

# Homework Set 4

Physics 631

J 9.22, 9.23

C. Greene's solutions

#9.22 (a, b)



We have a spherical resonant cavity inside a perfect conductor

From 9.122, the relevant solutions to Maxwell's equations are written in terms of electric (E) and magnetic (M) multipole coefficients as:

$$\vec{H} = \sum_{lm} \left( a_E^{lm} j_l(kr) \vec{X}_{lm} - \frac{i}{k} a_M^{lm} \nabla \times j_l(kr) \vec{X}_{lm} \right)$$

$$\vec{E} = Z_0 \sum_{lm} \left( \frac{i}{k} a_E^{lm} \nabla \times j_l(kr) \vec{X}_{lm} + a_M^{lm} j_l(kr) \vec{X}_{lm} \right)$$

And the boundary conditions at  $r=a$  are:

$$E_{||} = 0, \quad H_{\perp} = 0$$

Case (1) TE = M modes, i.e. set  $a_E^{lm} \rightarrow 0$

Now, since  $\vec{X}_{lm}$  has only angular components,

$$\Rightarrow \vec{E}_{lm} = Z_0 j_l(kr) \vec{X}_{lm}(\theta, \phi)$$

$$\vec{H}_{lm} = -\frac{i}{k} \nabla \times [j_l(kr) \vec{X}_{lm}(\theta, \phi)]$$

Therefore  $\vec{E}_{lm}$  will vanish at  $r=a$  if  $j_l(ka) = 0$

$$\Rightarrow k_{l,p} = \frac{\zeta_{l,p}}{a} \text{ where } \zeta_{l,p} = p^{\text{th}} \text{ zero of } j_l(s).$$

Note that since  $\vec{X}_{00} = \vec{Y}_{00} = 0$ ,  $l=0$  could only give the trivial solution, of no interest.

So the lowest 4 modes, obtained using Mathematica's command `BesselJZeros[L + 1/2, n]`, are

$l$	$p$	$\zeta_{l,p} = \frac{\omega_{l,p}}{c/a}$	$\lambda_{l,p}'' = \frac{2\pi a}{\zeta_{l,p}}$
1	1	4.49341	1.40a
2	1	5.76346	1.09a
3	1	6.98793	0.90a
1	2	7.72525	0.81a

Case (2) TM = E modes  $\Rightarrow a_M \rightarrow 0$

$$\text{and } \vec{H}_{lm} = j_l(kr) \vec{X}_{lm}$$

$$\vec{E}_{lm} = \frac{iZ_0}{k} \nabla \times [j_l(kr) \vec{X}_{lm}(\theta, \phi)]$$

Actually, let's first check whether the BC  $H_{\perp} = 0$  is satisfied in case (1):

$$\nabla \times (\psi \vec{v}) = \nabla \psi \times \vec{v} + \psi \nabla \times \vec{v}$$

$$\nabla j_{\ell} = \hat{r} \frac{\partial}{\partial r} j_{\ell}(kr)$$

$$\begin{aligned} \hat{r} \cdot \vec{H}_{\ell m} &= \hat{r} \cdot \left( -\frac{i}{k} \nabla \times [j_{\ell}(kr) \vec{X}_{\ell m}] \right) \\ &= \hat{r} \cdot \left( -\frac{i}{k} \left[ \nabla j_{\ell}(kr) \times \vec{X}_{\ell m} + j_{\ell}(kr) \nabla \times \vec{X}_{\ell m} \right] \right) \end{aligned}$$

0

$$\text{while } \nabla \times \vec{Y}_{\ell m} = -i \vec{r} \nabla^2 Y_{\ell m} + i \nabla \left( 1 + r \frac{\partial}{\partial r} \right) Y_{\ell m}$$

(using 9.125)

$$= -\frac{i \vec{r}}{r^2} [-\ell(\ell+1)] Y_{\ell m} + i \nabla Y_{\ell m}$$

and so finally,

$$\hat{r} \cdot \vec{H}_{\ell m} = \frac{i}{r} \ell(\ell+1) Y_{\ell m} j_{\ell}(kr)$$

so this means that the BC  $H_{\perp} = 0$

is satisfied at  $r = a$  after we have

satisfied the  $E_{\parallel} = 0$  BC for  $M$  modes

by choosing  $k = \frac{s_{\ell, p}}{a}$ !

Now, returning to the  $E$ -modes, they automatically obey  $H_{\perp} = 0 = \hat{r} \cdot \vec{H}_{\ell m}$  for any  $k$ , since  $\vec{H}_{\ell m}$  has only  $\hat{\theta}, \hat{\phi}$ -components.

Thus we must impose  $E_{\parallel} = 0$ , using some of the above formulas, i.e.

$$\nabla \times [j_l(kr) \vec{L} Y_{lm}] = \frac{\partial j_l(kr)}{\partial r} \hat{r} \times (\vec{L} Y_{lm}) + j_l(kr) \nabla \times \vec{L} Y_{lm}$$

$$= \frac{\partial j_l(kr)}{\partial r} \hat{r} \times (\vec{r} \times \vec{p}) Y_{lm} + j_l(kr) \left[ + \frac{i\vec{r}}{r^2} l(l+1) Y_{lm} + i \nabla Y_{lm} \right]$$

$$\hookrightarrow = \left[ \vec{r} (\hat{r} \cdot \vec{p}) - r \vec{p} \right] Y_{lm}$$

$\hat{r} \cdot \vec{p} = 0, \vec{p} = -i \nabla$

does not contribute to  $E_{||}$ , so ignore

||-components in  $\hat{\theta}, \hat{\phi}$

$$0 = i \left( r \frac{\partial j_l(kr)}{\partial r} \Big|_{r=a} + j_l(kr) \Big|_{r=a} \right) \nabla Y_{lm}$$

and therefore the  $E_{||} = 0$  BC for E modes will be satisfied if  $k$  obeys:

$$ka j_l'(ka) + j_l(ka) = 0, \text{ where } j_l'(z) \equiv \frac{d}{dz} j_l(z)$$

In Mathematica, find roots as follows:

$$jj[L_-, z_-] := \left(\frac{\pi}{2z}\right)^{1/2} \text{BesselJ}[L+1/2, z]$$

$$jjD[L_-, z_-] := D[jj[L, x], x] /. x \rightarrow z$$

$$\text{FindRoot}[z jjD[L, z] + jj[L, z] = 0, \{z, 3+L\}]$$

Call  $s'_{l,p} = p^{\text{th}}$  zero of  $s' j_l'(s') + j_l(s') = 0$   
 i.e.  $\frac{d}{ds'} [s' j_l(s')] = 0$

which gives the following for the lowest 4 or 5 roots!

$l$	$p$	$s'_{l,p} = \frac{\omega'_{l,p}}{c/a}$	$\lambda'_{l,p} / a$
1	1	2.744	2.29
2	1	3.870	1.62
3	1	4.973	1.26
4	1	6.062	1.04
1	2	6.117	1.03

(c) The lowest TM mode and the lowest TE mode both have  $l=1$ , and 3  $m$ -values  $m=0, \pm 1$ . Here I work out just the  $m=0$  cases, for simplicity:

$$\begin{aligned} \underline{TE_{p0m} = TE_{110}} \quad \vec{E}_{110}^{(TE)} &= z_0 j_1(kr) \vec{X}_{10}(\theta, \phi) \\ \vec{H}_{110}^{(TE)} &= -\frac{i}{k} \nabla \times [j_1(kr) \vec{X}_{10}(\theta, \phi)] \end{aligned}$$

I will solve this using the conventions of Edmonds and/or Varshalovic et al. <sup>p.217</sup> ~~7.3.6.55~~ or Edmonds 5.9.19, p.84:

$$\begin{aligned} \nabla \times (j_l \vec{X}_l) &= \nabla \times (j_l(kr) \vec{Y}_{lM}^l) \\ &= \left(\frac{l}{2l+1}\right)^{1/2} i \left( \frac{d}{dr} j_l(kr) - \frac{l}{r} j_l(kr) \right) \vec{Y}_{lM}^{l+1} + i \left( \frac{dj_l}{dr} + \frac{(l+1)}{r} j_l \right) \left(\frac{l+1}{2l+1}\right)^{1/2} \vec{Y}_{lM}^{l-1} \end{aligned}$$

and explicit expressions were obtained in Mathematica, Vector Spherical Harmonics.nb

e.g.  $\vec{Y}_{1,0}(\theta, \phi) = Y_{11} \hat{e}_{-1} \langle 11, 1-1 | 10 \rangle + Y_{1,-1} \hat{e}_1 \langle 1-1, 11 | 10 \rangle$   
 (note:  $\langle 10, 10 | 10 \rangle = 0$ )

$$\begin{aligned} &= \frac{1}{\sqrt{2}} \left( -\frac{1}{2} e^{i\phi} \sqrt{\frac{3}{2\pi}} \sin\theta \right) \left( \frac{1}{\sqrt{2}} (\hat{x} - i\hat{y}) \right) \\ &\quad - \frac{1}{\sqrt{2}} \left( \frac{1}{2} e^{-i\phi} \sqrt{\frac{3}{2\pi}} \sin\theta \right) \left( -\frac{1}{\sqrt{2}} (\hat{x} + i\hat{y}) \right) \\ &= (-\hat{x} \sin\phi + \hat{y} \cos\phi) \sqrt{\frac{3}{2\pi}} \frac{\sin\theta}{2} i = \vec{X}_{10}(\theta, \phi) \end{aligned}$$

The full expression for the  $\nabla \times (j_1(kr) \vec{X}_{10})$  is written out in full in the Mathematica notebook.

Then the same formulas also apply to the TM mode, with the different value of  $k$  substituted, and the following interchange:

$$\vec{H}_{110}^{TM} = \vec{E}_{110}^{TE} / Z_0$$

$$\vec{E}_{110}^{TM} = -Z_0 \vec{H}_{110}^{TE}$$

9.23 The cavity  $Q$  is given by:

$$Q \equiv \omega_0 \frac{\text{Stored Energy}}{\text{Rate of power loss}} = \frac{\omega_0 U}{\frac{dU}{dt}}$$

where the time-averaged energy inside for a TE mode is:

$$U = \int d^3x \left\{ \frac{\epsilon_0}{4} |\vec{E}|^2 + \frac{\mu_0}{4} |\vec{H}|^2 \right\}$$

$$= \int_0^a r^2 dr \int d\Omega \left\{ \frac{\epsilon_0}{4} \left| z_0 j_l \left( s_{l,p} \frac{r}{a} \right) \vec{X}_{lm} \right|^2 \right.$$

$$\left. + \frac{\mu_0}{4} \left| \frac{-i}{\frac{s_{l,p} z_0}{a}} \nabla \times \left( z_0 j_l \left( s_{l,p} \frac{r}{a} \right) \vec{X}_{lm} \right) \right|^2 \right\}$$

$$= \frac{\epsilon_0}{4} z_0^2 \int_0^a r^2 j_l^2 \left( s_{l,p} \frac{r}{a} \right) dr \rightarrow \frac{\epsilon_0 z_0^2}{4 a^{-3}} \int_0^1 x^2 j_l^2 \left( s_{l,p} x \right) dx$$

$$+ \frac{\mu_0}{4} \int_0^a r^2 dr \left\{ \left( \frac{d j_l \left( s_{l,p} \frac{r}{a} \right)}{dr} - l \frac{j_l \left( s_{l,p} \frac{r}{a} \right)}{r} \right)^2 \frac{l}{2l+1} \right.$$

$$\left. + \left( \frac{d j_l \left( s_{l,p} \frac{r}{a} \right)}{dr} + \frac{l+1}{r} j_l \left( s_{l,p} \frac{r}{a} \right) \right)^2 \frac{l+1}{2l+1} \right\}$$

↑  
(used last Eq. on p.05 above, + orthonormality)  
of the vector spherical harmonics

also recall that  $\epsilon_0 z_0^2 = \mu_0$ , or calling  $k = \frac{s_{l,p}}{a}$

$$\Rightarrow U = \frac{\mu_0}{4 k^3} \int_0^{s_{l,p}} x^2 \left\{ \left[ j_l'(x) - \frac{l j_l(x)}{x} \right]^2 \frac{l}{2l+1} + j_l^2(x) \right.$$

$$\left. + \left[ j_l'(x) + \frac{l+1}{x} j_l(x) \right]^2 \frac{l+1}{2l+1} \right\} dx$$

$$\Rightarrow U = \frac{\mu_0 a^3}{4 s_{l,p}^3} \left\{ s_{l,p}^3 [j_l'(s_{l,p})]^2 \right\} \quad (\text{see next pages})$$

To see how to evaluate this integral, return to the first equation for  $U^{TE}$  above on p.07, and set  $k = \omega \mu_0 / a$ :

$$U^{TE} = \int d^3x \left\{ \frac{\mu_0}{4} j_l^2(kr) |\vec{X}_{lm}|^2 + \frac{\mu_0}{4k^2} |\nabla \times j_l(kr) \vec{X}_{lm}|^2 \right\}$$

and through some vector spherical harmonic identities, e.g. Jackson 10.48, p. 472, this simplifies to:

$$U^{TE} = \frac{\mu_0}{4} \int_0^a r^2 dr \left\{ \underbrace{2 j_l^2(kr)}_{\text{integral}_1} + \frac{1}{k^2 r^2} \frac{\partial}{\partial r} \left[ r j_l(kr) \frac{\partial}{\partial r} (r j_l(kr)) \right] \right\}$$

First, notice that  $\text{integral}_2$  is easily done because the integrand is a perfect derivative, i.e.

$$\begin{aligned} \text{integral}_2 &= \frac{1}{k^2} r j_l(kr) \frac{\partial}{\partial r} (r j_l(kr)) \Big|_0^a \\ &= \frac{1}{k^3} \underbrace{(x j_l(x))}_{\substack{\text{vanishes for} \\ \text{TE modes}}} \underbrace{(x j_l(x))'}_{\substack{\text{vanishes for} \\ \text{TM modes}}} \Big|_{x=ka} = 0! \end{aligned}$$

lower limit  $\rightarrow 0$   
because  $r j_l(kr) \rightarrow 0$   
 $r \rightarrow 0$

$$\text{Thus } U^{TE} = 2 \frac{\mu_0}{4} \int_0^a dr (r j_l(kr))^2$$

To evaluate this, consider the slightly more general integral  $I(k, k') = \int_0^a dr [r j_l(kr)] [r j_l(k'r)]$

we will take the limit  $k \rightarrow k'$  later

Now recall the differential equation obeyed, namely

$$\hat{L}(r j_l(kr)) = k^2(r j_l(kr))$$

where  $\hat{L} = -\frac{d^2}{dr^2} + \frac{l(l+1)}{r^2}$ , and hence:

$$I(k, k') = \frac{\int_0^a dr \{ r j_l(kr) \hat{L}[r j_l(k'r)] - [\hat{L} r j_l(kr)] r j_l(k'r) \}}{k'^2 - k^2}$$

$$= (k'^2 - k^2)^{-1} \int_0^a dr \left\{ \frac{d^2}{dr^2} (r j_l(kr)) r j_l(k'r) - r j_l(kr) \frac{d^2}{dr^2} (r j_l(k'r)) \right\}$$

$$= (k'^2 - k^2)^{-1} \int_0^a \frac{d}{dr} W(r j_l(k'r), r j_l(kr)) dr$$

$$= (k'^2 - k^2)^{-1} W[r j_l(k'r), r j_l(kr)]_{r=a}$$

(the  $r=0$  contribution vanishes at all  $k, k'$ )

(recall, the Wronskian definition is  $W(y, z) = yz' - y'z$ )

→ to evaluate this  $\frac{0}{0}$  limit as  $k' \rightarrow k$ ,

set  $k' = k + \Delta$  and expand the numerator and denominator to 1<sup>st</sup> order in  $\Delta$ ,

which is akin to de l'Hospital's rule.

Using  $j_l[(k+\Delta)r] = j_l(kr) + r\Delta j_l'(kr) + O(\Delta^2)$ ,

$$\Rightarrow I(k, k) = \lim_{\Delta \rightarrow 0} \left( \frac{1}{2k\Delta} \right) \left\{ r j_l[(k+\Delta)r] \frac{d}{dr} (r j_l(kr)) - \frac{d}{dr} [r j_l((k+\Delta)r)] r j_l(kr) \right\}_{r=a}$$

or simplifying,

$$I(k, k) = \frac{1}{2k} \left\{ a j_l'(ka) \frac{d}{dx} (x j_l(kx)) - a j_l(ka) \frac{d}{dx} (x j_l'(x)) \right\}_{x=ka}$$

⇒ Through this device, all integrals needed for this problem can be evaluated analytically.

e.g. For TE modes,

$$U^{TE} = \frac{\mu_0}{2} I(k, k) \Big|_{k = s_{l,p}/a} \quad \text{where } j_l(s_{l,p}) = 0$$

$$= \frac{\mu_0}{4} a^3 \left( \frac{d}{dx} [j_l(x)] \right)_{x \rightarrow s_{l,p}}^2$$

Now, to complete this problem we need to evaluate the rate of power loss,

$$P_{\text{loss}} = \frac{\mu_0 \omega^2}{4} \int_S da |\hat{n} \times \vec{H}|^2$$

where for TE modes,  $\vec{H}_{plm}^{TE} = \frac{-i}{k} \nabla \times j_l(kr) \vec{X}_{lm}$   
and  $k = s_{l,p}/a$

and

$$\begin{aligned} \nabla \times [j_l \vec{X}_{lm}] &= i \left( \frac{d j_l}{dr} - \frac{l j_l}{r} \right) \left( \frac{l}{2l+1} \right)^{1/2} \vec{Y}_{lm}^{l+1} \\ &+ i \left( \frac{d j_l}{dr} j_l + j_l \frac{l+1}{r} \right) \left( \frac{l+1}{2l+1} \right)^{1/2} \vec{Y}_{lm}^{l-1} \end{aligned}$$

and using Varshalovich 7.3.73 on p 220:

$$\hat{n} \times \vec{Y}_{lm}^{l+1} = i \left( \frac{l}{2l+1} \right)^{1/2} \vec{X}_{lm}$$

$$\hat{n} \times \vec{Y}_{lm}^{l-1} = i \left( \frac{l+1}{2l+1} \right)^{1/2} \vec{X}_{lm}, \quad \text{giving}$$

$$P_{\text{loss}}^{\text{TE}} = \frac{\mu_0 \omega \delta}{4} a^2 \frac{1}{k^2} \left\{ \frac{-l}{2l+1} \left( \frac{dj_l}{dr} - \frac{l j_l}{r} \right) - \frac{(l+1)}{2l+1} \left( \frac{dj_l}{dr} + \frac{l+1}{r} j_l \right) \right\}^2$$

$$= \frac{\mu_0 \omega \delta}{4} a^2 [j_l'(s_{l,p})]^2$$

so

$$Q^{\text{TE}} = \frac{\omega U^{\text{TE}}}{P_{\text{loss}}^{\text{TE}}} = \frac{a}{\delta}$$

for all  $l, m!$   
amazing simplification!

Next work out  $Q^{\text{TM}}$ , and we see at

once that  $U_E^{\text{TM}} = \left( \frac{Z_0^2 \epsilon_0}{\mu_0} \right) U_B^{\text{TE}}, U_B^{\text{TM}} = \left( \frac{\mu_0}{Z_0^2 \epsilon_0} \right) U_E^{\text{TE}}$

except  $s_{l,p} \rightarrow s'_{l,p}$

$$\text{and } P_{\text{loss}}^{\text{TM}} = \frac{\mu_0 \omega \delta}{4} a^2 [j_l(s'_{l,p})]^2$$

$$\text{and } U^{\text{TM}} = \frac{\mu_0 a^3}{4 s'_{l,p}{}^3} \left\{ s'_{l,p}{}^3 [j_l(s'_{l,p})]^2 \left( 1 - \frac{l(l+1)}{s'_{l,p}{}^2} \right) \right\}$$

$$\Rightarrow Q_l^{\text{TM}} = \frac{\omega U^{\text{TM}}}{P_{\text{loss}}^{\text{TM}}} = \left( 1 - \frac{l(l+1)}{s'_{l,p}{}^2} \right) \frac{a}{\delta}$$

$\Rightarrow$  again, amazingly simple in the end!