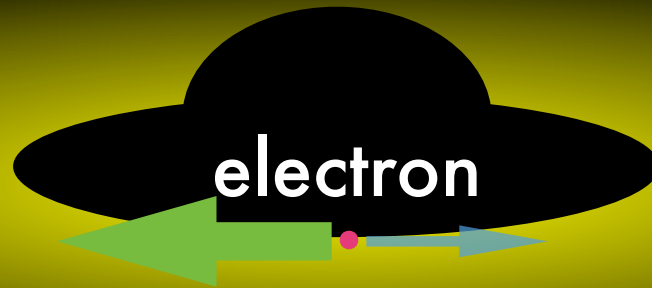
The Cosmic Battery

Koutsantoniou & Contopoulos 2014 (in prep.)



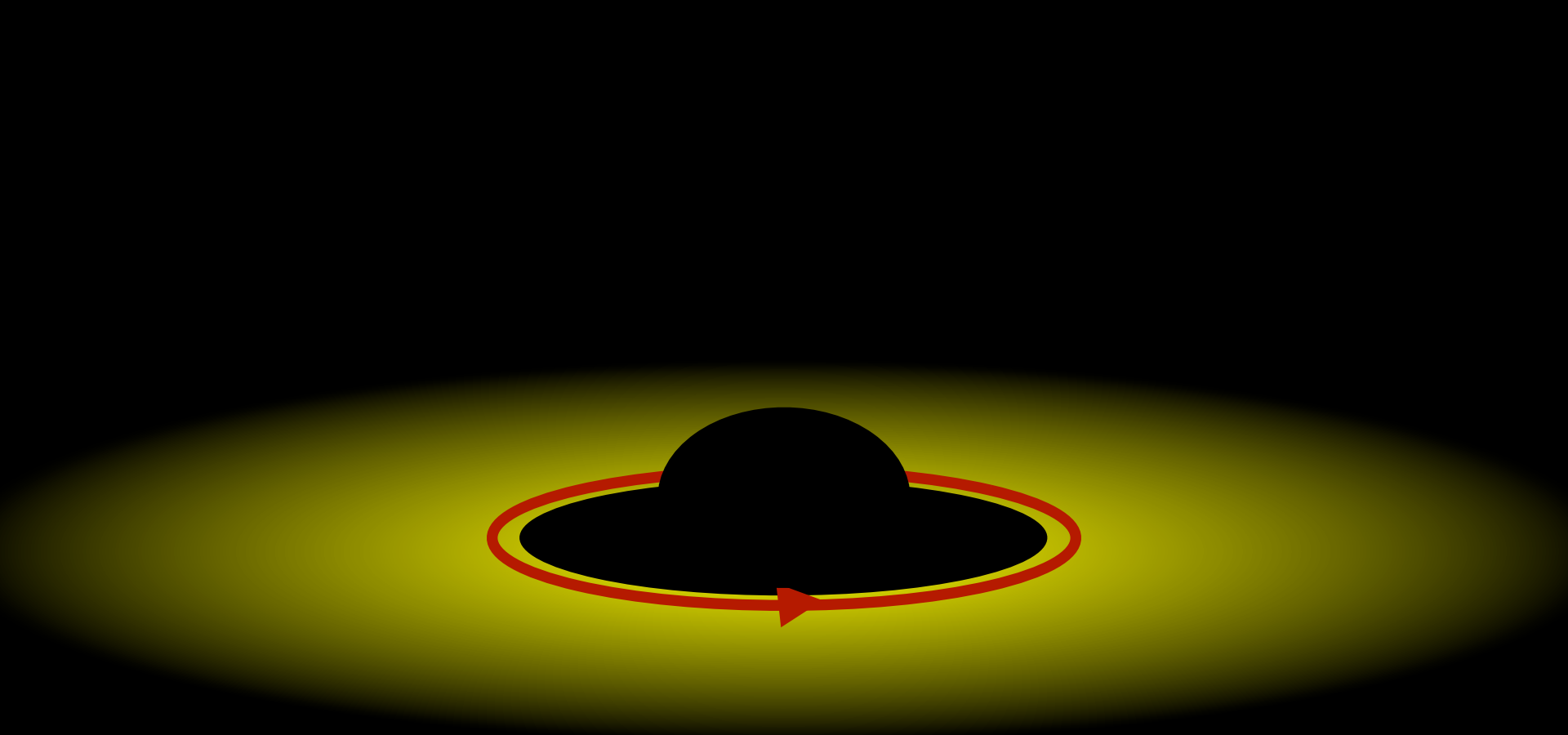
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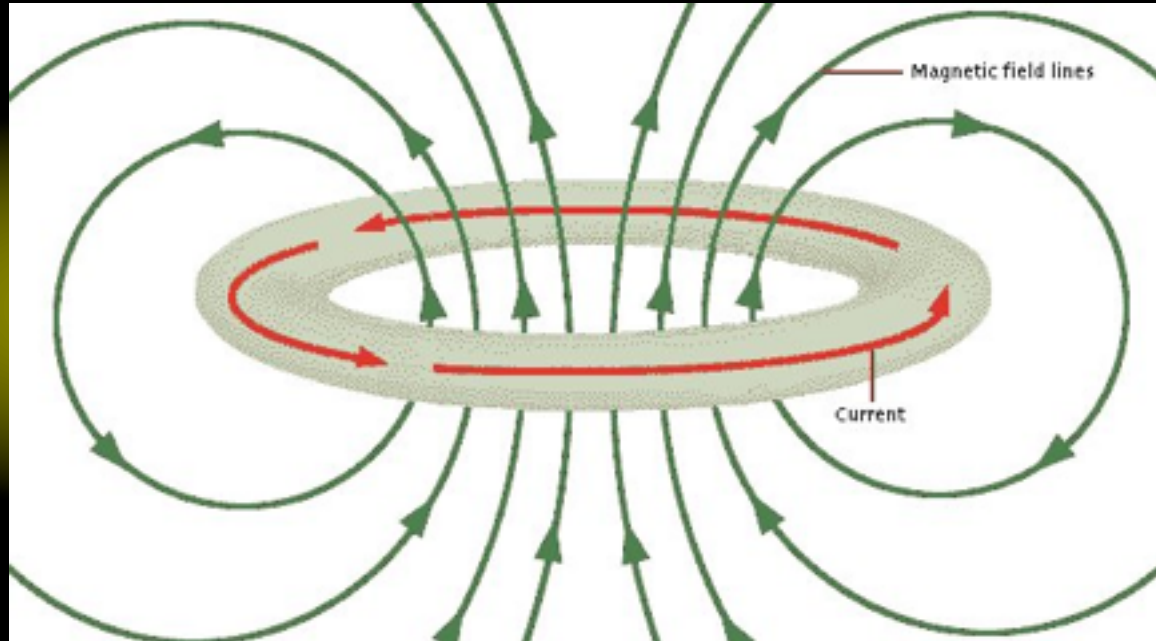
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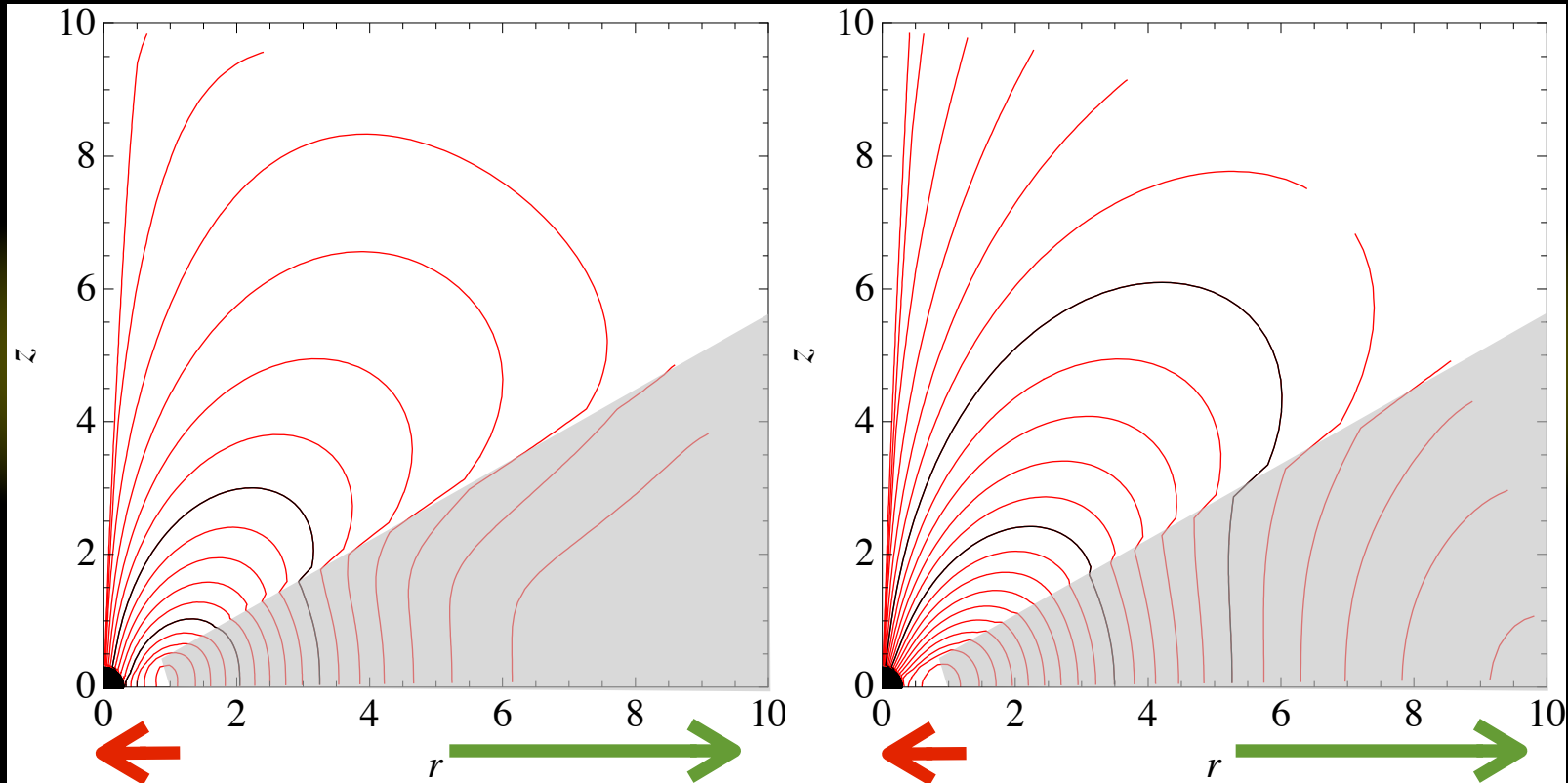
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The Cosmic Battery

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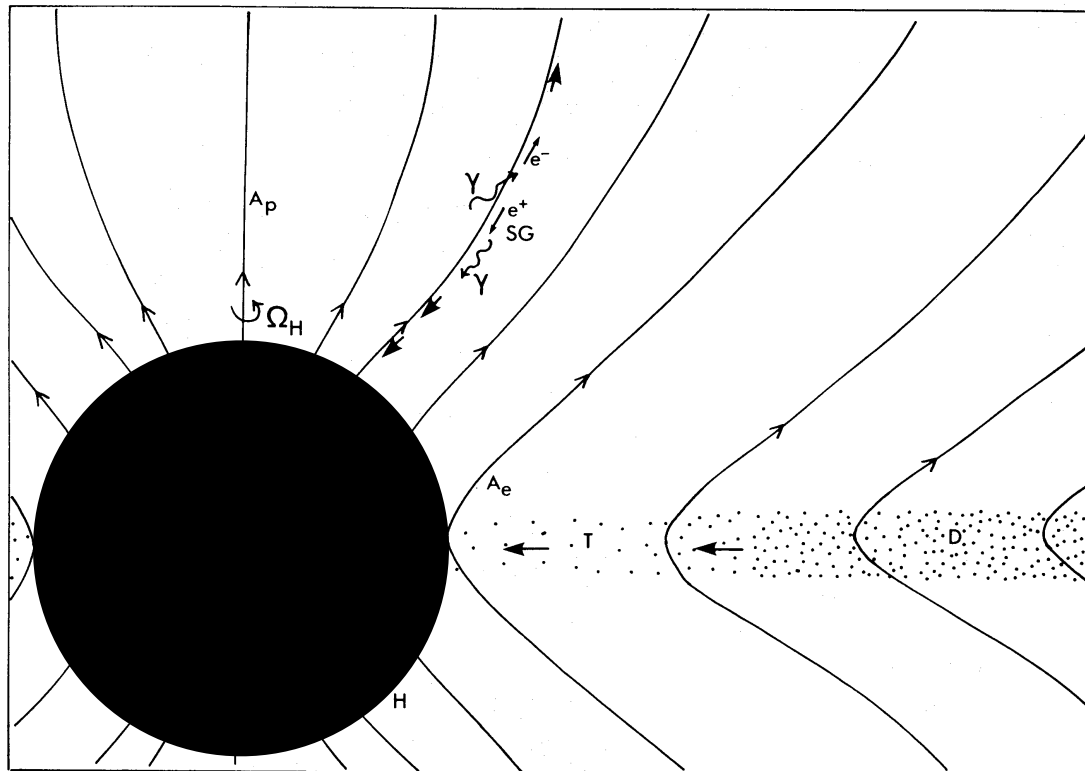
$$\frac{\partial \Psi_m}{\partial t} = 2\pi\alpha r \underbrace{(v \times B)}_{\leftarrow} - \frac{f_{\text{rad}}^\phi}{e} + \eta \underbrace{\nabla \times B}_{\rightarrow}$$



The Cosmic Battery

Katsanikas & Contopoulos 2014 (in prep.)

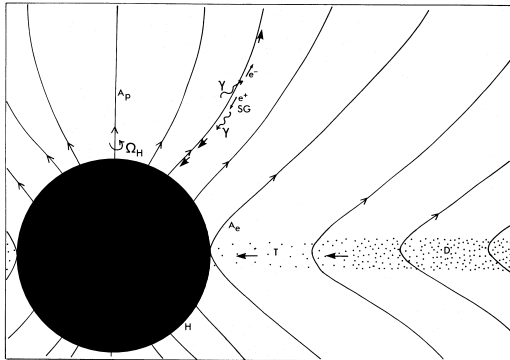
$$\dot{E}_{\text{BZ}} \propto \omega(\Omega_{\text{BH}} - \omega)\Psi_m^2 \sim \Omega_{\text{BH}}^2 \Psi_m^2$$



Blandford
Znajek 1977

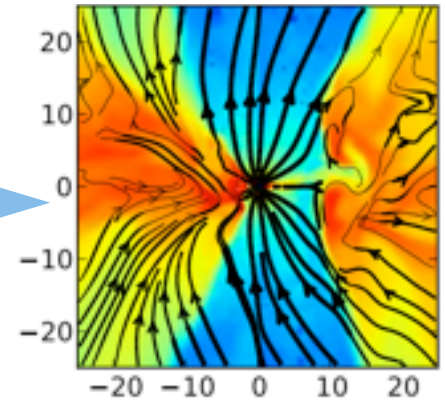
Blandford
Znajek 1977

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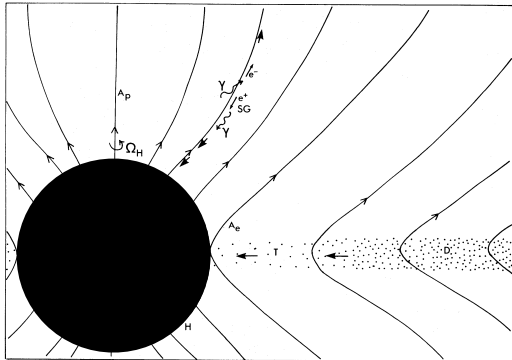
Blandford
Znajek 1977

+ 33 years



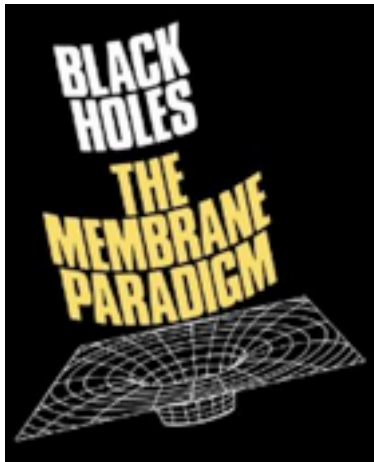
Narayan
Tchekhovskoy
McKinney 2010...

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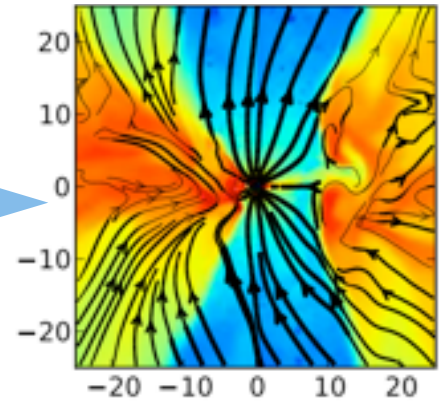


Blandford
Znajek 1977

Price
Thorne 1988

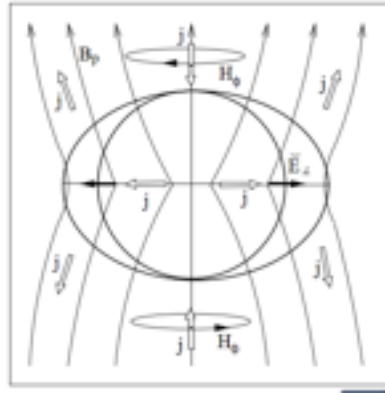
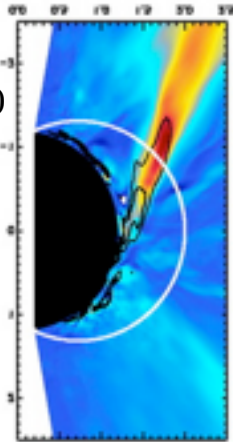


+ 33 years



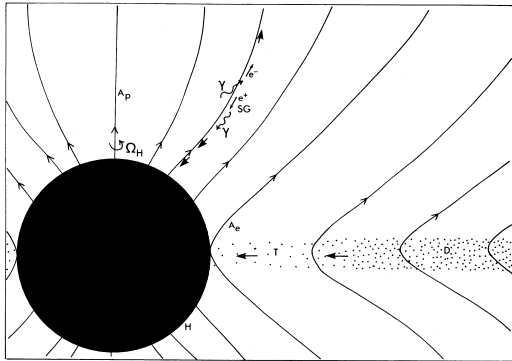
Narayan
Tchekhovskoy
McKinney 2010...

Punsly
Coroniti 1990

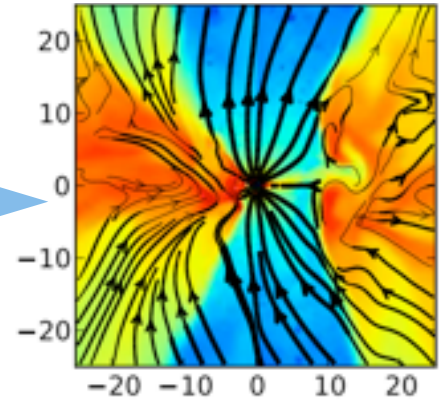


Komissarov 2004

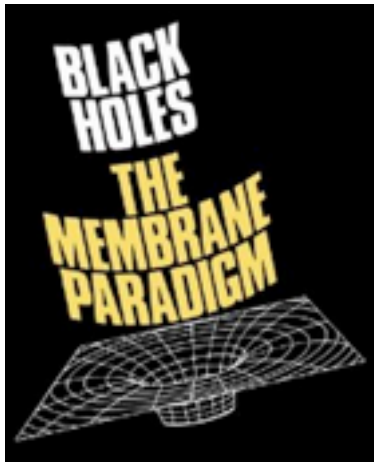
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+ 33 years



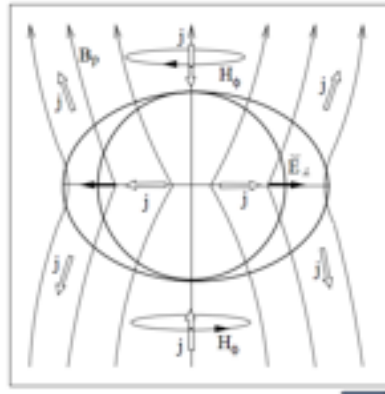
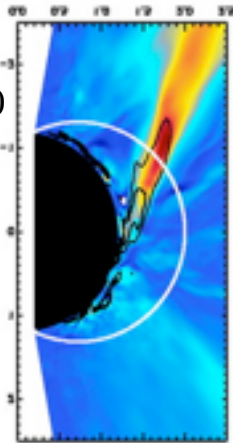
Blandford
Znajek 1977



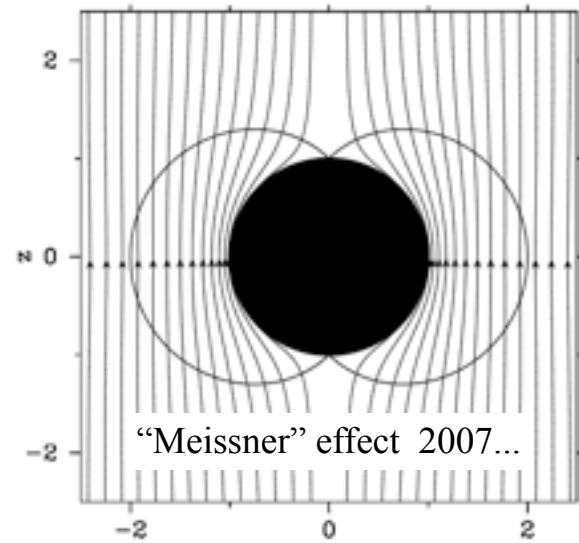
Price
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Narayan
Tchekhovskoy
McKinney 2010...

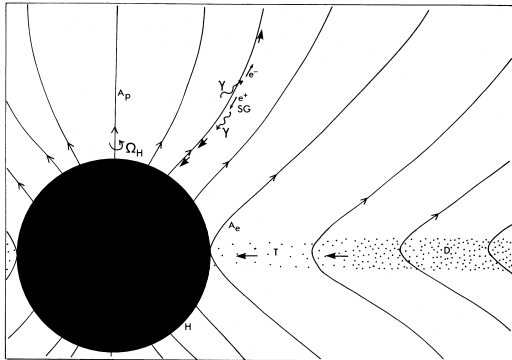
Punsly
Coroniti 1990



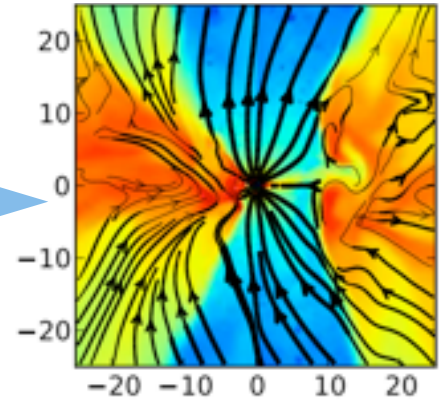
Komissarov 2004



$$\dot{E}_{\text{BZ}} \propto \omega(\Omega_{\text{BH}} - \omega)\Psi_m^2 \sim \Omega_{\text{BH}}^2 \Psi_m^2$$

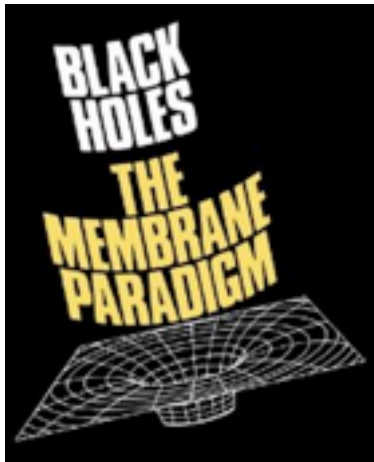


+ 33 years

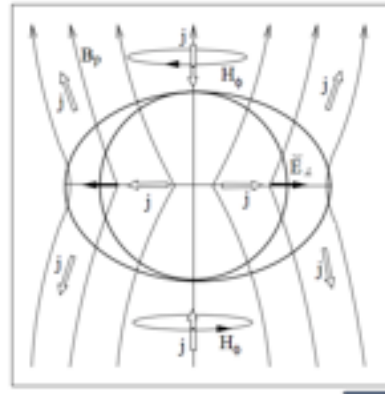


Narayan
Tchekhovskoy
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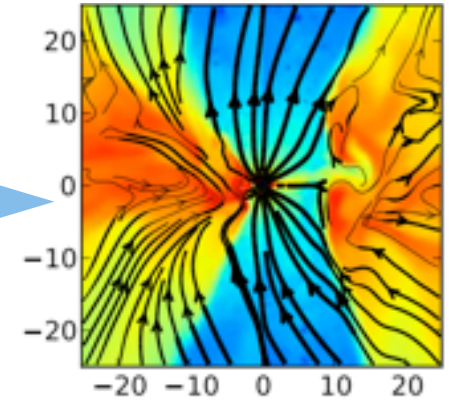

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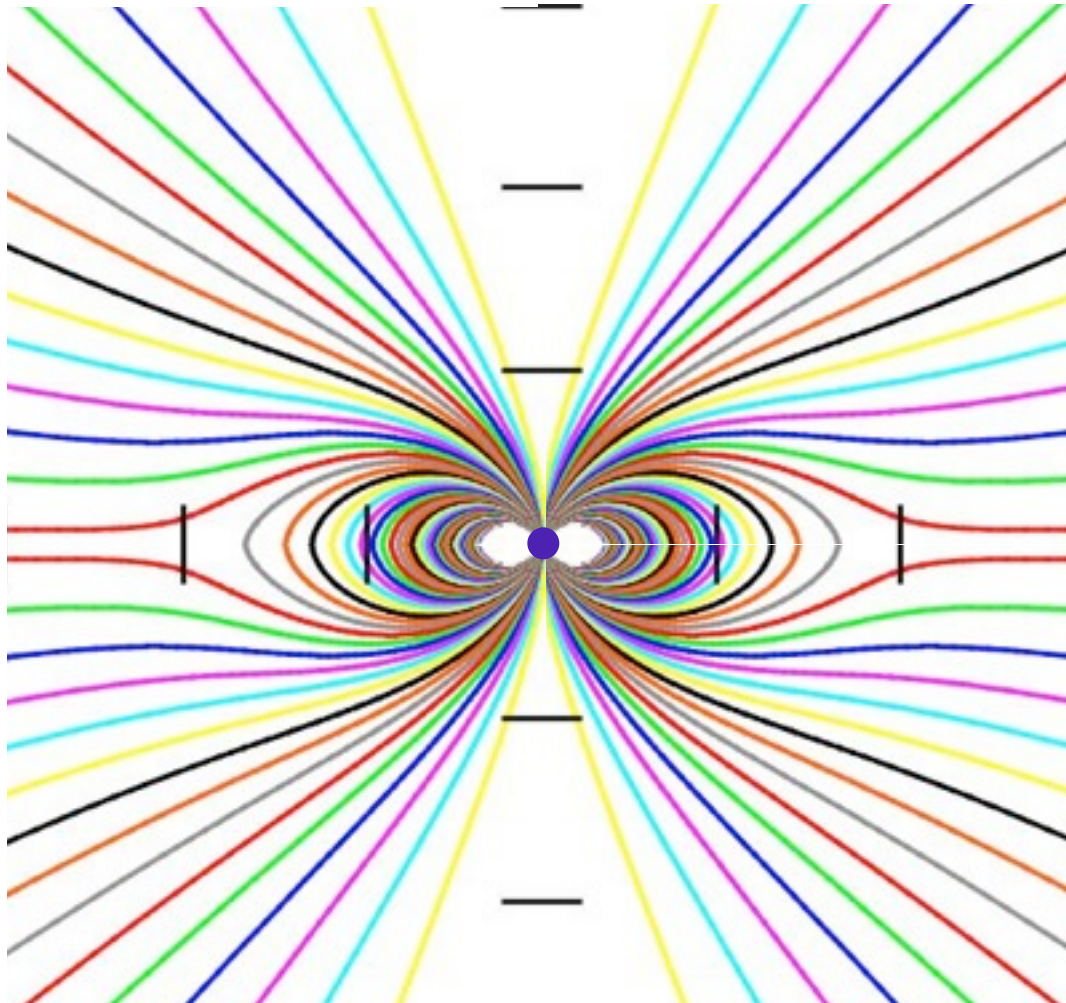


“Meissner” effect 2007...

[illegible]

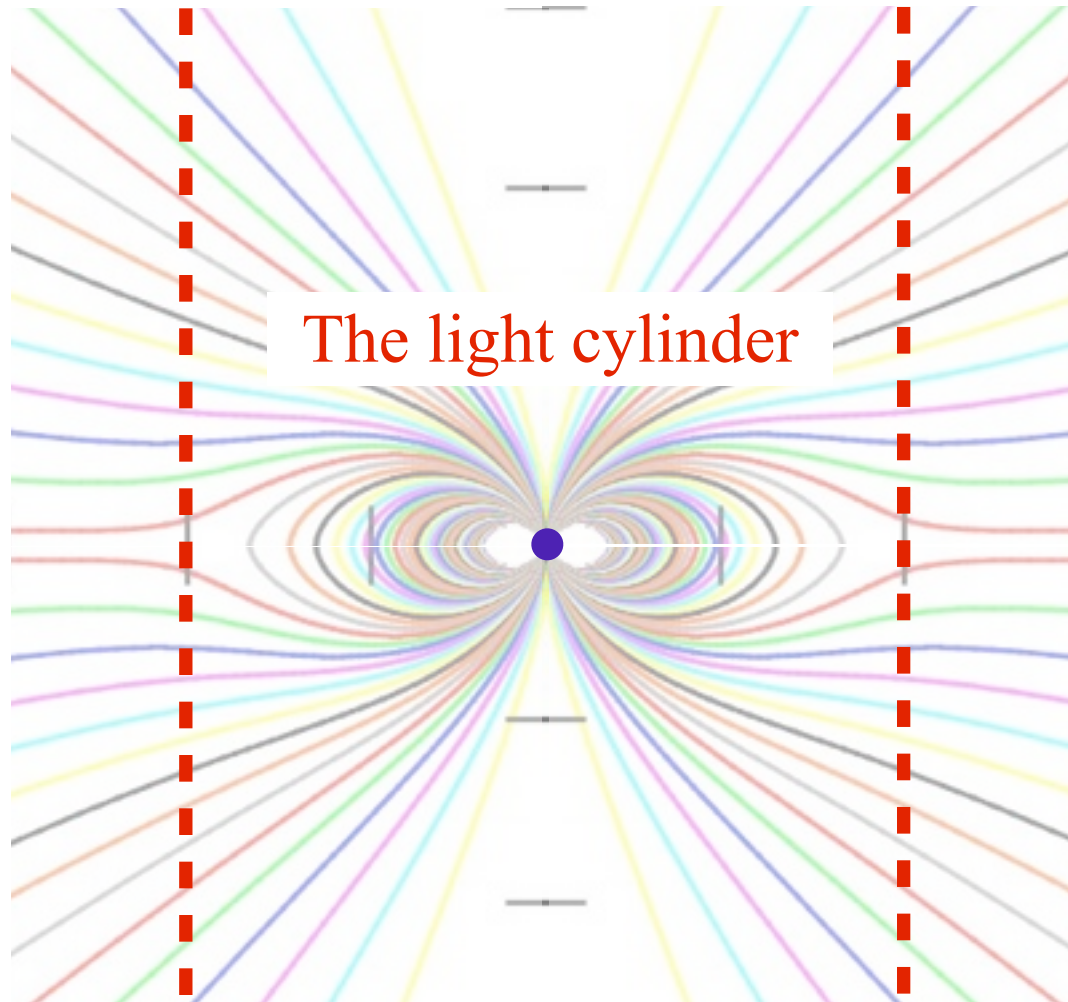
Narayan & McClintock 2012
Russell, Gallo & Fender 2013

The analogy with pulsars



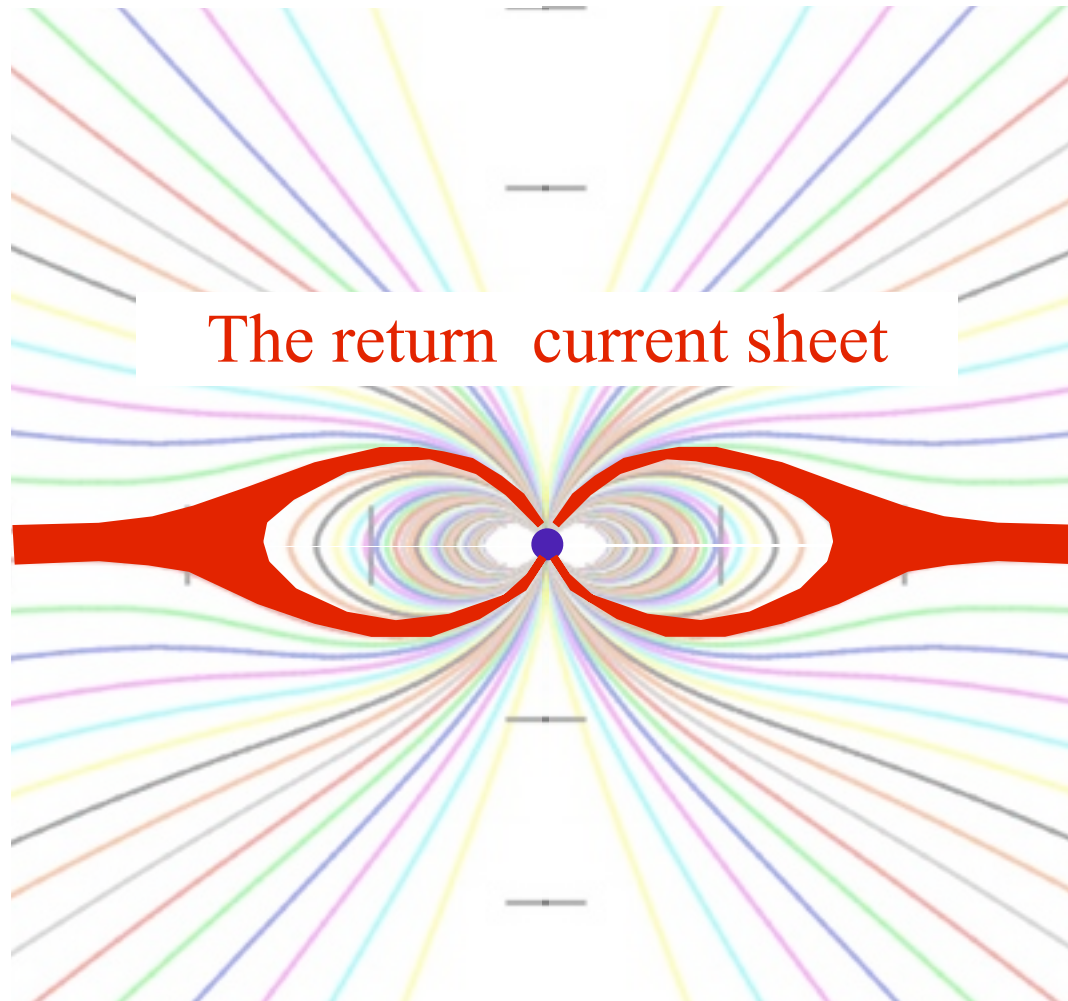
Contopoulos, Kazanas & Fendt 1999

The analogy with pulsars



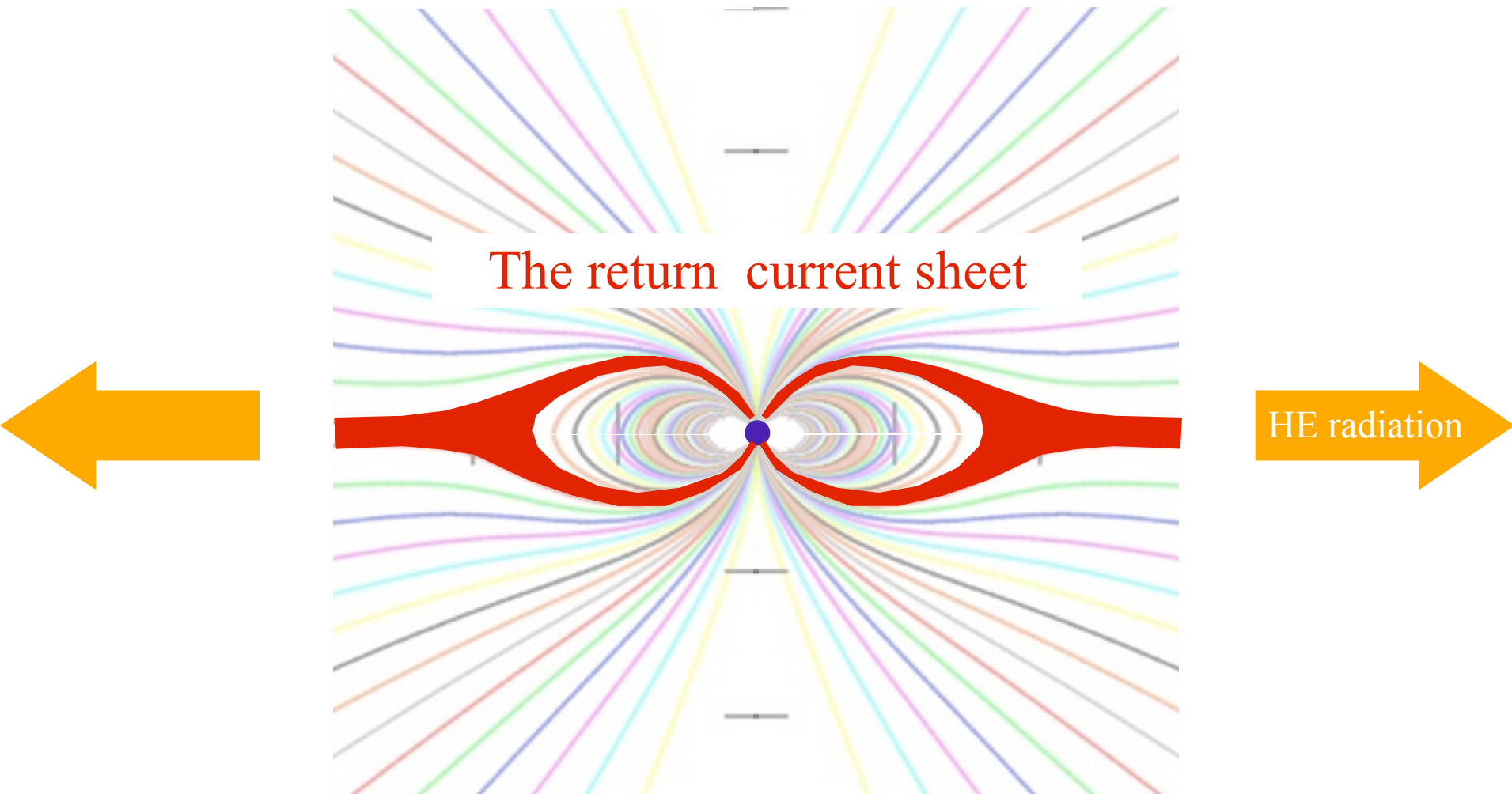
Contopoulos, Kazanas & Fendt 1999

The analogy with pulsars

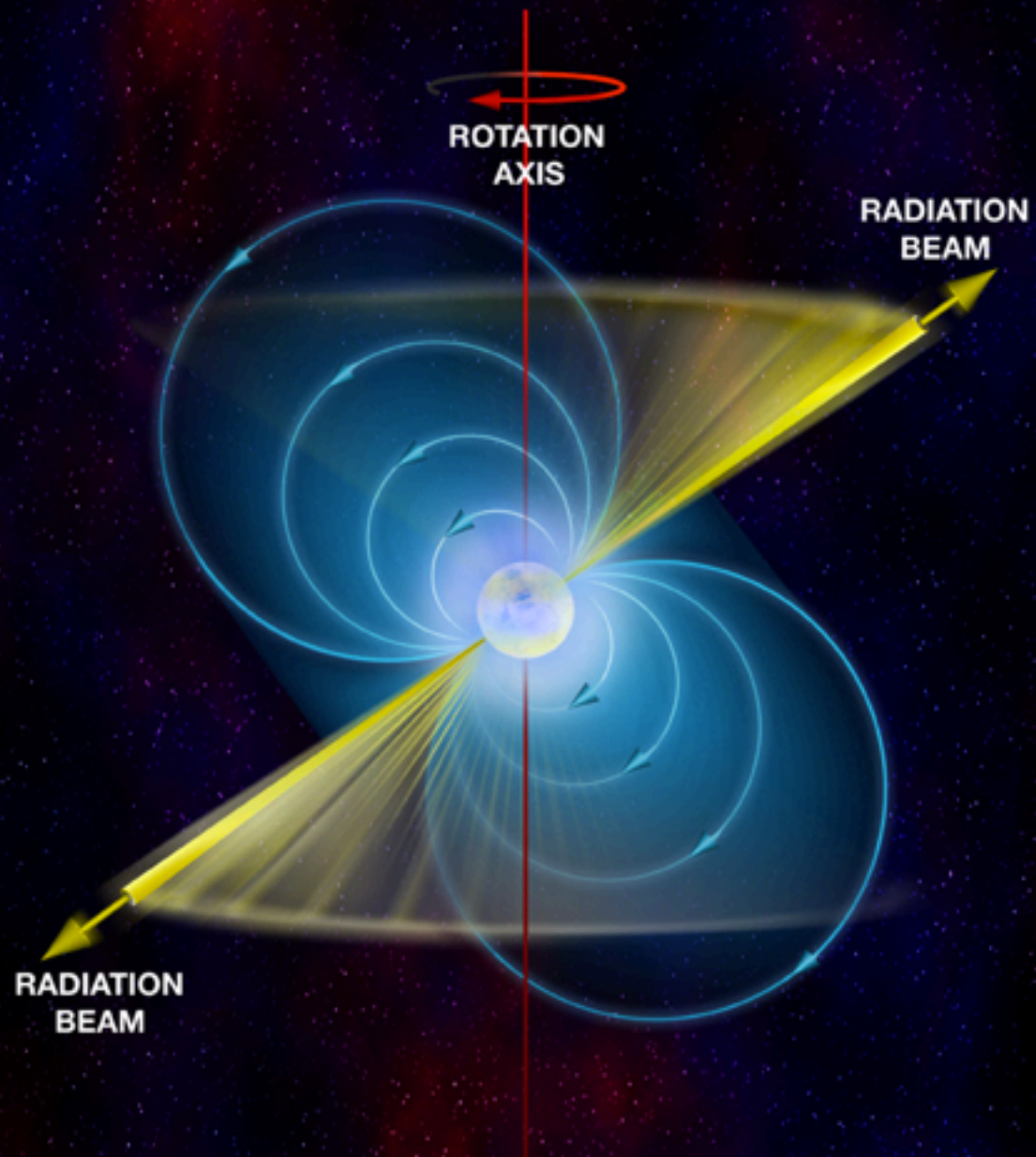


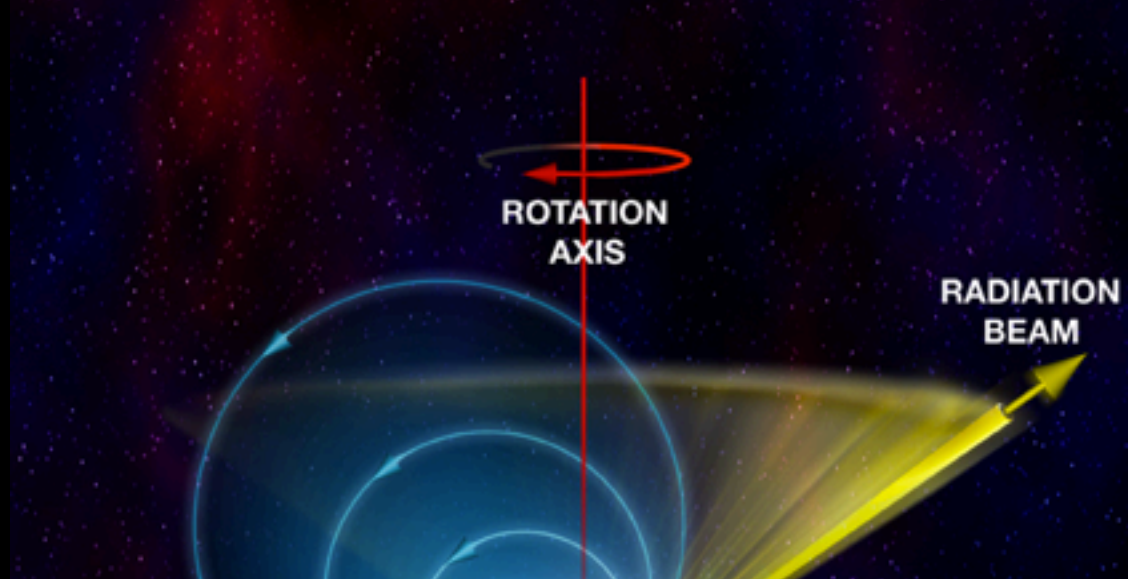
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The analogy with pulsars

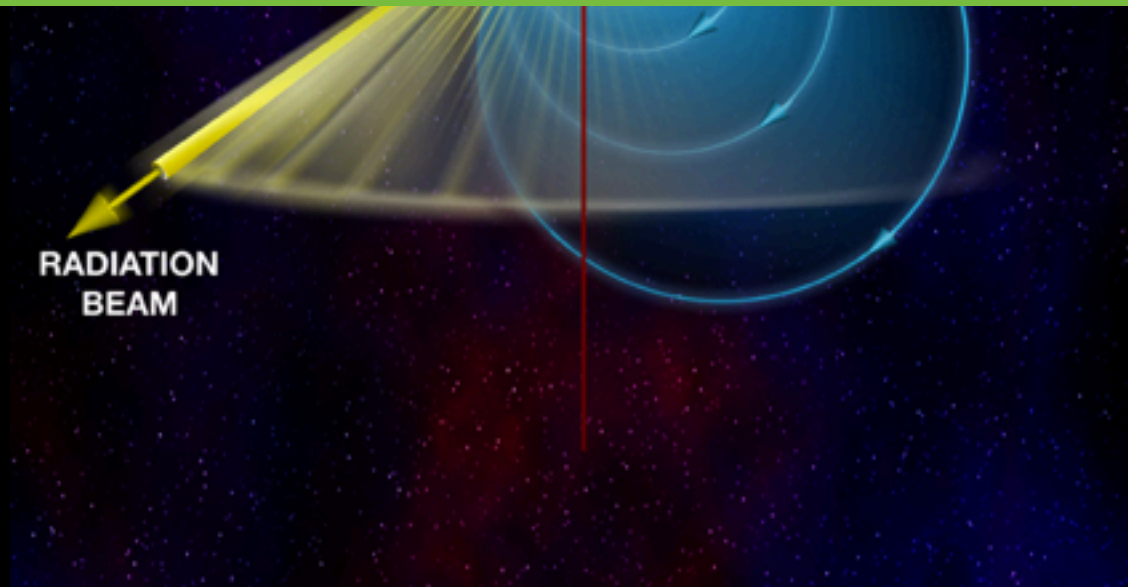


Contopoulos, Kazanas & Fendt 1999



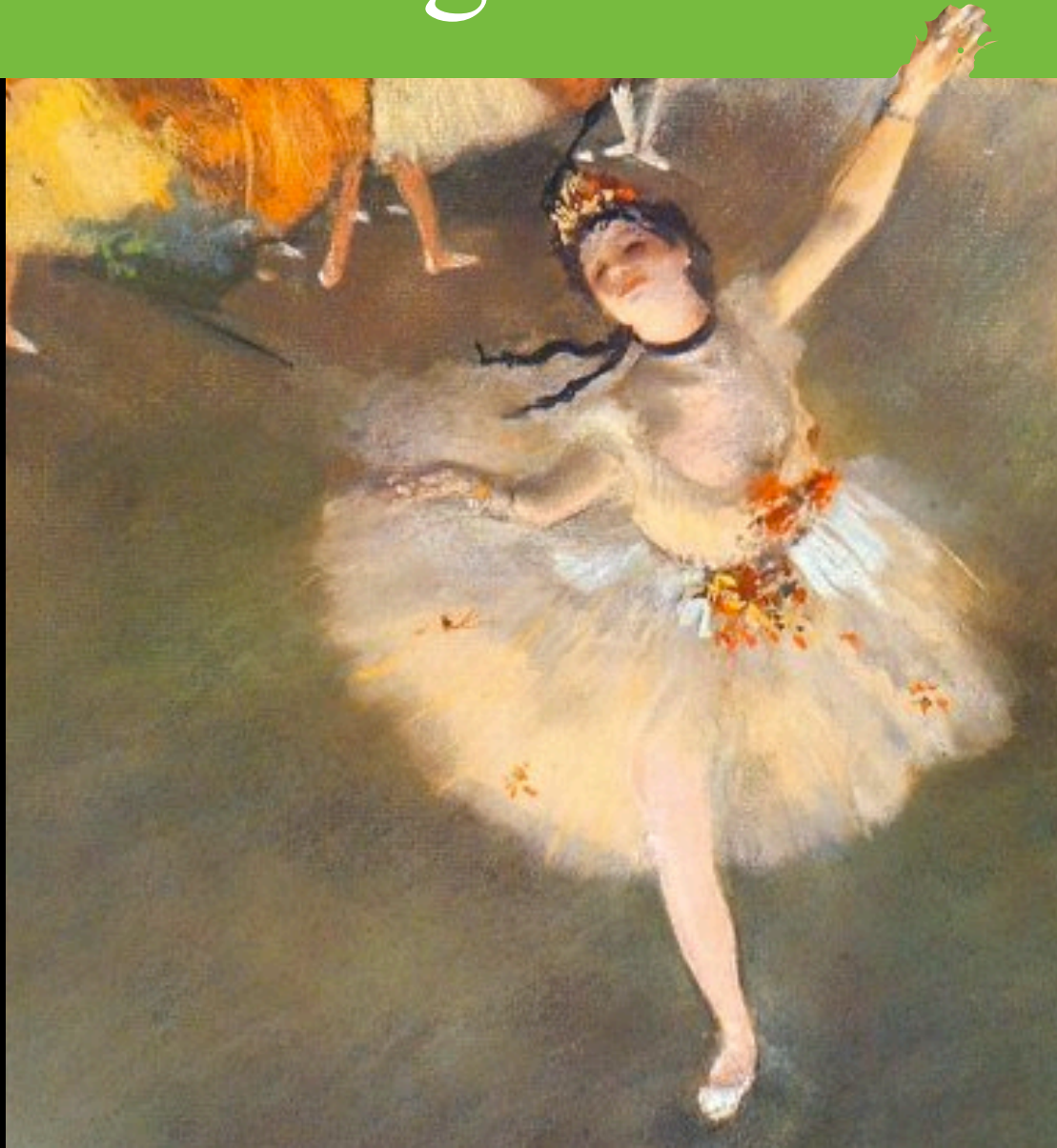


Not a lighthouse!





An “orthogonal” emitter



The GR pulsar equation

$$ds^2 = -\alpha^2 dt^2 + \varpi^2 (d\phi - \Omega dt)^2 + \frac{\Sigma}{\Delta} dr^2 + \Sigma d\theta^2$$

$$\nabla \cdot \mathbf{B} = 0$$

$$\nabla \cdot \mathbf{E} = 4\pi\rho_e$$

$$\nabla \times (\alpha\mathbf{B}) = 4\pi\alpha\mathbf{J}$$

$$\nabla \times (\alpha\mathbf{E}) = 0$$

$$\mathbf{B}(r, \theta) = \frac{1}{\sqrt{A} \sin \theta} \left\{ \Psi_{,\theta} \mathbf{e}_{\hat{r}} - \sqrt{\Delta} \Psi_{,r} \mathbf{e}_{\hat{\theta}} + \frac{2I\sqrt{\Sigma}}{\alpha} \mathbf{e}_{\hat{\phi}} \right\}$$

$$\mathbf{E}(r, \theta) = \frac{\Omega - \omega}{\alpha\sqrt{\Sigma}} \left\{ \sqrt{\Delta} \Psi_{,r} \mathbf{e}_{\hat{r}} + \Psi_{,\theta} \mathbf{e}_{\hat{\theta}} + 0 \mathbf{e}_{\hat{\phi}} \right\}$$

$$\rho_e \mathbf{E} + \mathbf{J} \times \mathbf{B} = 0$$

The GR pulsar equation

$$\begin{aligned}
 & \left\{ \Psi_{,rr} + \frac{1}{\Delta} \Psi_{,\theta\theta} + \Psi_{,r} \left(\frac{A_{,r}}{A} - \frac{\Sigma_{,r}}{\Sigma} \right) - \frac{\Psi_{,\theta} \cos \theta}{\Delta \sin \theta} \right\} \cdot \left[1 - \frac{\omega^2 A \sin^2 \theta}{\Sigma} + \frac{4M\alpha\omega r \sin^2 \theta}{\Sigma} - \frac{2Mr}{\Sigma} \right] \\
 & - \left(\frac{A_{,r}}{A} - \frac{\Sigma_{,r}}{\Sigma} \right) \Psi_{,r} - \left(2 \frac{\cos \theta}{\sin \theta} - \frac{A_{,\theta}}{A} + \frac{\Sigma_{,\theta}}{\Sigma} \right) (\omega^2 A \sin^2 \theta - 4M\alpha\omega r \sin^2 \theta + 2Mr) \frac{\Psi_{,\theta}}{\Delta \Sigma} \\
 & + \frac{2Mr}{\Sigma} \left(\frac{A_{,r}}{A} - \frac{1}{r} \right) \Psi_{,r} + \frac{4\omega M\alpha r \sin^2 \theta}{\Sigma} \left\{ \Psi_{,r} \left(\frac{1}{r} - \frac{A_{,r}}{A} \right) - \frac{\Psi_{,\theta} A_{,\theta}}{\Delta A} \right\} \\
 & - \frac{\omega' \sin^2 \theta}{\Sigma} (\omega A - 2\alpha M r) \left(\Psi_{,r}^2 + \frac{1}{\Delta} \Psi_{,\theta}^2 \right) = -\frac{4\Sigma}{\Delta} I I'
 \end{aligned}$$

The GR pulsar equation

$$\begin{aligned}
 & \left\{ \Psi_{,rr} + \frac{1}{\Delta} \Psi_{,\theta\theta} + \Psi_{,r} \left(\frac{A_{,r}}{A} - \frac{\Sigma_{,r}}{\Sigma} \right) - \frac{\Psi_{,\theta} \cos \theta}{\Delta \sin \theta} \right\} \cdot \left[1 - \frac{\omega^2 A \sin^2 \theta}{\Sigma} + \frac{4M\alpha\omega r \sin^2 \theta}{\Sigma} - \frac{2Mr}{\Sigma} \right] \\
 & - \left(\frac{A_{,r}}{A} - \frac{\Sigma_{,r}}{\Sigma} \right) \Psi_{,r} - \left(2 \frac{\cos \theta}{\sin \theta} - \frac{A_{,\theta}}{A} + \frac{\Sigma_{,\theta}}{\Sigma} \right) (\omega^2 A \sin^2 \theta - 4M\alpha\omega r \sin^2 \theta + 2Mr) \frac{\Psi_{,\theta}}{\Delta \Sigma} \\
 & + \frac{2Mr}{\Sigma} \left(\frac{A_{,r}}{A} - \frac{1}{r} \right) \Psi_{,r} + \frac{4\omega M\alpha r \sin^2 \theta}{\Sigma} \left\{ \Psi_{,r} \left(\frac{1}{r} - \frac{A_{,r}}{A} \right) - \frac{\Psi_{,\theta} A_{,\theta}}{\Delta A} \right\} \\
 & - \frac{\omega' \sin^2 \theta}{\Sigma} (\omega A - 2\alpha M r) \left(\Psi_{,r}^2 + \frac{1}{\Delta} \Psi_{,\theta}^2 \right) = -\frac{4\Sigma}{\Delta} II'
 \end{aligned}$$

The GR pulsar equation

$$\left(\Psi_{,rr} + \frac{1}{r^2} \Psi_{,\theta\theta} + \frac{2\Psi_{,r}}{r} - \frac{1}{r^2} \frac{\cos \theta}{\sin \theta} \Psi_{,\theta} \right) \cdot [1 - \omega^2 r^2 \sin^2 \theta] - \frac{2\Psi_{,r}}{r} - 2\omega^2 \cos \theta \sin \theta \Psi_{,\theta} - \omega\omega' r^2 \sin^2 \theta \left(\Psi_{,r}^2 + \frac{1}{r^2} \Psi_{,\theta}^2 \right) = -4II'$$

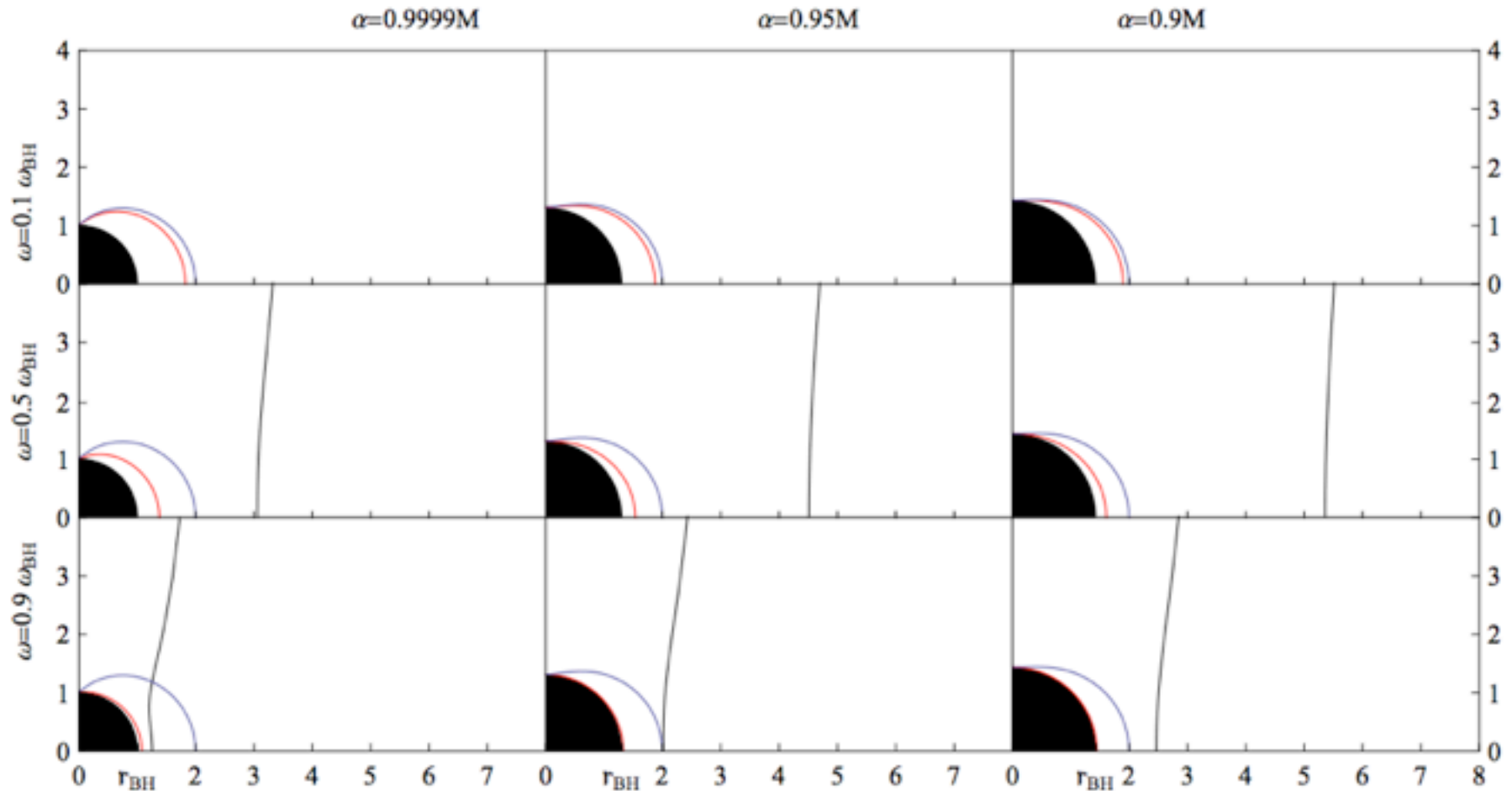
- The pulsar light cylinder: $r \sin \theta = c/\omega$
- The electric current $I(\Psi)$ must be determined self-consistently

The GR pulsar equation

$$\alpha^{-1}(\omega - \Omega)\varpi = \pm 1$$

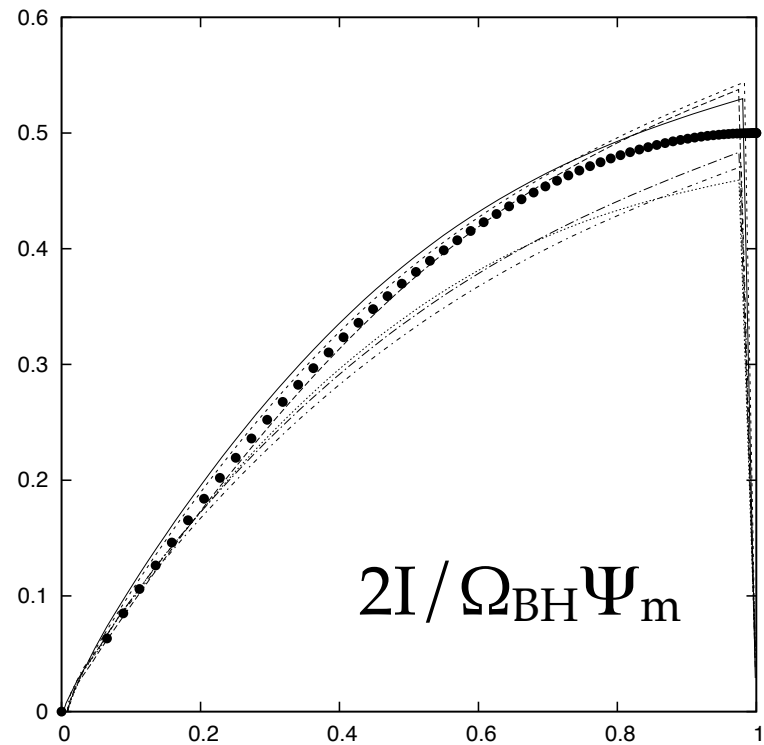
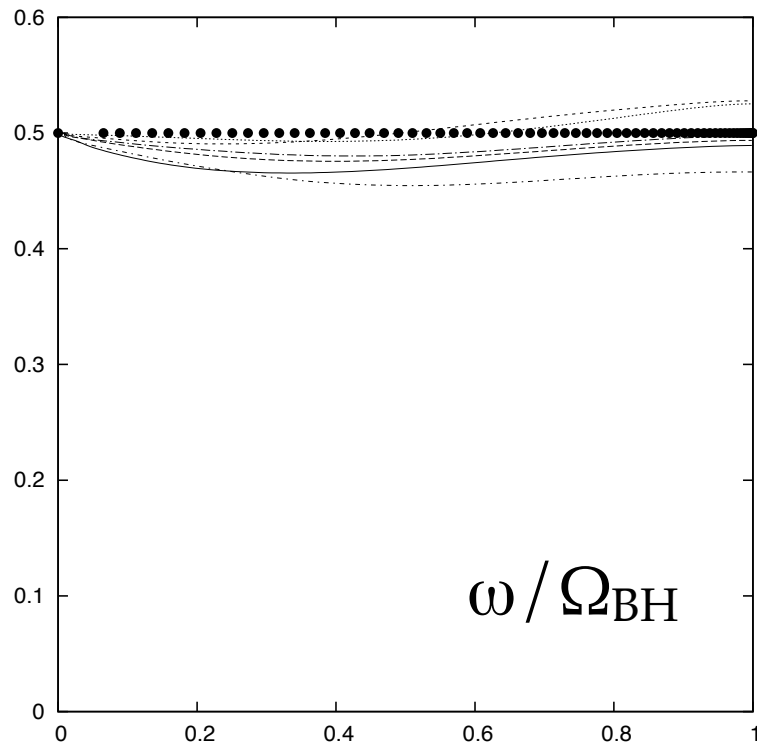
- The black hole possesses **two light surfaces**
- The electric current $I(\Psi)$ must be determined self-consistently together with the angular velocity of the magnetic field $\omega(\Psi)$

The GR pulsar equation

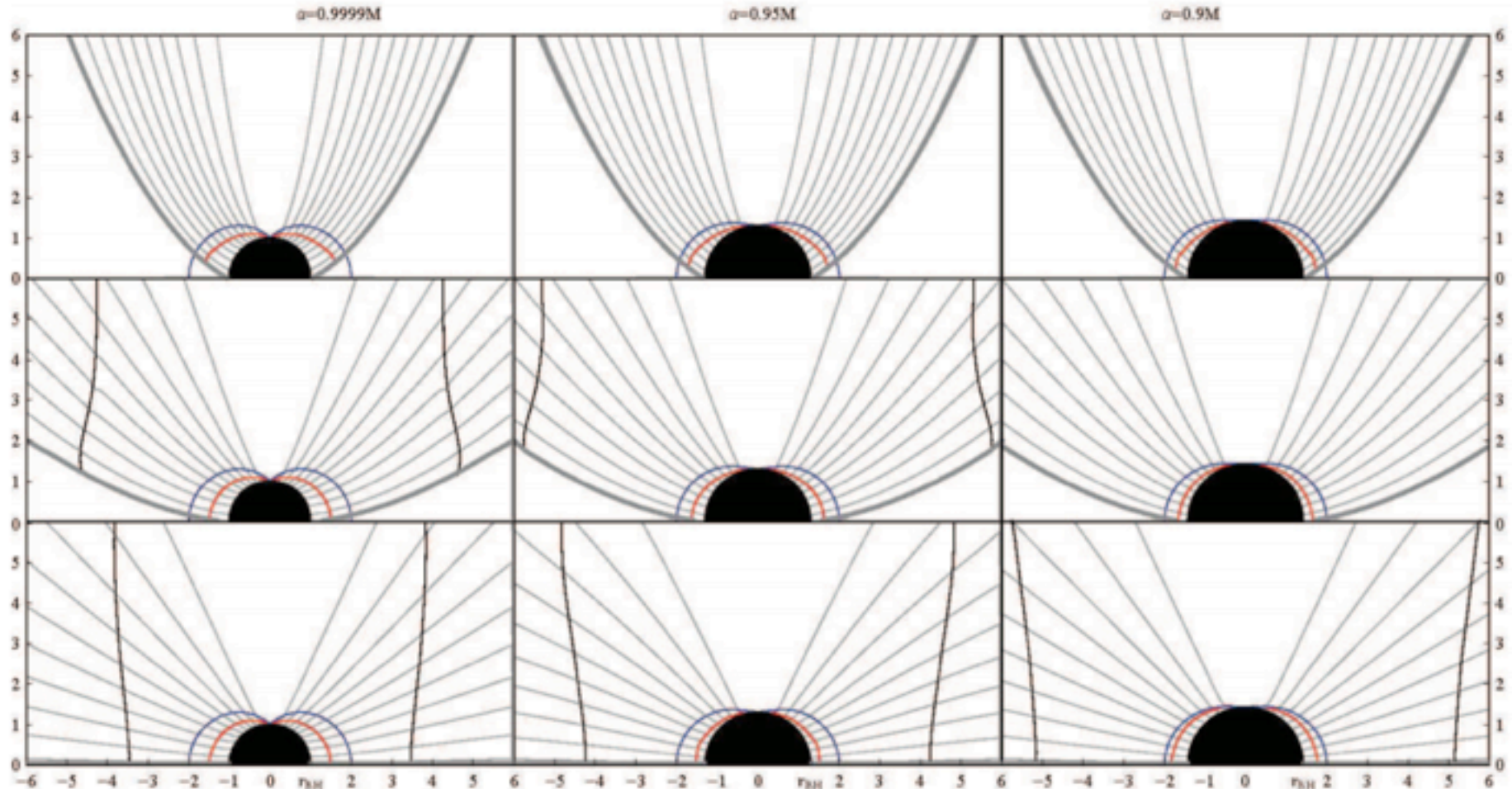


The GR pulsar equation

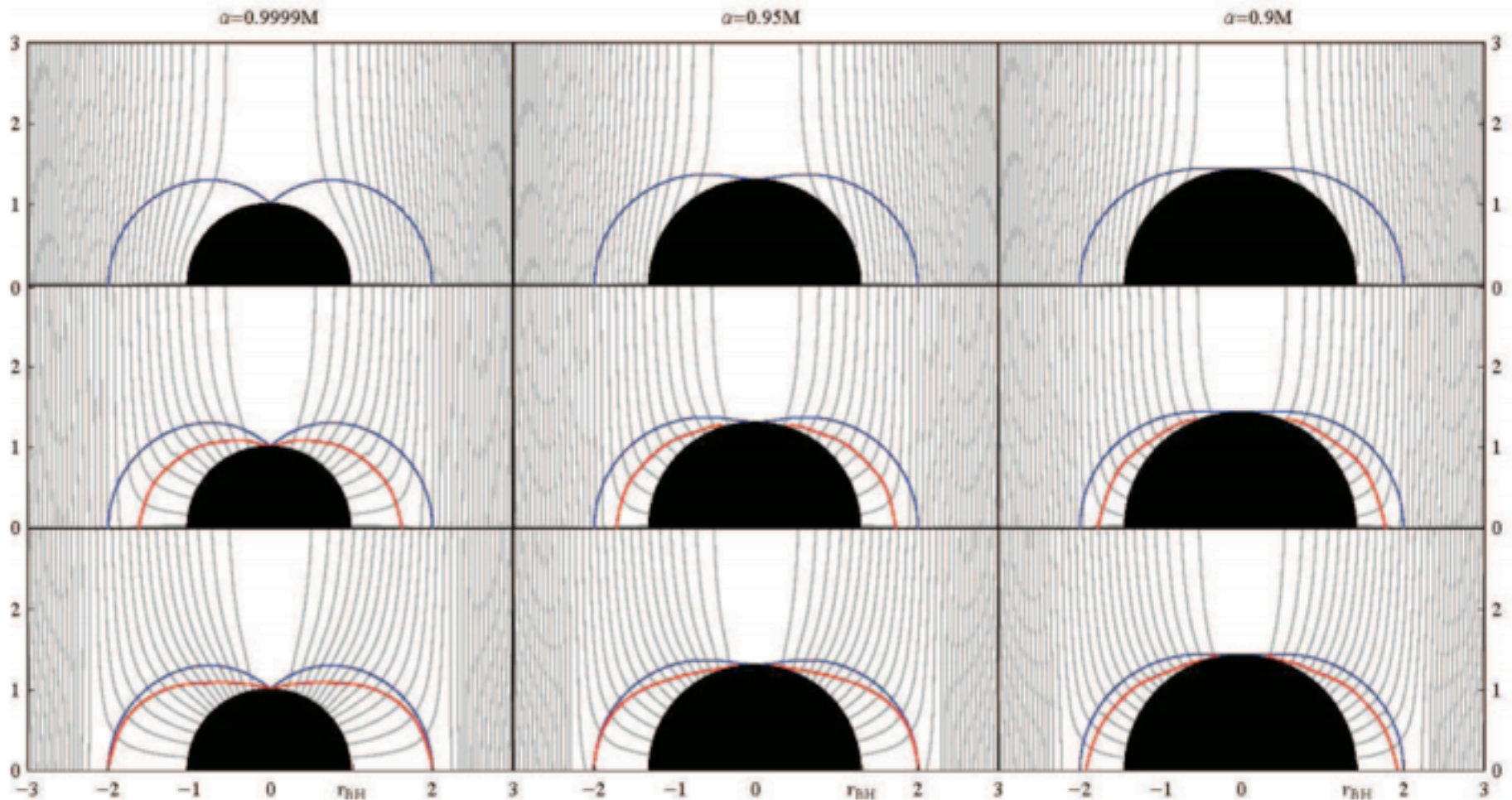
$$\alpha=0.7-1 \text{ M}, \omega \sim 0.5 \Omega_{\text{BH}}$$



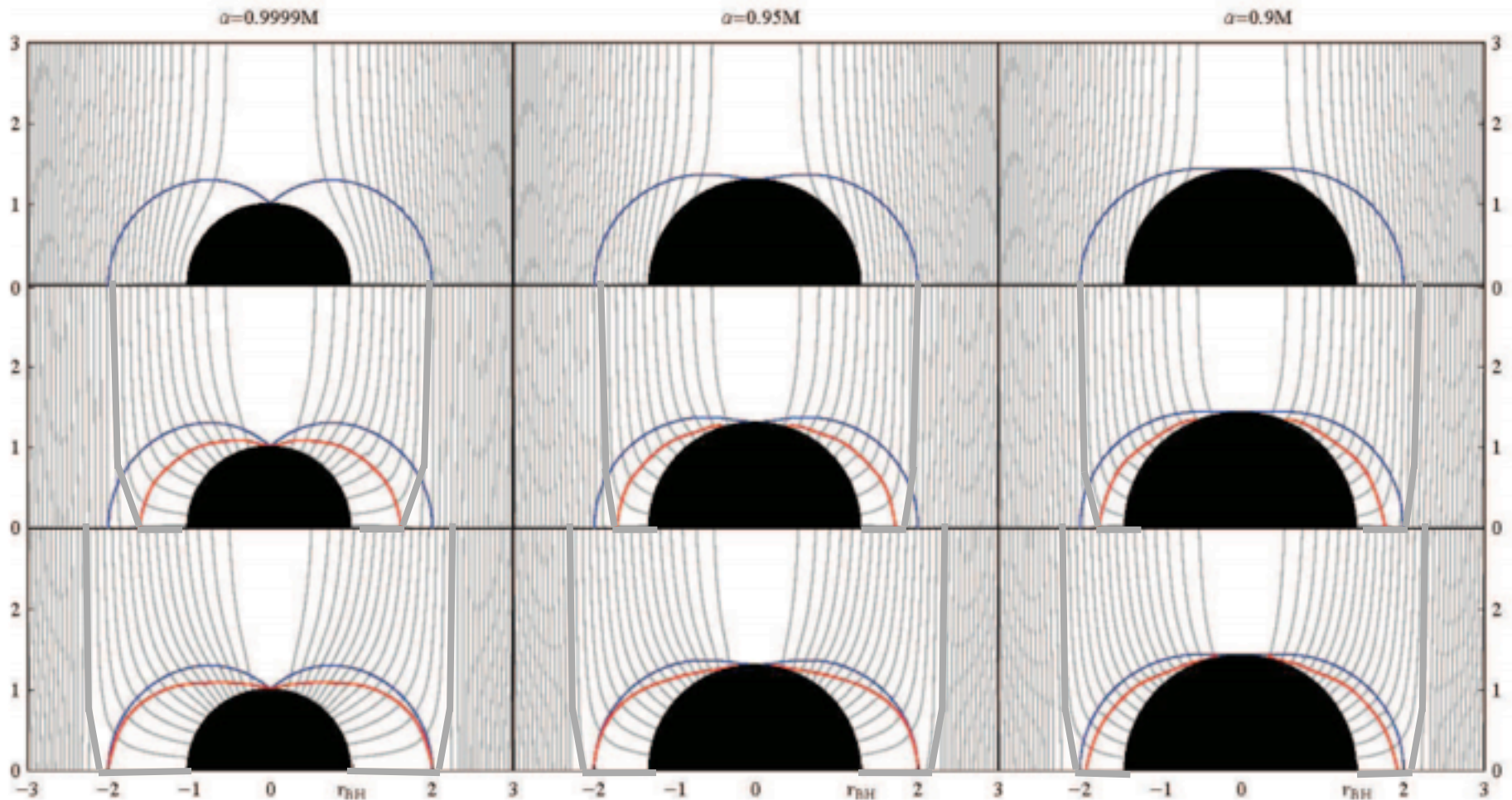
The GR pulsar equation



The GR pulsar equation



The GR pulsar equation



The GR pulsar equation

$$E^\theta = \pm B^\phi \text{ at 'infinity'}$$

$$I(\Psi) = -0.5\omega(\Psi)\Psi_{,\theta}\sin\theta$$

Infinity

$$I(\Psi) = -0.5(\Omega_{\text{BH}} - \omega(\Psi))\frac{\sqrt{A}}{\Sigma}\Psi_{,\theta}\sin\theta$$

Horizon

The GR pulsar equation

$$E^\theta = \pm B^\phi \text{ at 'infinity'}$$

$$I(\Psi) = -0.5\omega(\Psi)\Psi_{,\theta} \sin \theta$$

$$4II' = \left(\omega^2 \Psi_{,\theta\theta} + \omega^2 \frac{\cos \theta}{\sin \theta} \Psi_{,\theta} + \omega \omega' \Psi_{,\theta}^2 \right) \sin^2 \theta$$

Infinity

$$I(\Psi) = -0.5(\Omega_{\text{BH}} - \omega(\Psi)) \frac{\sqrt{A}}{\Sigma} \Psi_{,\theta} \sin \theta$$

$$4II' = \left((\omega - \Omega_{\text{BH}})^2 \Psi_{,\theta\theta} + (\omega - \Omega_{\text{BH}})^2 \frac{\cos \theta}{\sin \theta} \Psi_{,\theta} - \frac{\Sigma_{,\theta}}{\Sigma} \Psi_{,\theta} + \omega' (\omega - \Omega_{\text{BH}}) \Psi_{,\theta}^2 \right) \frac{A}{\Sigma^2} \sin^2 \theta$$

Horizon

Black hole spindown

Reducible rotational BH energy

$$E \sim M r_{\text{BH}}^2 \Omega^2 \sim 10^{53} M / M_{\odot} \text{ erg}$$

Black hole spindown

Reducible rotational BH energy

$$E \sim M r_{\text{BH}}^2 \Omega^2 \sim 10^{53} M/M_{\odot} \text{ erg}$$

+ magnetic field

Black hole spindown

$$\dot{E}_{\text{BZ}} \sim -\Psi_m^2 \Omega^2 \sim -\Omega^2$$

Black hole spindown

$$\dot{E}_{\text{BZ}} \sim -\Psi_m^2 \Omega^2 \sim -\Omega^2$$

$$\dot{E} \sim M r_{\text{BH}}^2 \Omega \dot{\Omega}$$

Black hole spindown

$$\dot{E}_{\text{BZ}} \sim -\Psi_m^2 \Omega^2 \sim -\Omega^2$$

$$\dot{E} \sim M r_{\text{BH}}^2 \Omega \dot{\Omega}$$

$$\sim e^{-t/\tau_{\text{BZ}}}$$

Black hole spindown

$$\dot{E}_{\text{BZ}} \sim -\Psi_m^2 \Omega^2 \sim -\Omega^2$$

$$\dot{E} \sim M r_{\text{BH}}^2 \Omega \dot{\Omega}$$

$$\sim e^{-t/\tau_{\text{BZ}}}$$

$$B_o \sim 10^7 \text{ G} \qquad \tau_{\text{BZ}} \sim 1000 \text{ Gyr}$$

BZ simulations in stationary spacetime (fixed BH spin)

How about
Gamma-ray bursts?

How about
Gamma-ray bursts?

Black hole spindown

$$\dot{E}_{\text{BZ}} \sim -\Psi_m^2 \Omega^2 \sim -\Omega^2$$

$$\dot{E} \sim M r_{\text{BH}}^2 \Omega \dot{\Omega}$$

$$\sim e^{-t/\tau_{\text{BZ}}}$$

$$B_o \sim 10^{15} \text{ G} \qquad \tau_{\text{BZ}} \sim 10^3 - 10^4 \text{ sec}$$

Collapsar simulations in stationary spacetime (fixed BH spin)

Black hole spindown

$$\dot{E}_{\text{BZ}} \sim -\Psi_m^2 \Omega^2 \sim -\Omega^2$$

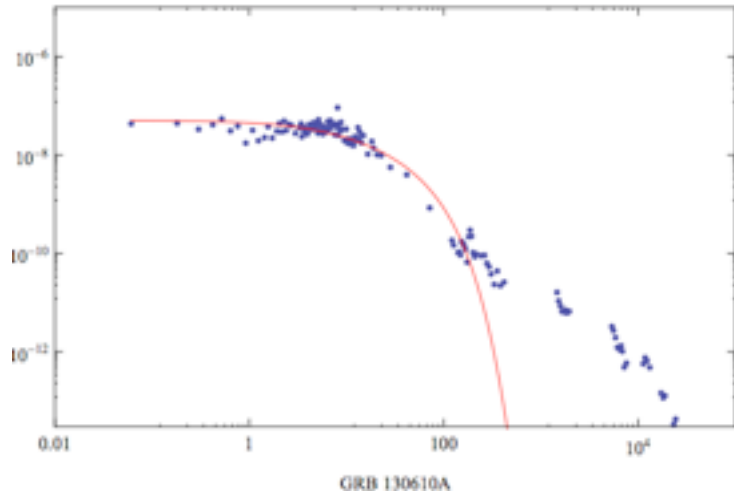
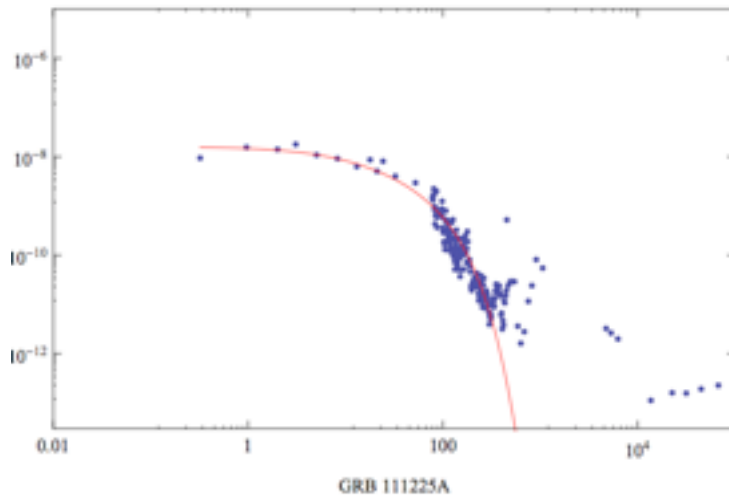
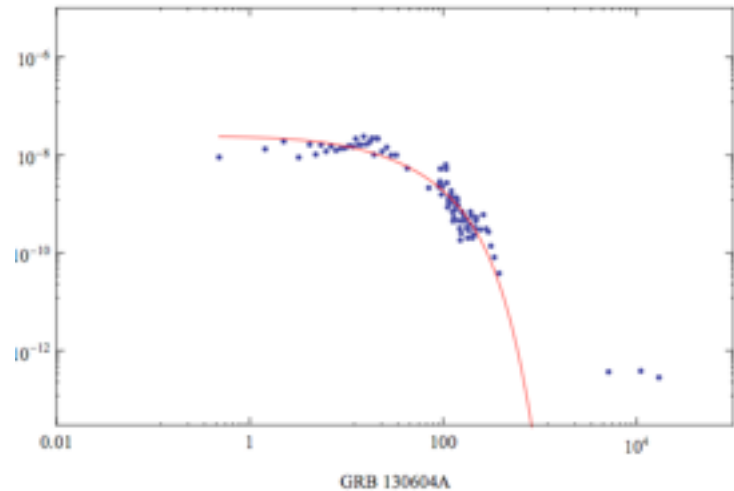
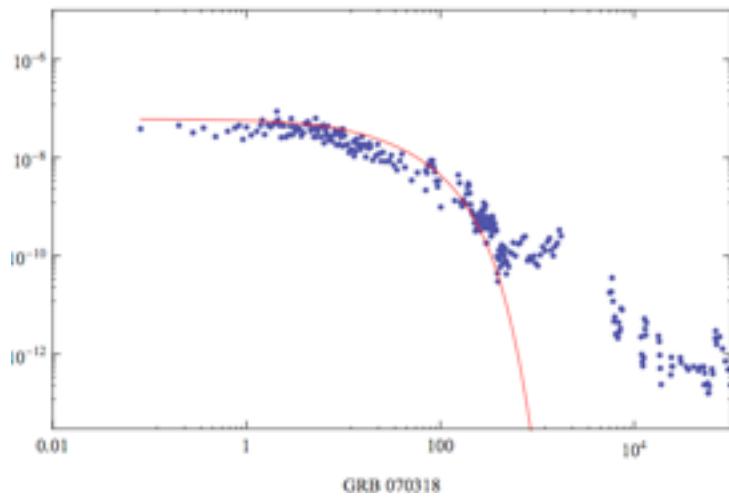
$$\dot{E} \sim M r_{\text{BH}}^2 \Omega \dot{\Omega}$$

$$\sim e^{-t/\tau_{\text{BZ}}}$$

$$B_o \sim 10^{16} \text{ G} \qquad \tau_{\text{BZ}} \sim 10 - 100 \text{ sec}$$

The “orthogonal” GRB: Contopoulos, Nathanail & Pugliese 2014

Probe of BH magnetosphere

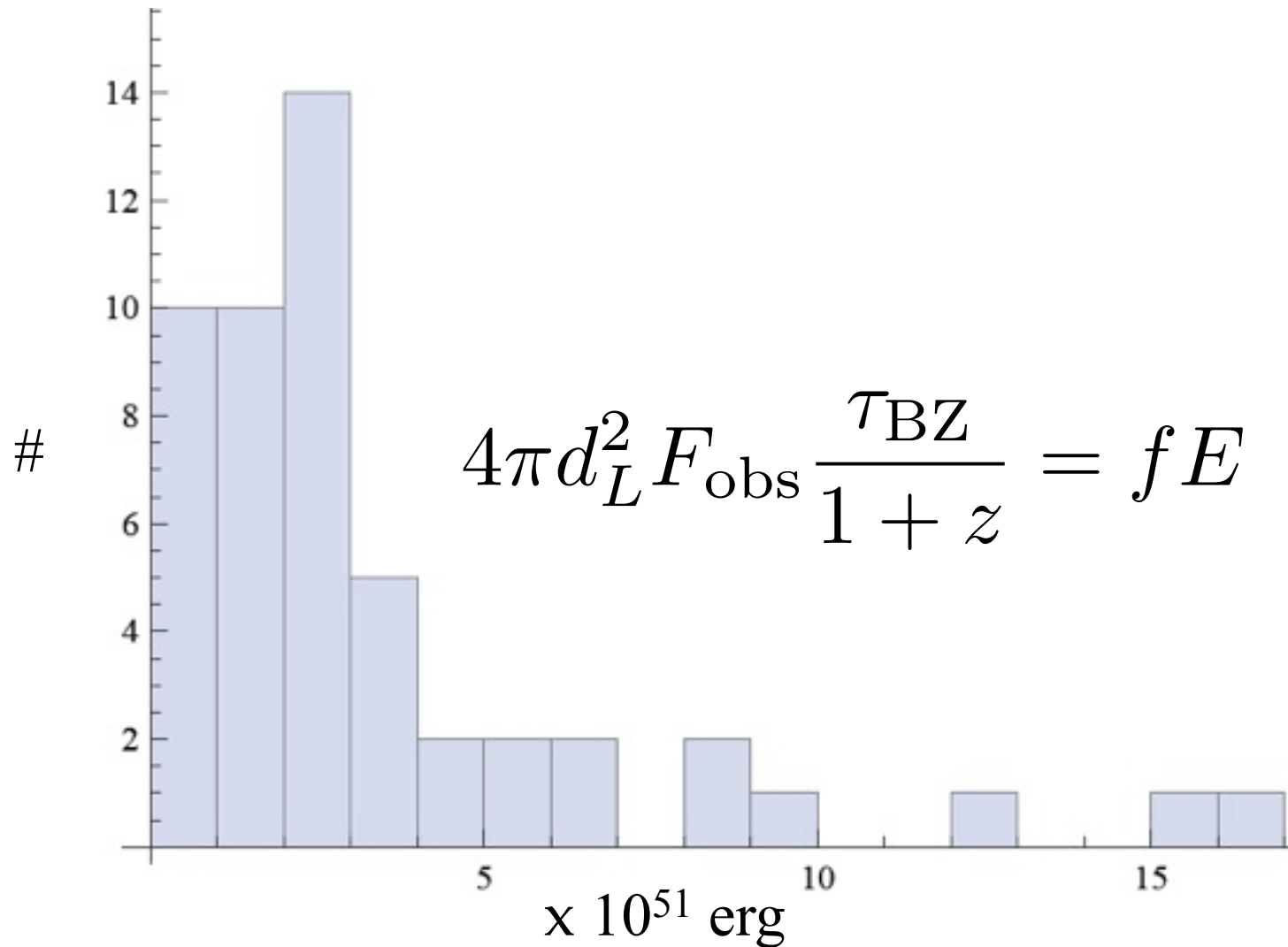


Contopoulos, Nathanail & Pugliese 2014

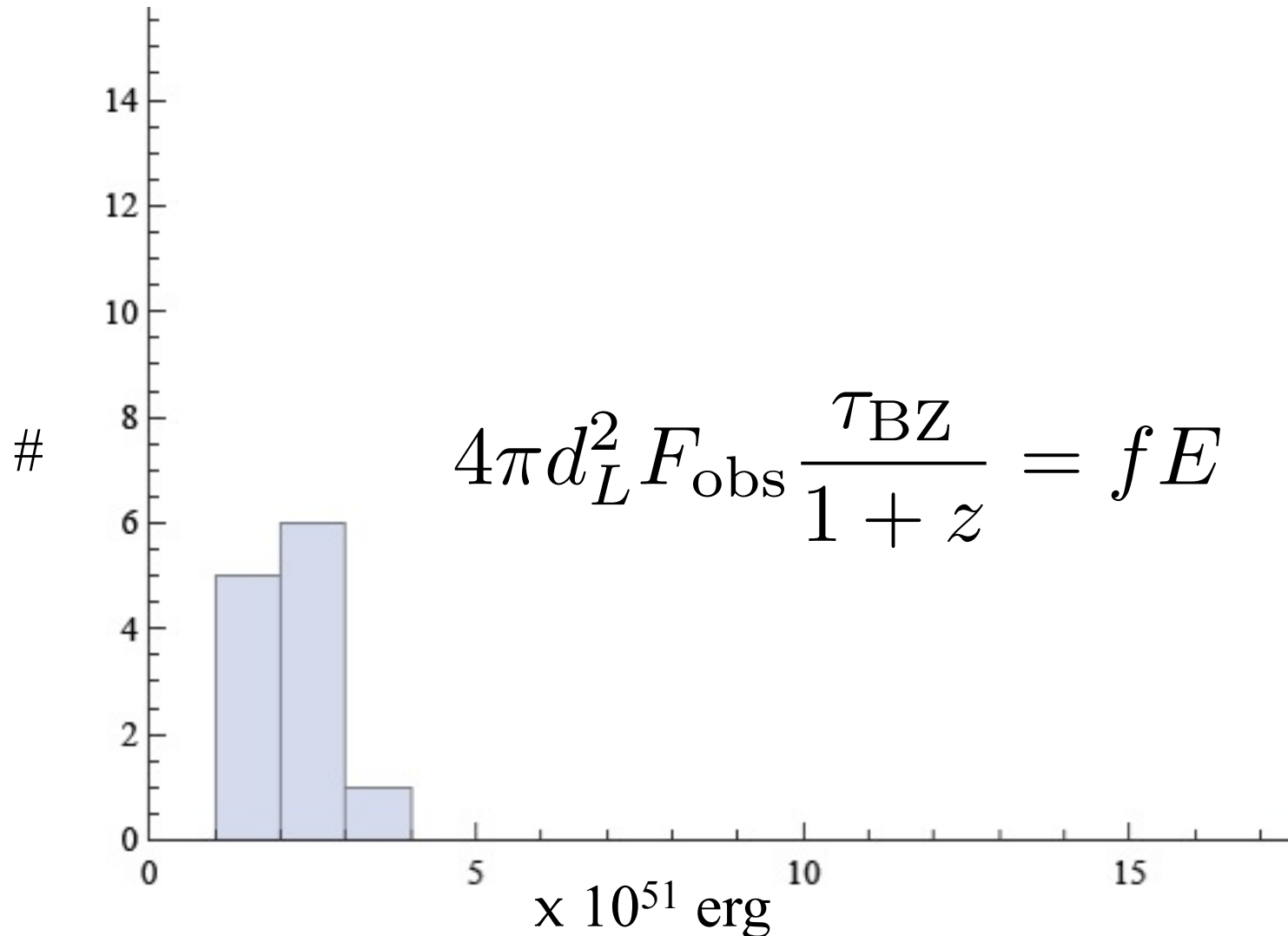
Probe of BH magnetosphere

$$4\pi d_L^2 F_{\text{obs}} \frac{\tau_{\text{BZ}}}{1+z} = f E$$

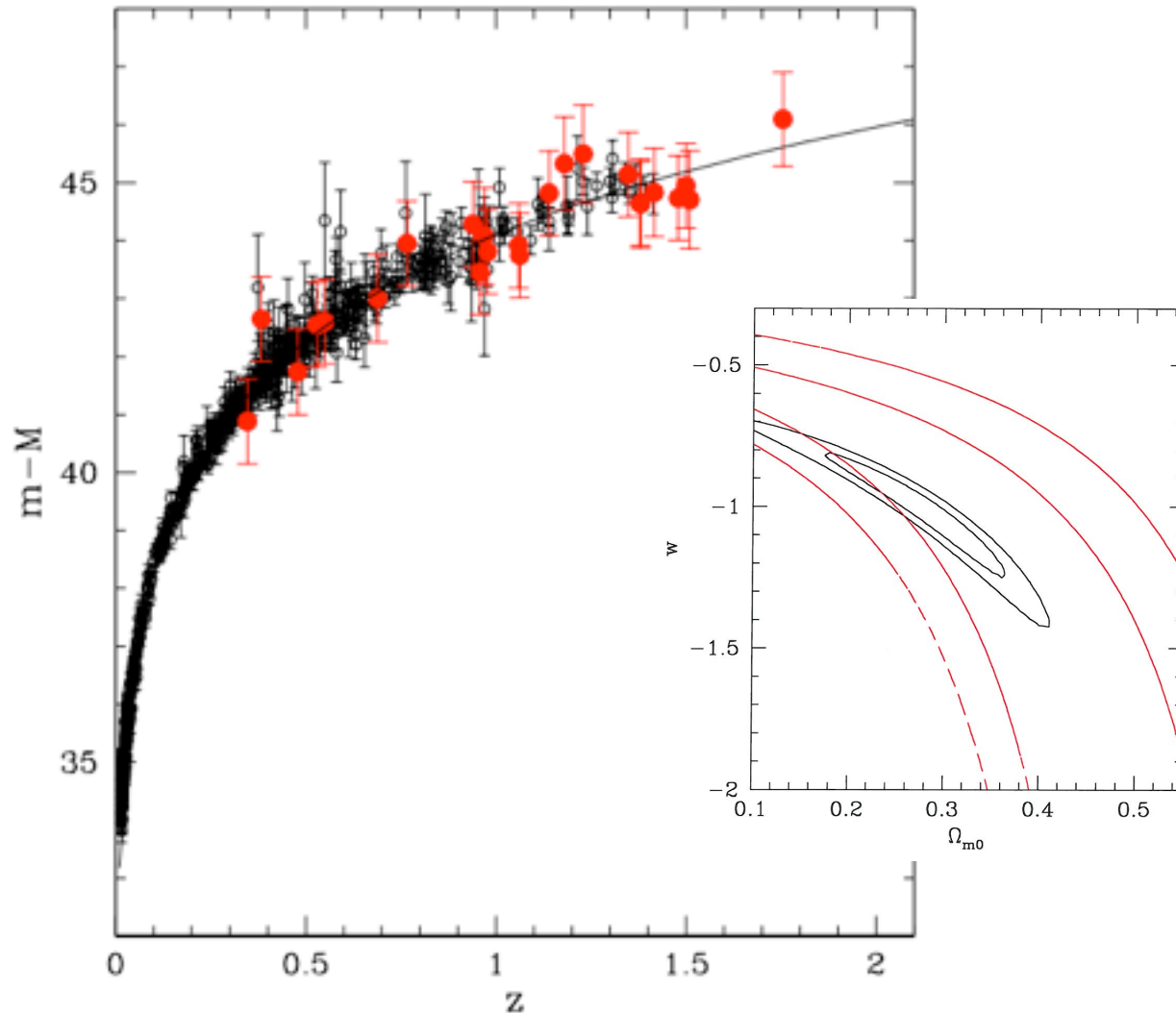
Probe of BH magnetosphere



Are GRBs standard candles?



Are GRBs standard candles?



Summary

- The Cosmic Battery: astrophysical magnetic fields generated at the inner edge of the accretion disk. The field is held by the inner disk. The return field diffuses outward through the outer disk.

Summary

- Black hole and pulsar magnetospheres:
 - The “membrane” does not teach us much
 - Inner Light Surface
 - Electric current sheet
 - Orthogonal emitters

Summary

- GRBs as standard candles in Cosmology?

Thank you



European Union
European Social Fund



Co- financed by Greece and the European Union

