

# Physics 56400 Introduction to Elementary Particle Physics I

Lecture 7
Fall 2019 Semester

Prof. Matthew Jones

# **Liquid Scintillator**

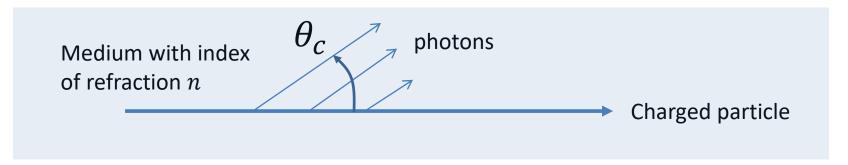
- Scintillating chemicals can be dissolved in, for example, purified mineral oil
- Can be used to fill a very large volume
- Might not be rad-hard
  - Chemical impurities can be created by ionizing radiation that reduce light output or decrease attenuation length
  - You might end up with a very large volume of radioactive waste
  - Disposal might be expensive...

#### **Particle Detectors**

- Detecting charged particles using the scintillation process
  - Photon yield is proportional to the ionization energy loss
  - Photons are converted to photo-electrons at the cathode of a photomultiplier tube
  - Photomultiplier tubes amplify the photo-current
  - Alternatively, solid state detectors serve the same purpose
- Charged particles can also emit Cherenkov light

#### **Cherenkov Radiation**

 When a charged particle moves through a medium with a speed that is faster than the speed of light in the medium, it emits Cherenkov radiation



$$\cos \theta_c = \frac{1}{n\beta}$$

$$\frac{dN}{dxd\lambda} = \frac{2\pi\alpha z^2}{\lambda^2} \left( 1 - \frac{1}{\beta^2 n^2(\lambda)} \right)$$

# **Cherenkov Light**

• Consider a muon with an energy of 10 GeV moving through water (n=1.33)

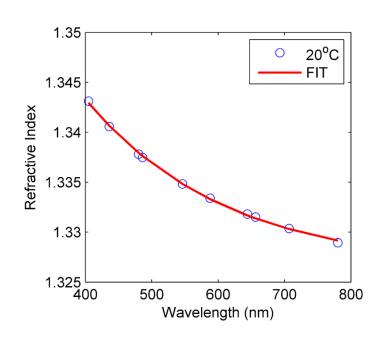
$$\beta = \frac{p}{E} = \frac{\sqrt{E^2 - m^2}}{E} = \sqrt{1 - m^2/E^2} = 0.99994$$

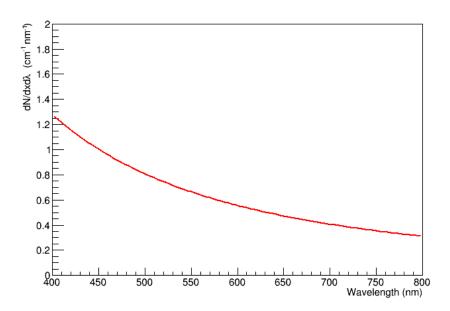
Angle of Cherenkov light emission:

$$\cos \theta_c = \frac{1}{n\beta} = \frac{1}{1.33} = 0.75$$

- Cherenkov light is emitted immediately
  - Scintillation light is slower and can have long components
- How many photons are produced?

# **Cherenkov Light**





Light yield integrated over 400-800 nm:

$$\frac{dN}{dx} = 252 \text{ cm}^{-1}$$

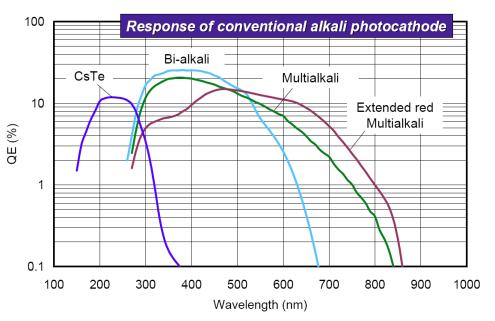
If the muon travels through 10 meters of water, then

$$N = 252,000$$

Much less light per unit length than for scintillator.

# **Cherenkov Light**

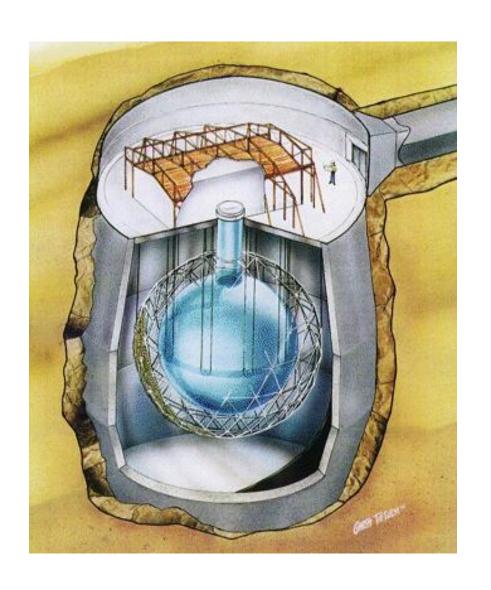
Spectral response of photomultiplier tubes:

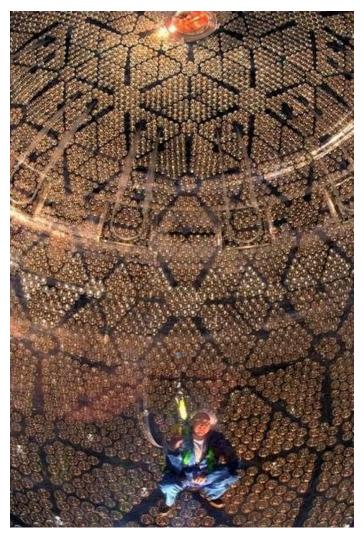




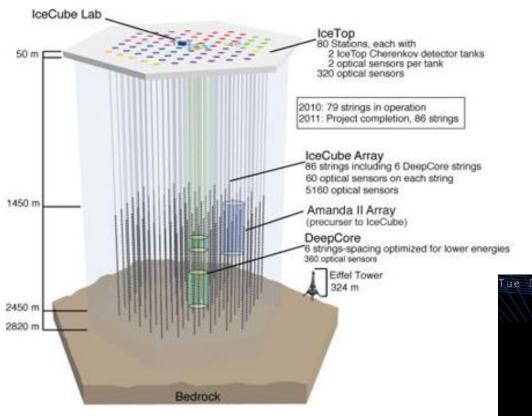
- Average quantum efficiency is about 5-10%
- Light collection efficiency might be 20%
- Number of photoelectrons: 2500 5000.

#### **Water Cherenkov Detectors**

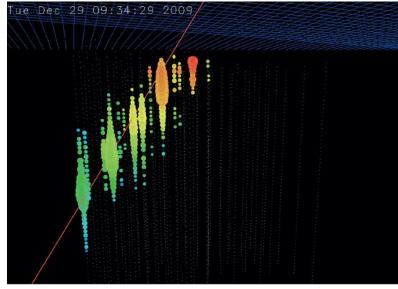




## **Large Cherenkov Detectors**



Color indicates time (red is early, green is late)
Size indicates light yield



#### **Threshold Cherenkov Counters**

- If  $\beta < 1/n$  then no light is emitted.
- Thin volumes of gas with  $n \approx 1$  can be used to distinguish between pions (m=139.5 MeV) and kaons (m=493.7 MeV)
- Suppose a magnet selects charged particles with momentum p < 2 GeV.

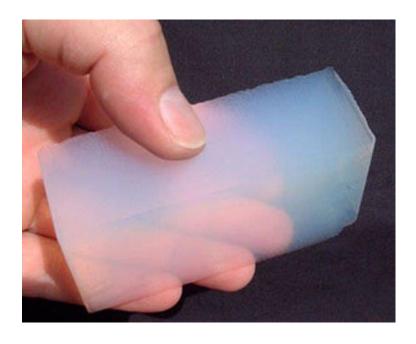
$$\beta_{\pi} = 1/\sqrt{1 + m_{\pi}^2/p^2} < 0.9976$$

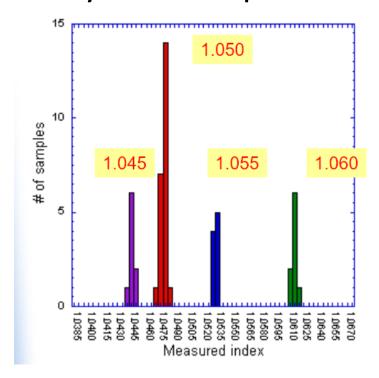
$$\beta_{K} = 1/\sqrt{1 + m_{K}^2/p^2} < 0.9709$$

• We need a material with  $n \approx 1.05$  so that pions will emit Cherenkov light but kaons won't.

#### **Threshold Cherenkov Counters**

Such materials do exist and they are transparent

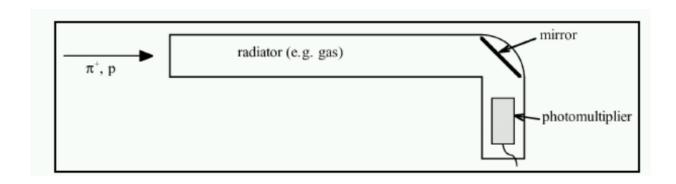


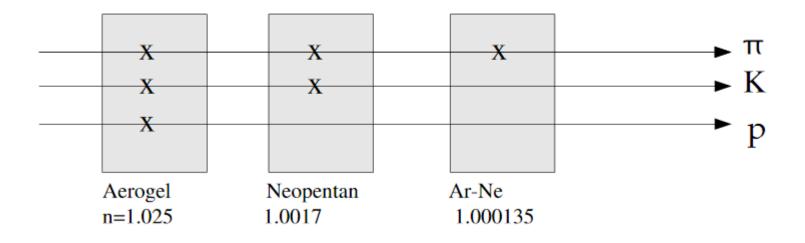


• See, for example:

https://www.youtube.com/watch?v=SWnqOc-cWpU

#### **Threshold Cherenkov Counters**

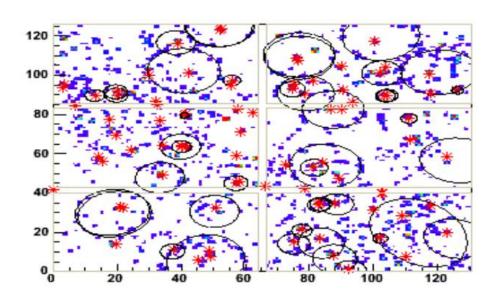


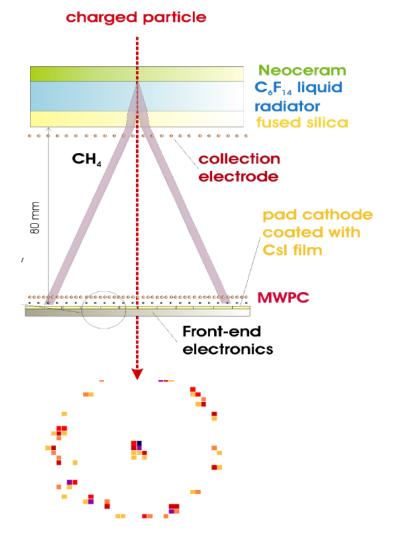


# **Ring Imaging Cherenkov Detectors**

The radius of the ring determines the angle of light emission,  $\theta_c$  which gives the velocity:

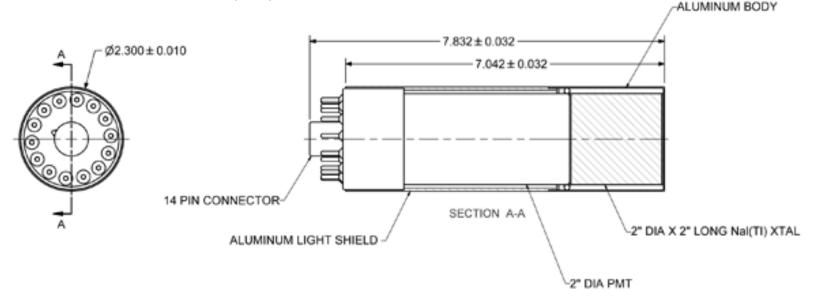
$$\beta = \frac{1}{n \cos \theta_c}$$





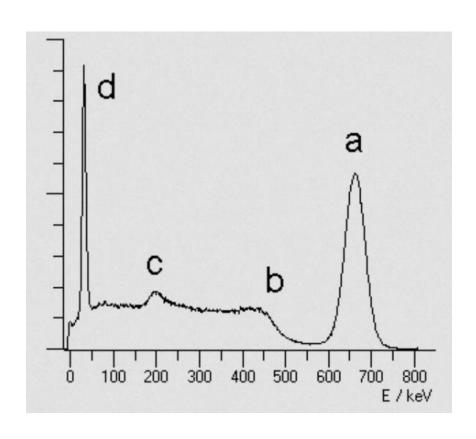
#### **Detection of Photons**

- After Compton scattering, we can detect the ionization of the scattered electron
- If the photon is above threshold, we can detect both the electron and positron after pair production
- Consider a NaI(TI) counter:



## **Gamma Ray Spectrum**

Cs-137 spectrum:  $E_{\gamma} = 662 \text{ keV}$ 



- a) All of the photon's energy is absorbed within the crystal.
- b) Compton edge: photon scatters out of the crystal.
- c) Backscatter peak.

## **Gamma Ray Spectrum**

Photon energy:

$$\frac{1}{E'} - \frac{1}{E} = \frac{1 - \cos \theta}{m_e}$$

Electron energy:

$$\frac{\Delta E}{E} = \frac{E - E'}{E} = 1 - \frac{1}{1 + E/m_e(1 - \cos \theta)}$$

• When the photon back-scatters at  $\theta = \pi$ ,

$$\frac{\Delta E}{E} = 0.7215$$

• The Compton edge is located at  $E_c = 478 \text{ keV}$ 

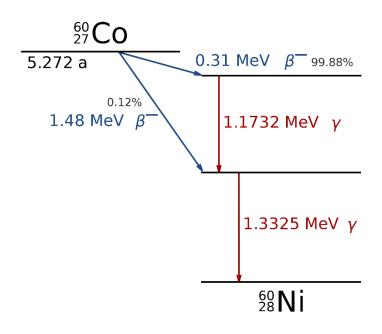
#### **Gamma Ray Spectrum**

 The backscatter peak results from the gamma ray scattering off of passive detector material back into the crystal, where all its energy is deposited.

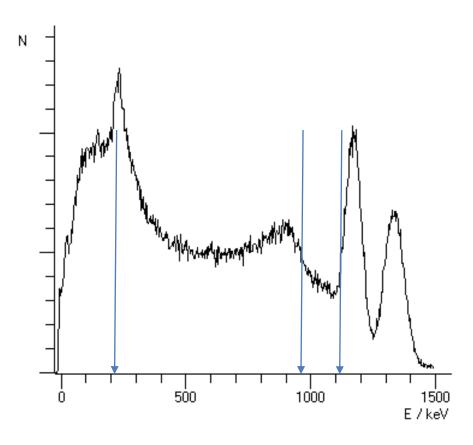
$$E' = \frac{m_e}{(1 - \cos \theta) + m_e/E}$$

- When  $\theta = \pi$ , E' = 184.4 keV
- In the limit as  $E \gg m_e$ ,  $E' \to m_e/2$

# Other Gamma Ray Spectra



Compton edges are at  $E_1 = 963 \text{ keV}$   $E_2 = 1118 \text{ keV}$  Backscatter peak is at  $E_b = 210 \text{ keV}$ 

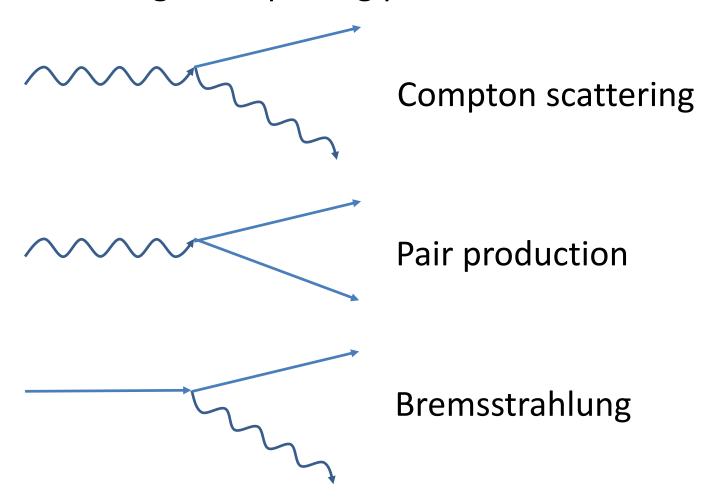


## **Electromagnetic Calorimeters**

- If we want to measure the energy of a high energy photon, we need it to Compton scatter or pair produce
- We also need to contain all the scattered particles within the detector and measure their energy.
- A simple model for electron/photon interactions with matter is the "electromagnetic shower".
- The mean free path of a high energy electron or photon in material is about one radiation length.
  - Example: For lead,

$$\frac{X_0}{\rho} = \frac{6.37 \text{ g} \cdot \text{cm}^{-2}}{11.350 \text{ g} \cdot \text{cm}^{-3}} = 0.56 \text{ cm}$$

• Electromagnetic splitting processes:



- Sequential branchings can be modeled as a stochastic process:
  - 1. Propagate particle a distance x with probability distribution

$$p(x) = \frac{1}{X_0} e^{-x/X_0}$$

2. Split into two new particles with energy fraction y distributed according to some probability distribution

$$f(y) = 1$$
 (uniform distribution)

One particle has energy  $E \cdot y$  and the other has energy  $E \cdot (1-y)$ 

- 3. Repeat until the energy of a particle falls below the critical energy cutoff,  $E_c$ , which is (roughly) the energy at which  $E < X_0 \cdot \left(\frac{dE}{dx}\right)_{min}$
- These are essentially the main ideas behind Monte Carlo calculations like GEANT and EGS.

## **Critical Energy**

 The critical energy depends on the material but can be parameterized as a function of Z:

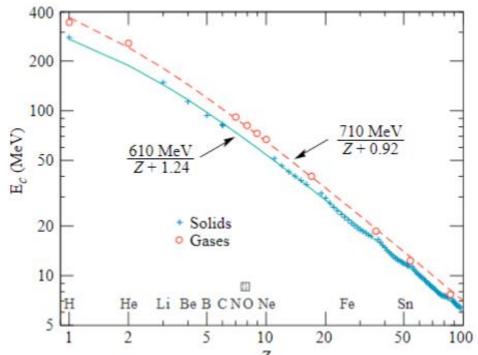
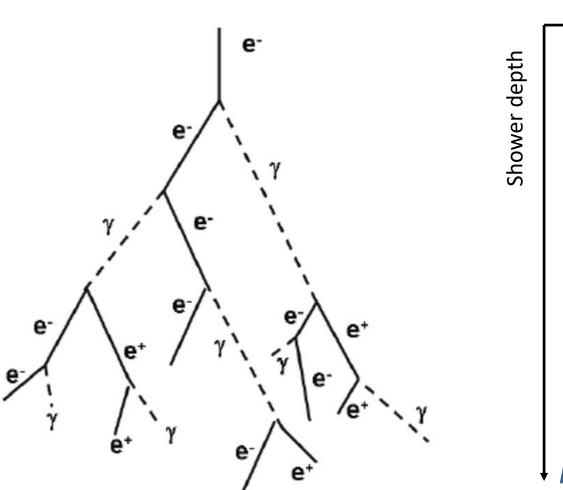
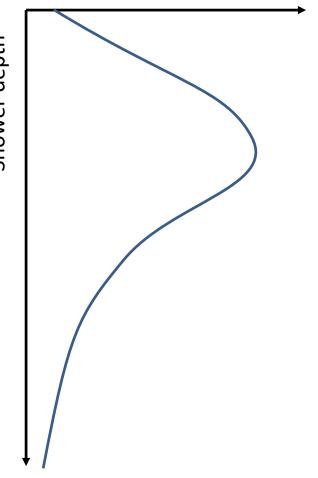


Figure 33.14: Electron critical energy for the chemical elements, using Rossi's definition [2]. The fits shown are for solids and liquids (solid line) and gases (dashed line). The rms deviation is 2.2% for the solids and 4.0% for the gases. (Computed with code supplied by A. Fassó.)



Number of particles



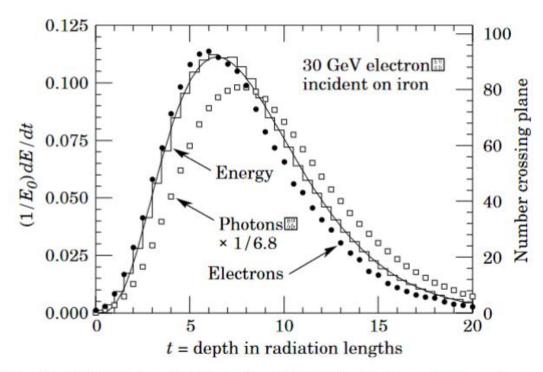


Figure 33.20: An EGS4 simulation of a 30 GeV electron-induced cascade in iron. The histogram shows fractional energy deposition per radiation length, and the curve is a gamma-function fit to the distribution. Circles indicate the number of electrons with total energy greater than 1.5 MeV crossing planes at  $X_0/2$  intervals (scale on right) and the squares the number of photons with  $E \geq 1.5$  MeV crossing the planes (scaled down to have same area as the electron distribution).

• The longitudinal profile of an electromagnetic shower can be described empirically by a gamma distribution:

$$\frac{dE}{dt} = E_0 b \frac{(bt)^{a-1} e^{-bt}}{\Gamma(a)}$$

The maximum occurs at

$$t_{max} = \frac{a-1}{b} = \log \frac{E}{E_c} \pm 0.5$$

for electrons (-) and photons (+).

The transverse size of the shower scales with the Moliere radius,

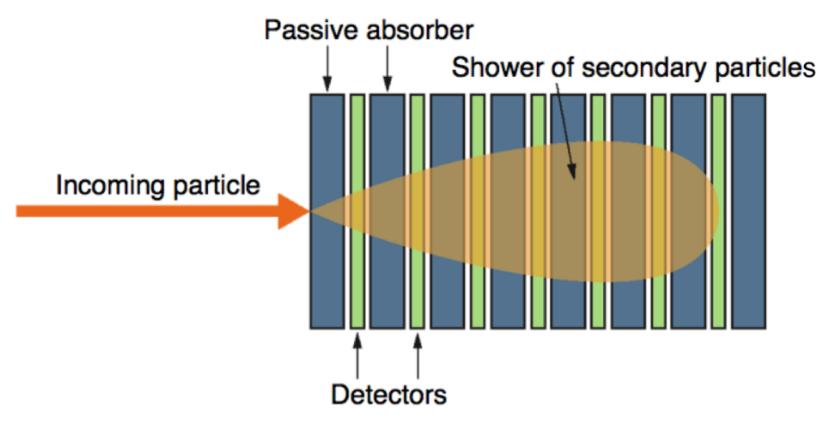
$$R_M = X_0 \left( \frac{21 \text{ MeV}}{E_c} \right)$$

- 90% of the shower energy is contained within a cylinder of radius  $R_M$
- 99% of the shower energy is contained within  $3.5 R_M$

#### **Electromagnetic Calorimeters**

- Electromagnetic calorimeters should be several radiation lengths thick to contain the highest energy showers.
- Short  $X_0$  will also restrict the transverse size of showers, facilitating transverse segmentation.
- The ionization energy deposited by electrons and positrons in the shower is proportional to the incident energy.

# **Sampling Calorimeters**



Energy resolution is dominated by the number of photoelectrons detected:

$$\frac{\sigma_E}{E} \propto \frac{\sigma_N}{N} \approx \frac{\sqrt{N}}{N} = \frac{1}{\sqrt{N}} \sim \frac{1}{\sqrt{E}}$$

#### **Sampling Calorimeters**

- Sampling calorimeters can be made rather compact
  - Lead/scintillator,  $X_0/\rho \sim 0.56$  cm
  - Tungsten/silicon,  $X_0/\rho \sim 0.35$  cm
  - A depth of 20 radiation lengths is only 11.2 cm or 7 cm, plus the thickness of scintillator.
- Sampling calorimeters can be constructed with some very clever geometries.
- Not all the energy deposited in the detector is measured which degrades the energy resolution.
- A homogeneous calorimeter is sensitive to all ionization deposited in the calorimeter volume.

#### **Homogeneous Calorimeters**

• If a scintillator is dense enough and large enough, it can contain the entire electromagnetic shower.

#### Examples:

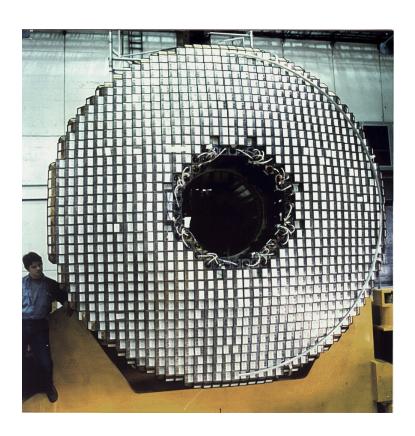
- Bismuth germinate ( $Bi_4Ge_3O_{12}$ ) BGO,  $\rho = 7.13 \text{ g} \cdot \text{cm}^{-3}$
- Barium Fluoride (BaF<sub>2</sub>),  $\rho = 4.88 \,\mathrm{g \cdot cm^{-3}}$
- Cerium Fluoride (CeF<sub>3</sub>),  $\rho = 6.16 \,\mathrm{g\cdot cm^{-3}}$
- Lead Tungstate (PbWO<sub>4</sub>),  $\rho = 8.28 \,\mathrm{g \cdot cm^{-3}}$



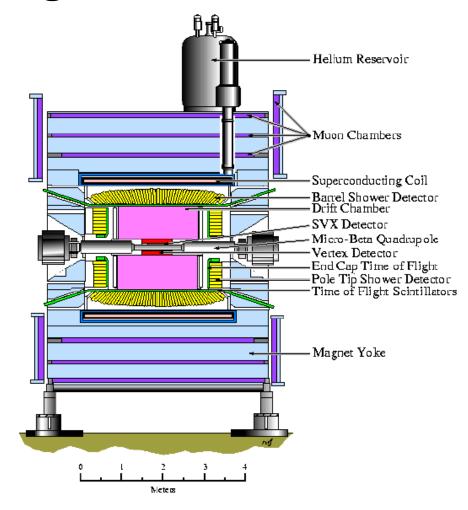
#### **Lead Glass Calorimeters**

- Lead glass is very dense ( $\rho \sim 4~{\rm g\cdot cm^{-3}}$ )
- It is NOT a scintillator
- Its index of refraction is about 1.7
- Any electron with  $\beta > 0.588 c$  will emit Cherenkov light
- This corresponds to  $E_c \sim 0.63 \text{ MeV}$

#### **Examples of Electromagnetic Calorimeters**



OPAL electromagnetic end-cap calorimeter (at CERN in the 1990's)

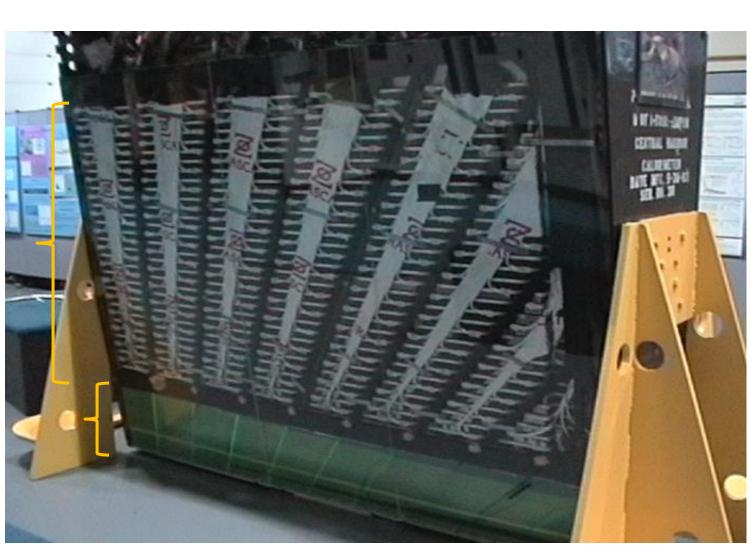


CLEO II Detector (at Cornell in the 1990's)

#### **Examples of Electromagnetic Calorimeters**

Hadronicsection

Electromagnetic section



#### **Examples of Electromagnetic Calorimeters**

• A "spaghetti" calorimeter:





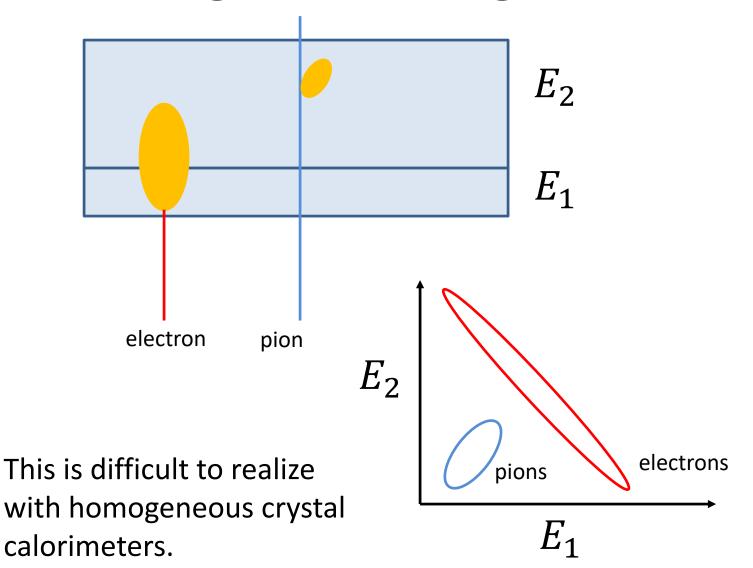
(ALICE zero-degree calorimeter)

- Quartz fibers are embedded between metal plates.
- Electrons emit Cherenkov light which propagates to the ends of the fibers

# **Longitudinal Segmentation**

- Even limited longitudinal segmentation can be very useful for electron identification.
- Both electrons and charged pions produce tracks leading into the calorimeter
- Electrons will shower but pions will mostly just deposit ionization energy
- A calorimeter with two layers can help to distinguish between electrons and pions.

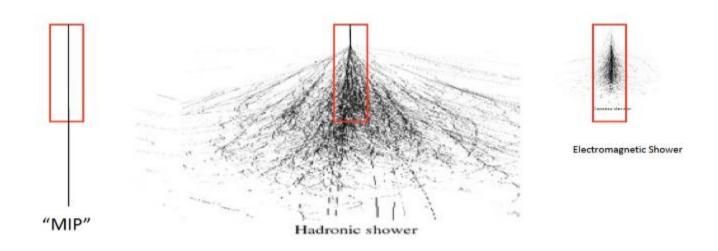
# **Longitudinal Segmentation**



#### **Hadronic Calorimeters**

- Same idea except that the shower size scales with the nuclear interaction length,  $\lambda_I$  which is much longer than  $X_0$ .
- Hadronic showers are not as uniform as electromagnetic showers.
  - Fewer secondary particles
  - More random secondary particle production
  - Large fluctuations in deposited energy
  - Multiple scattering

#### **Hadronic Calorimeters**



- Fluctuations in the number of  $\pi^0$  produced lead to large observed energy fluctuations.
- Neutral pions decay immediately to  $\gamma\gamma$  which deposit all their energy in localized electromagnetic showers.