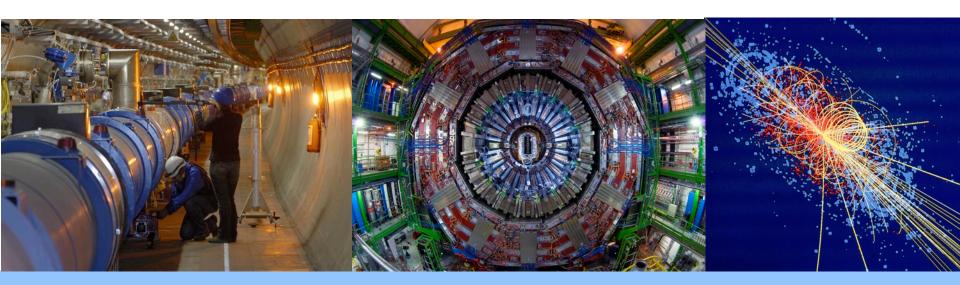
The Large Hadron Collider The Discovery Machine

Norbert Neumeister

Department of Physics Purdue University



Purdue SURF Research Seminar, June 2009

Outline

- Introduction
 - What is Particle Physics
 - The Standard Model
 - Why do we go to the energy frontier
- CERN
- The Large Hadron Collider
 - The machine and the physics
- The Compact Muon Solenoid detector
 - Construction
 - Preparation for data taking
- Search for Physics Beyond the Standard Model
- Summary

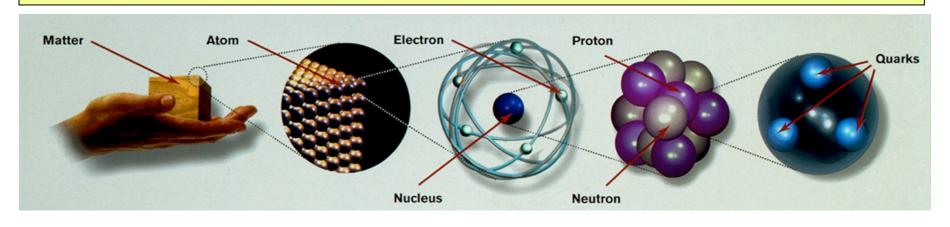
Particle Physics

Aim to answer the two following questions:

- What are the elementary constituents of matter?
- ➤ What are the fundamental forces that control their behavior at the most basic level?

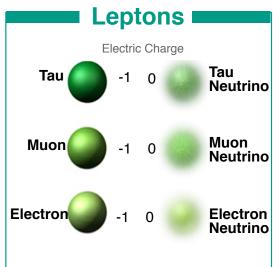
Tools:

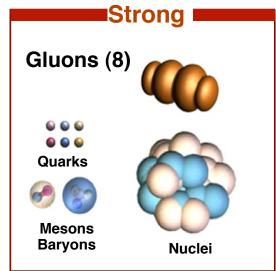
Particle Accelerators, Particle Detectors, Computers

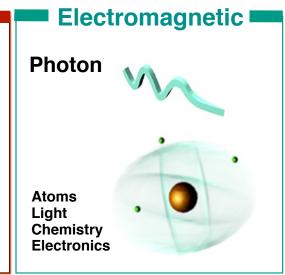


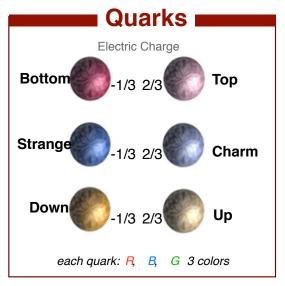
atom 10⁻¹⁰ m nucleus 10⁻¹⁴ m nucleon 10⁻¹⁵ m quark 10⁻¹⁸ m

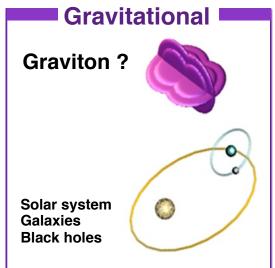
Particles and Forces

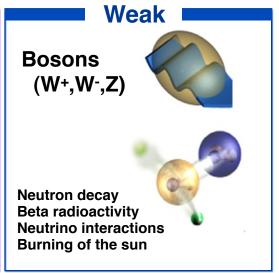




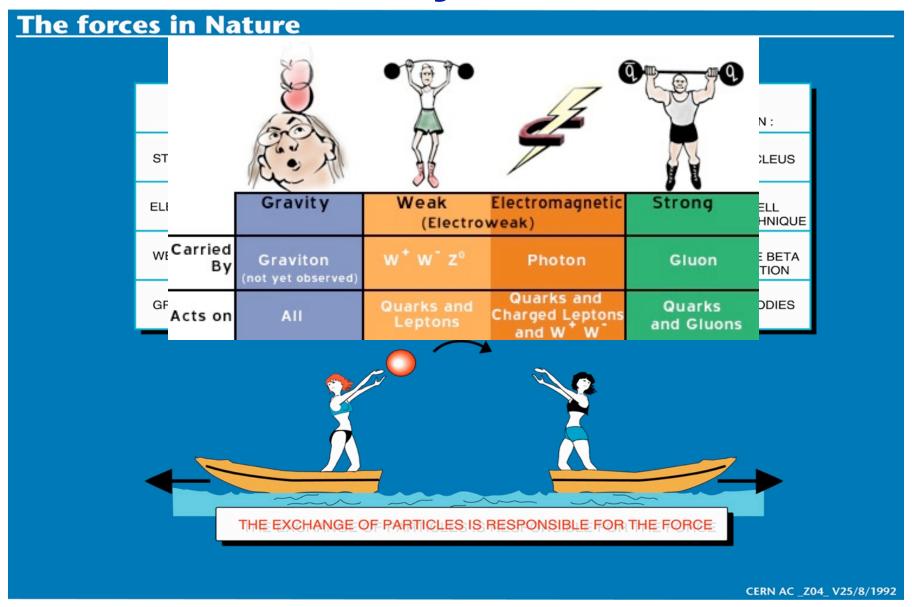




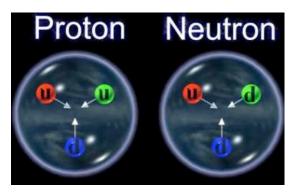




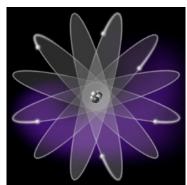
Hierarchy of Fields



Hierarchy of Structure



 $R \sim 10^{-15} \text{ m (strong)}$



 $R \sim 10^{-10} \text{ m}$ (electromagnetic)

 $R \sim 1/m_e$

The mass of the electron determines the size of atoms

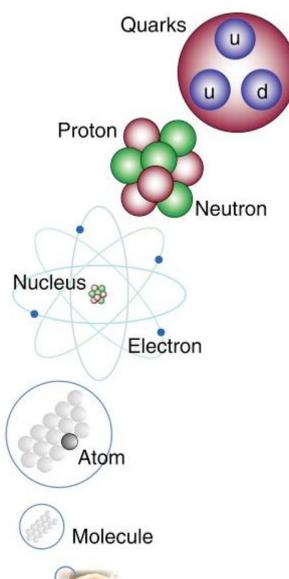


 $R > 10^6 \, m$ (gravitational)

The Standard Model

- Very successful model which contains all known particles in particle physics today.
- Describes the interaction between spin 1/2 particles (quarks and leptons) mediated by spin 1 gauge bosons (gauge symmetry).
- Electroweak physics tested up to per mill level.
- BUT: We know the Standard Model is incomplete:
 - One important element still missing: The Higgs boson (spin 0)
 - We know everything about the Higgs except for its mass
 - It is the last piece of the SM and also the key to understanding any beyond-the-SM physics
 - So we need to find the Higgs; and theory does NOT provide (precise) information on its mass!

The Standard Model



matter particles

	1st gen.	2nd gen.	3rd gen.
Q	(II)	(2)	
U	up	charm	top
A R	(A)		CA
K	down	strange	bottom
L E	(Ve)	(V)	
Р	e neutrino	μ neutrino	τ neutrino
T 0	P		7
N	electron	muon	tau

gauge particles



scalar particle(s)



Elements of the Standard Model

The Periodic Table

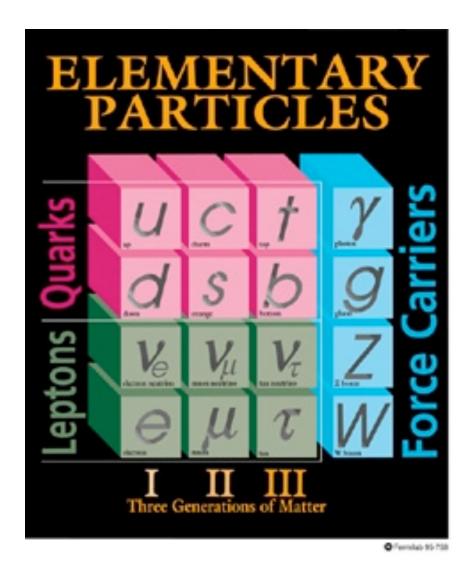
_ 1	2		3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18
hydrogen 1	l																	helium 2
н	l																	He
1.0079	0.00000		- 1	Key:	Y.La. Commission		4						CHEMICAL PROPERTY	. 200000.000	90 J. 9530C A.D.	A 17 10 475 A	0.000.000.000	4.0026
Ithium 3	beryllium.	atomic number											boron 5	carbon 6	ntrogen 7	oxygen 8	fluorine 9	10
ı i	D.	symbol											В	C	N	Ô		Ne
	Be																F	
6.941 sodium	9.0122 magnesium	atomic weight (mean relative mass)											10.811 guminium	12.011 sãoon	14.007 phosphorus	15.999 suffur	18.998 chlorine	20.180 argon
11	12											13	14	15	16	17	18	
Na	Mg													Si	Р	S	CI	Ar
22.990	24.305	5		****				-				-	26.982	28.086	30.974	32.086	35.453	39.948
potassium 19	calcium 20		21	thanium 22	vanadum 23	24	manganese 25	26	27	28	copper 29	30	gallium 31	germanium 32	amenic 33	selenium 34	35	36
K	Ca		Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr
59.098	40.078		44.956	47.887	50.942	51.996	54.938	55.845	58.933	58.693	63.546	65.30	69.723	72.61	74.922	78.96	79.904	83.60
nubidium 37	atrontium 38		yttrium 39	ziroonium 40	niobium 41	molybdenum 42	technetium 43	ruthenium 44	rhodum 45	paladum 46	47	cadmium 48	49	50	antimony 51	tellurium 52	53	54
Rb	Sr		v	Zr	Nb	Mo	Tc	Ru	Rh	Pd	_	Cd		Sn	Sb	Te	ï	Xe
											Ag		In					
85.468 cessium	87.62 berium	-	88.906 Metium	91.224 hatnium	92.906 tentalum	95.94 Sungsten	[97.907] rhenium	101.07 camium	102.91 indust	108.42 platinum	107.87 gold	112.41 mercury	114.82 fhallium	118.71 lead	121.76 bismuth	127.80 polonium	128.90 astatre	131.29 redon
55	56	57-70	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86
Cs	Ba	*	Lu	Hf	Ta	w	Re	Os	- Ir	Pt	Au	Hg	TI	Pb	Bi	Po	At	Rn
US	Da												200	444.4	200 00	200.000.000		[222.02]
132.91	137.33		174.97	178.49	180.95	183.84	188.21	190.23	192.22	195.08	198.97	200.59	204.38	201.2	208.98	[208.96]	[209.90]	
7	700	89-102		178.49 rutherfordium	dubnium	183.84 seaborgium 106	188.21 bohrlum 107	190.23 hassium 108	192.22 metherium 109	195.08 ununnilium 110	198.97 unununium 111	200.59 ununbium 112	204.36	urunquadum 114	208.98	116	[209.90]	ununcetium 118
132.91 francium 87	137.33 mdum 88	89-102 **	174.97 lawrencium 103	178.49 rutherfordium 104	dubnium 105	seatorgium 106	107	hassium 108	metherium 109	ununnitium 110	111	112	204.36	114		ununhexium 116	ord narradical	ununcetium 118
132.91 francium	137.33 radium		174.97 lawrencium	178.49 rutherfordium	dubnium	seaborgium	bohrlum	hassium	meitnerlunt	ununnitium 110	unununium	112	204.36			ununhexium	ord narradical	ununcetium

*lanthanoids

**actinoids

57	58	praecdymium 59	neodymium 60	61	62	europium 63	gadolinium 64	65	dysprosium 66	67	erbium 68	thulium 69	70
La	Ce	Pr	Na	Pm	Sm	Eu	Ga	Tb	Dy	Но	Er	Tm	T D
138.91	140.12	140.91	144.24	[144.91]	150.36	151.96	157.25	158.93	162.50	164.93	167.26	168.93	173.04
actinium 89	90	protectinium 91	92	neptunium 93	plutonium 94	americium 95	96	97	osifomium 98	eimaninium 99	100	101	nobelium 102
Ac	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No
[227.03]	232.04	231.04	238.03	[237.05]	[244.06]	[243.06]	[247.07]	[247.07]	[251.06]	[262.06]	[257.10]	[258.10]	[259.10]

The "Newer" Periodic Table



What is Matter?

Particles in various combinations

Quarks up charm top down strange bottom

Leptons





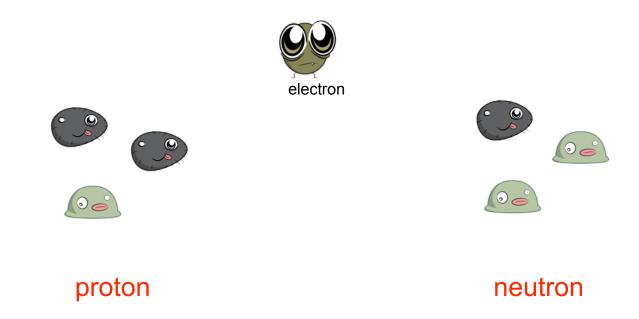






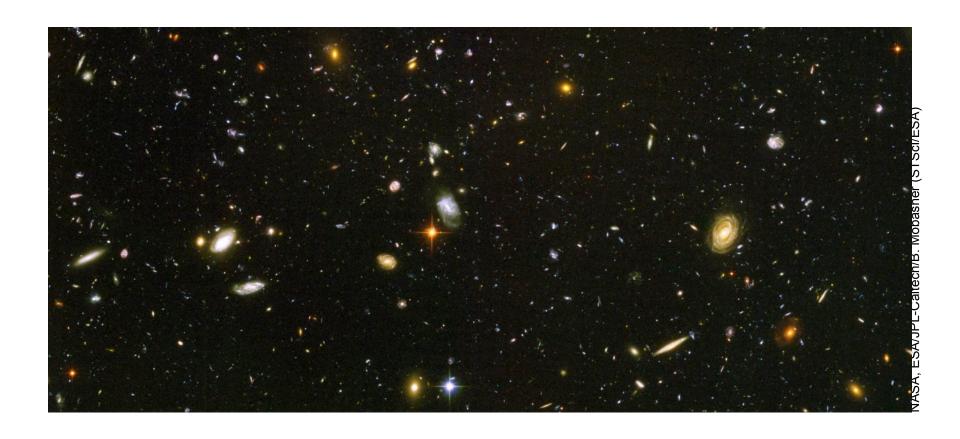


Building a Universe



Multiply by billions and billions and billions and billions...

Building a Universe



The Universe

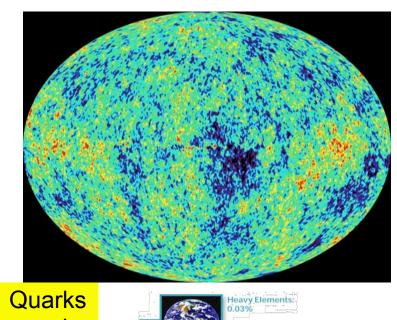
What shapes the cosmos?

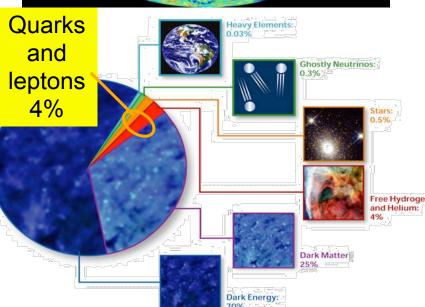
Old answer: Gravity – through mass and energy observed in the universe

- But now we know that:
 - ❖ There is much more mass than we expect from just the stars we see, or from the amount of helium that was formed in the early universe... <u>Solution</u>: **Dark matter**
 - ❖ The velocity of distant galaxies indicates that, in addition to the mass that slows down expansion of the universe, there is also some new kind of energy driving that expansion outward!

Solution: Dark Energy

- So, if we're so smart, why don't we know what 96% of the universe is made of?
 - Back to the accelerators!





Probing the TeV Energy Scale

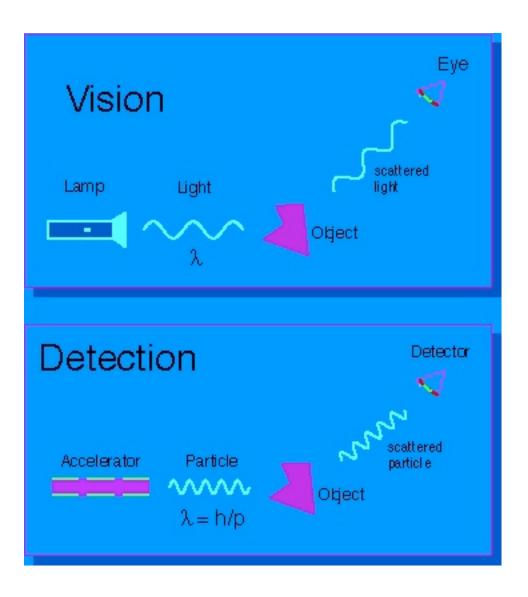
- Higher energy: Reproduce conditions of early Universe
- TeV energy scale: Expect breakdown of current calculations unless a new interaction or phenomenon appears
- Many theories, but need data to distinguish between them

What might we find:

- The mechanism that generates mass for all elementary particles
 - In Standard Model, masses generated through interaction with a new particle the Higgs boson
 - Other options possible, but we know that the phenomena occurs somewhere between 100 and 1000 GeV
- A new Symmetry of Nature
 - Supersymmetry gives each particle a partner
 - Would provide one source of the Dark Matter observed in the Universe
- Extra Space-Time Dimensions
 - String theory inspired
 - This would revolutionize Physics!
- These are only some of the possibilities

Why Accelerate Particles (1)

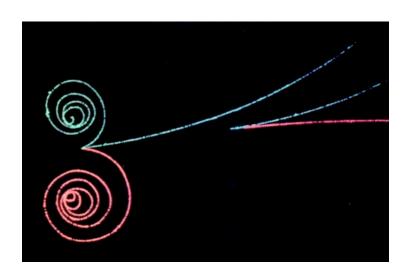
To look inside



Why Accelerate Particles (2)

- E = mc²: "Mass is condensed energy"
- Concentrate energy on one particle

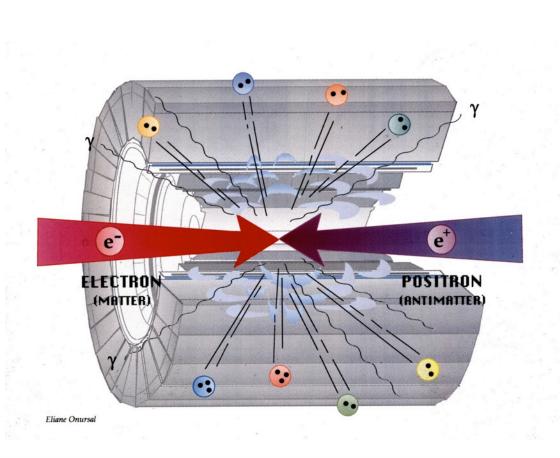
Photon -----

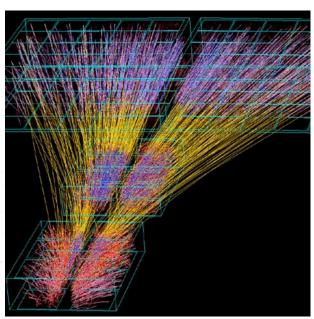


Creation of new particles

Particles from Energy

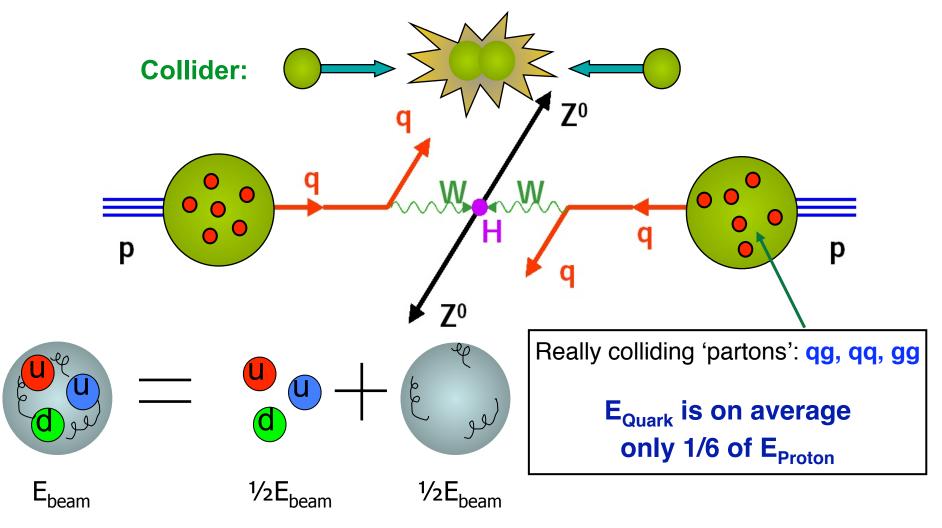
More energy - more (and new) particles are created





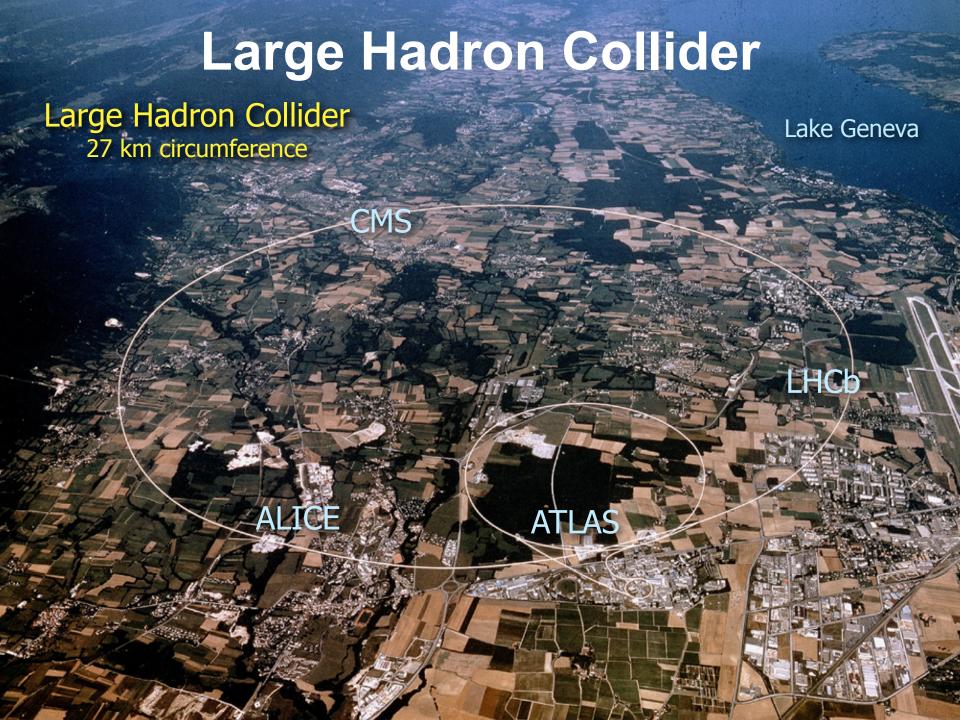
Particle collisions

Higgs Production in pp Collisions



 $E_{CM} \sim 1/3 E_{beam}$

→ Proton-proton collider with E_p ≥ 7 TeV



Conseil Européen pour la Recherche Nucléaire

European Laboratory for Particle Physics Founded in 1954

- 2415 staff*
- 730 Fellows and Associates*
- 9133 users*
- Budget (2007) \$900M (610M Euro)

*5 February 2008



- Member States: Austria, Belgium, Bulgaria, the Czech Republic, Denmark, Finland, France, Germany, Greece, Hungary, Italy, Netherlands, Norway, Poland, Portugal, Slovakia, Spain, Sweden, Switzerland and the United Kingdom.
 - Observers to Council: India, Israel, Japan, the Russian Federation, the United States of America, Turkey, the European Commission and Unesco

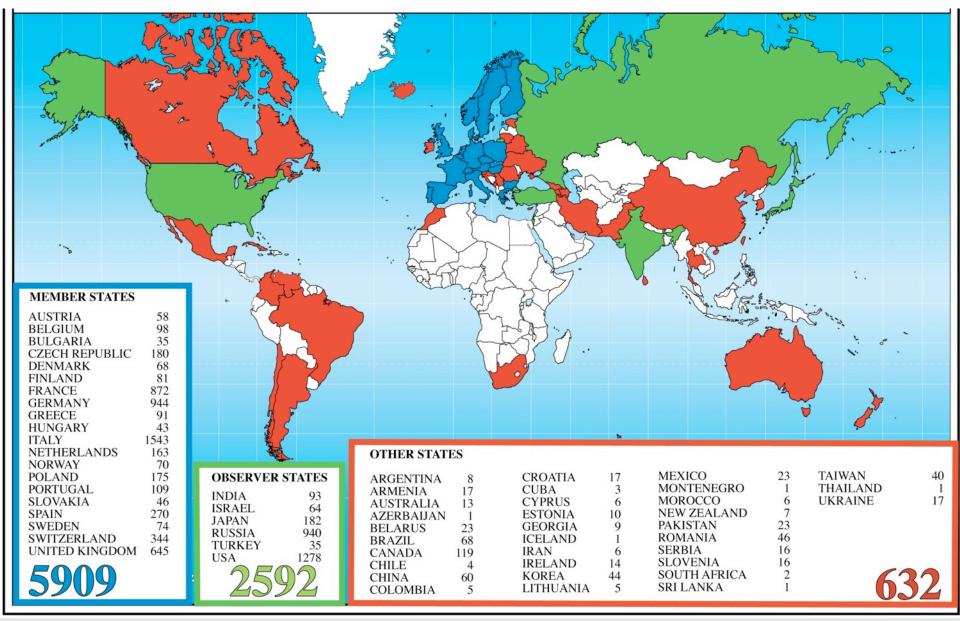
CERN



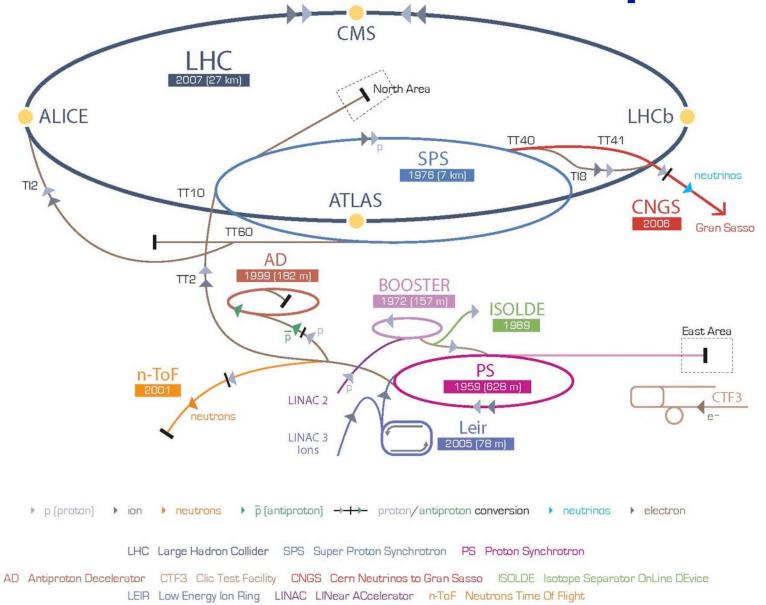
Near Geneva, Switzerland



CERN in Numbers



CERN Accelerator Complex



The World Wide Web



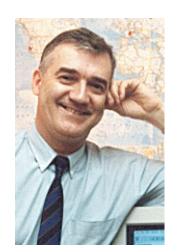
1990: Tim Berners-Lee, a CERN computer scientist invented the World Wide Web

The "Web" as it is affectionately called, was originally conceived and developed for the large high-energy physics collaborations which have a demand for instantaneous information sharing between physicists working in different universities and institutes all over the world. Now it has millions of academic and commercial users.

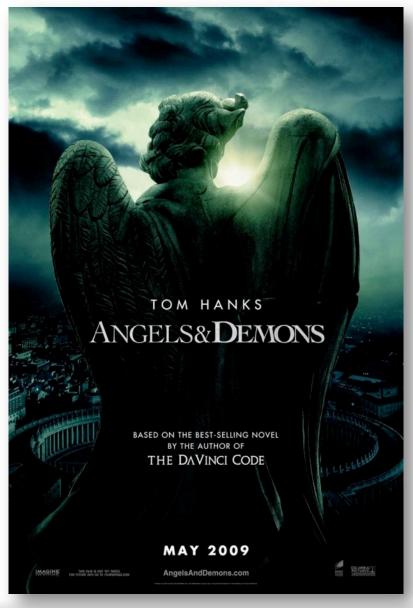


Tim together with Robert Cailliau, another CERN computer scientist, wrote the first WWW client (a browser-editor running under NeXTStep) and the first WWW server along with most of the communications software, defining URLs, HTTP and HTML.

In December 1993 WWW Tim received the IMA award and in 1995 Tim and Robert shared the Association for Computing (ACM) Software System Award for developing the World-Wide Web.



CERN and Hollywood



Angels and Demons

The plot:

- Antimatter is stolen from CERN's Large Hadron Collider (LHC) and hidden in Vatican City.
- Countdown to Vatican annihilation begins.
- Race through Rome to avert death and destruction.





The real LHC

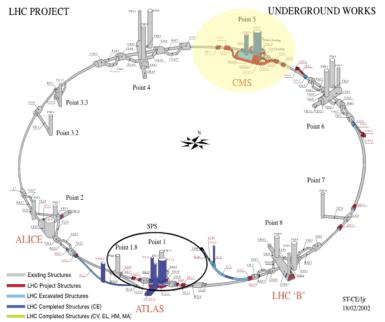
- The world's most powerful particle accelerator
- 16.8 miles around, 330 feet underground



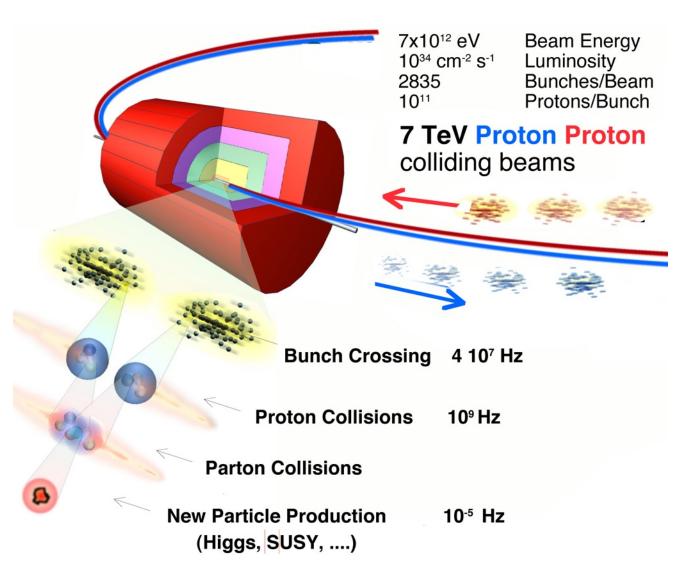
The Large Hadron Collider

- 7 TeV on 7 TeV proton-proton collider, 27 km ring
 - 7 times higher energy than the Tevatron at Fermilab
 - 100 times higher design luminosity than Tevatron (L=10³⁴ cm⁻²s⁻¹)
- 1232 superconducting 8.4 T dipole magnets @ T=1.9° K
 - Largest cryogenic structure, with 40 ktons of mass to cool
- 4 experiments
 - ATLAS
 - CMS
 - LHCb
 - ALICE





Collisions at the LHC



Selection of 1 event in 10,000,000,000,000

Stored Energy

$$E = 2E_{\text{proton}}N_{\text{bunches}}N_{\text{protons}}$$

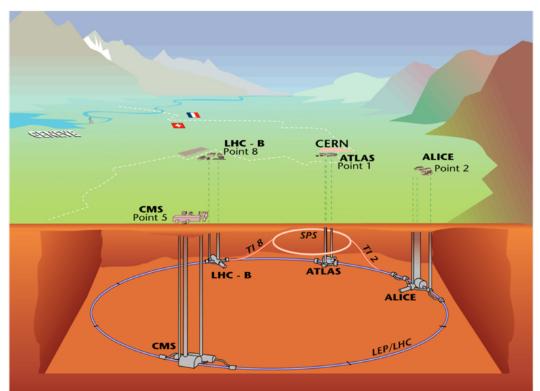
$$E = 2(7 \times 10^{12} \text{eV})(1.6 \times 10^{-19} \text{J/eV})(2808)(1.1 \times 10^{11}) \sim 700 \text{ MJ}$$

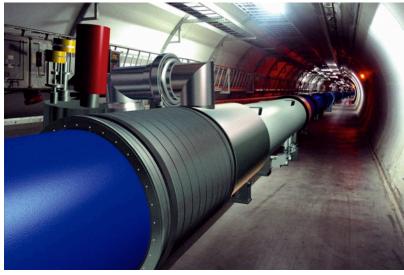
350 MJoule per beam: unprecedented!

- Kinetic energy
 - 1 small aircraft carrier of 10⁴ tonnes going 15 knots
 - 450 automobiles of 2 tonnes going 100 kph
- Chemical energy
 - 80 kg of TNT
 - 70 kg of (swiss?) chocolate
- Thermal energy
 - melt 500 kg of copper
 - raise 1 cubic meter of water 85 C



LHC tunnel, magnets

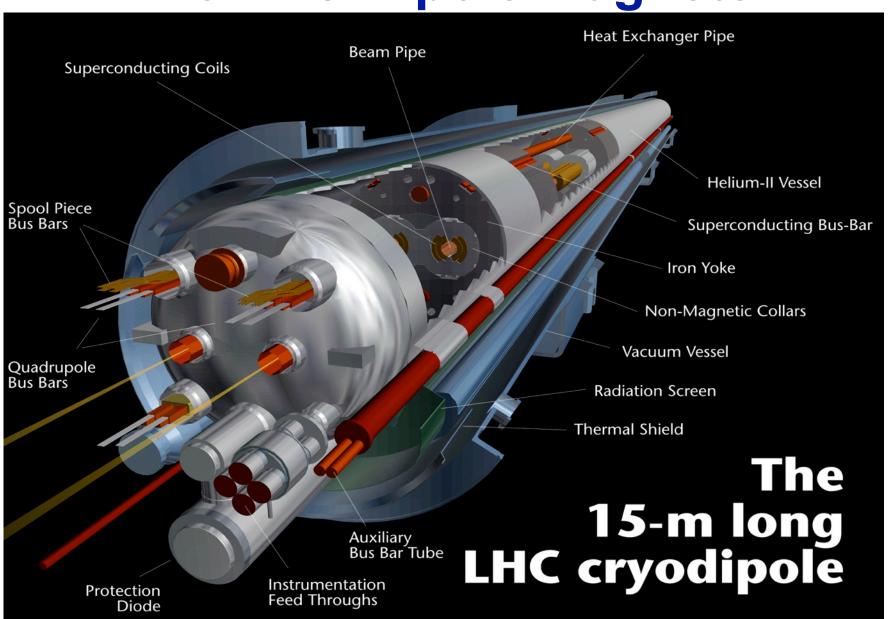




Proton-proton collisions E = 7000 + 7000 GeV 800 million collisions/sec Largest cryogenic system (1.8 K, suprafluid helium)

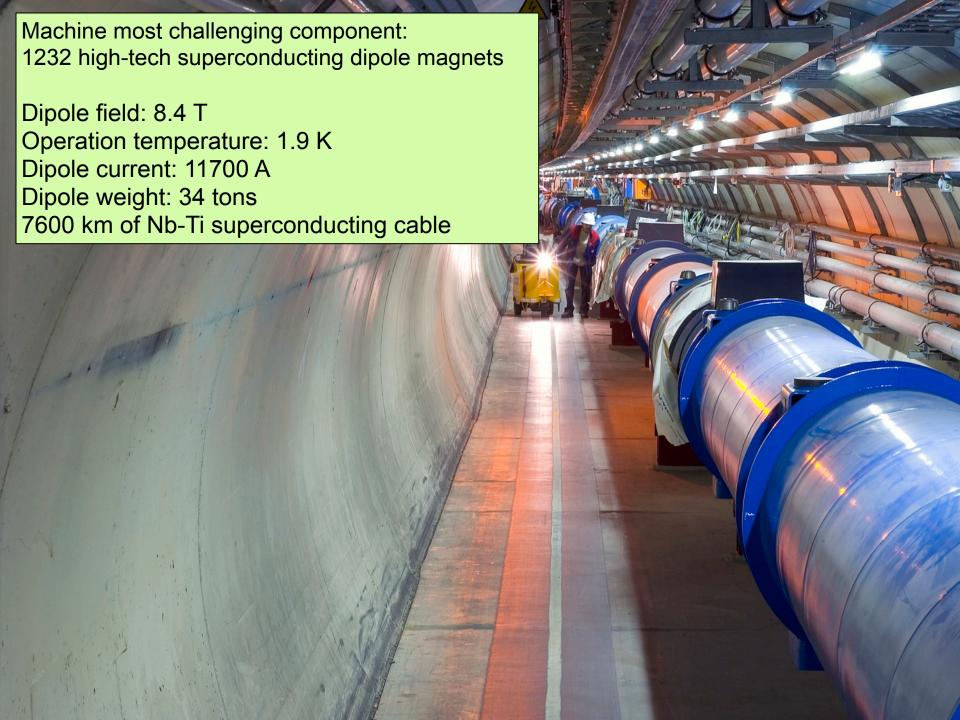
27 km of 8 T magnets 100 m below surface

The LHC Dipole Magnets

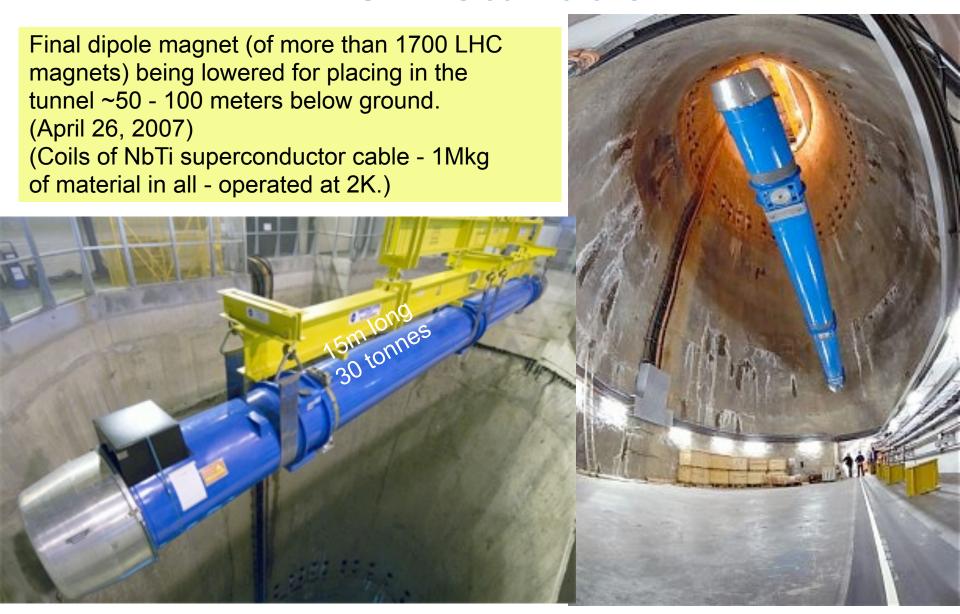


The LHC Dipole Magnets





LHC Installation



LHC History

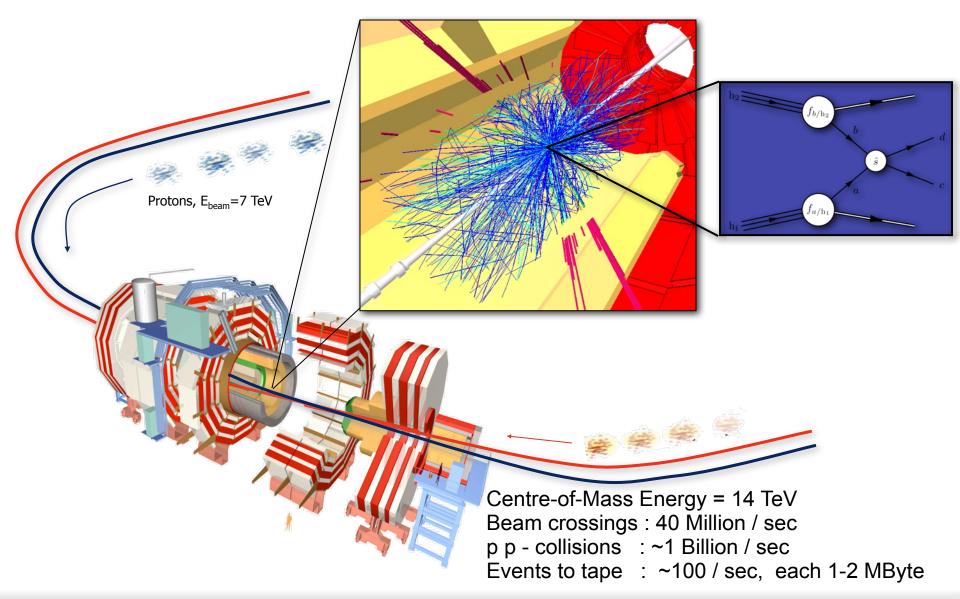
The most ambitious project in high-energy physics ever, and one of the most ambitious in science more generally

1984 Workshop on a Large Hadron Collider in the LEP tunnel, Lausanne 1987 Rubbia "Long-Range Planning Committee" recommends Large Hadron Collider as the right choice for CERN's future 1990 ECFA LHC Workshop, Aachen 1992 General Meeting on LHC Physics and Detectors, Evian les Bains 1993 Letters of Intent (ATLAS and CMS selected by LHCC) 1994 Technical Proposals approved Approval to move to construction (ceiling of 475 MCHF) 1996 1998 Memorandum of Understanding for construction signed 1998 Construction begins (after approval of Technical Design Reports) 2000 CMS assembly begins above ground. LEP closes 2004 CMS underground caverns completed 2008 CMS ready for first proton-proton Collisions

2008

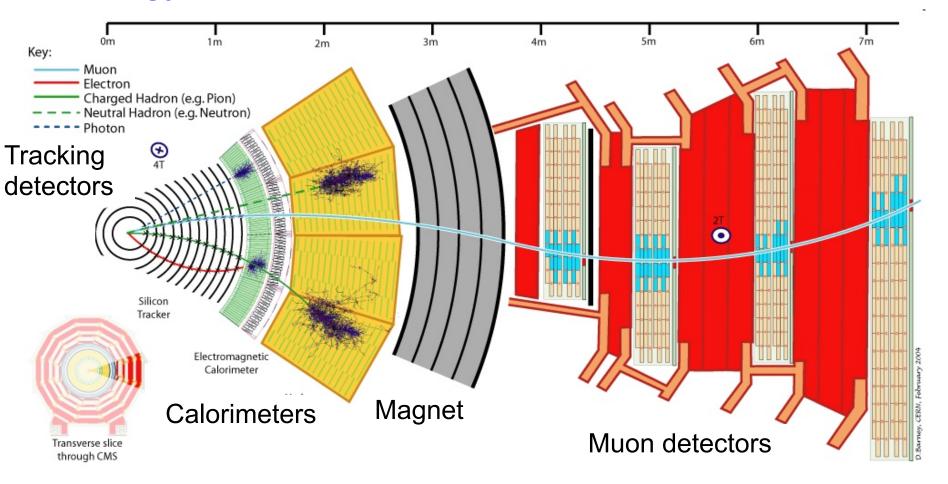
Official LHC inauguration: 21 Oct 2008

Collisions at the LHC



Particle Detection

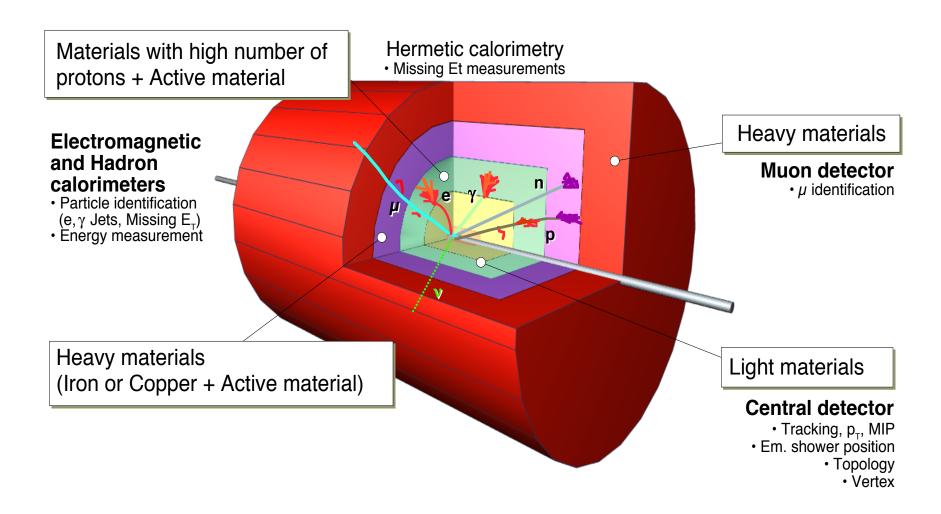
 Surround the collision point with instruments in order to identify the types of particles and their energy or momentum to reconstruct collisions



Detector Design

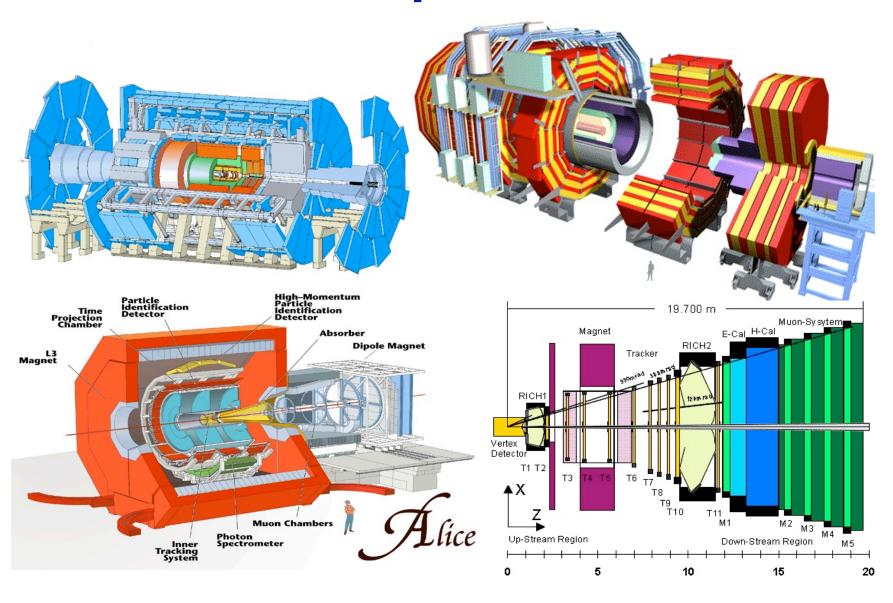
- LHC detectors must have fast response:
 - otherwise will integrate over many bunch crossings
 - typical response time: 20-40 ns
 - integrate over 1-2 bunch crossings → pile-up of 25-50 minimum bias events
 - very challenging readout electronics
- LHC detectors must be highly granular:
 - minimize probability that pile-up particles be in the same detectors element as interesting object
 - large number of readout channels ≈ O(107)
 - high cost
- LHC detectors must be radiation resistant:
 - high flux of particles from pp collisions → high radiation environment
 - up to 1017 n/cm2 in 10 years of LHC operation
 - up to 10 Gy (1 Gy = unit of absorbed energy = 1 Joule/kg)
- Can store data at ≈100 Hz:

Detectors at the LHC

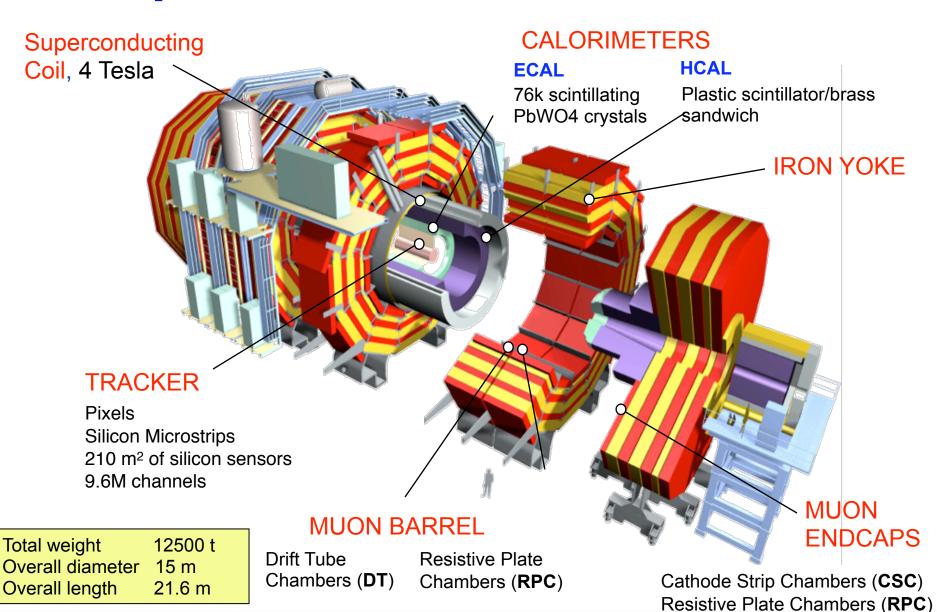


Each layer identifies and enables the measurement of the momentum or energy of the particles produced in a collision

LHC Experiments



Compact Muon Solenoid Detector



CMS Site Under Construction



Experiment Cavern

2003 2004

55 Cavern - Point 4 Headwall - 17-03-2003 - CERN ST-CE

CMS

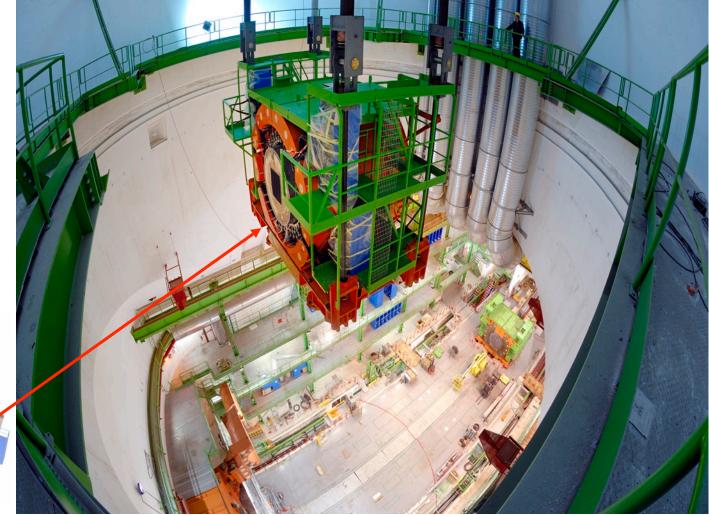


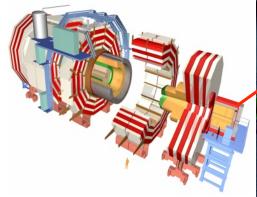
CMS Assembled & Tested on Surface



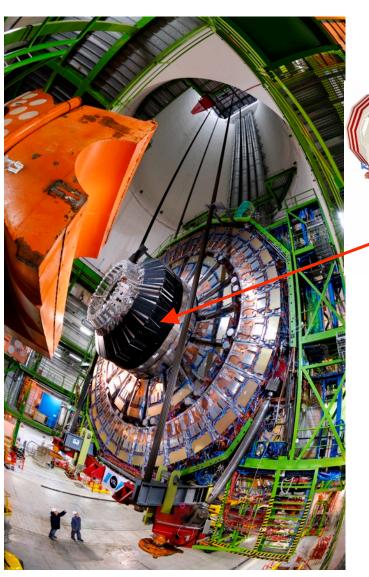
Lowering the Detector

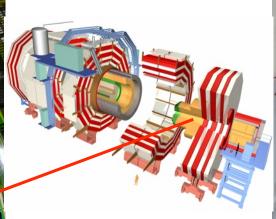
HF Nov 2006

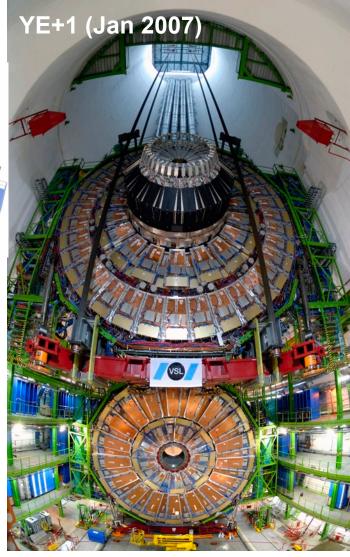




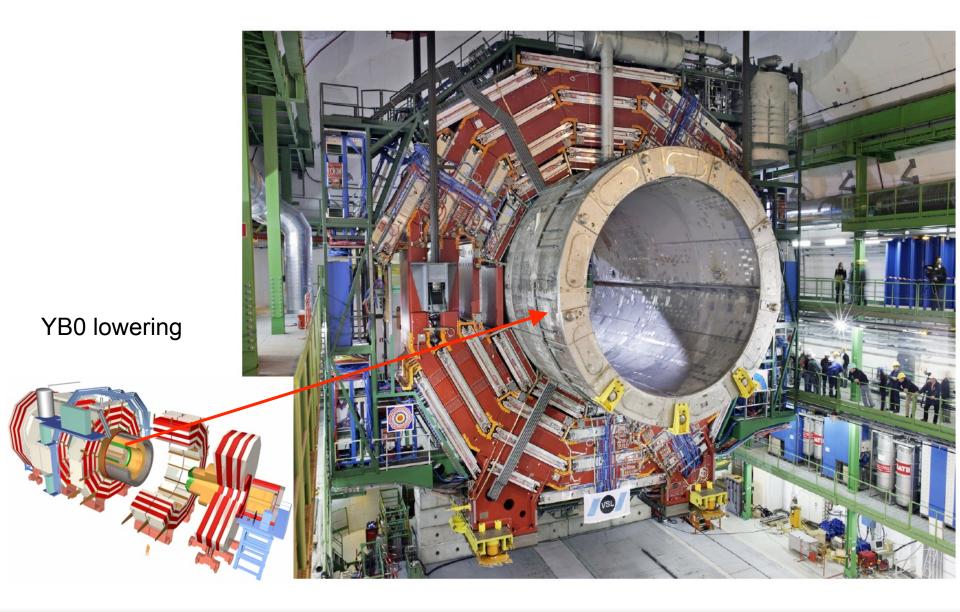
Lowering the Detector





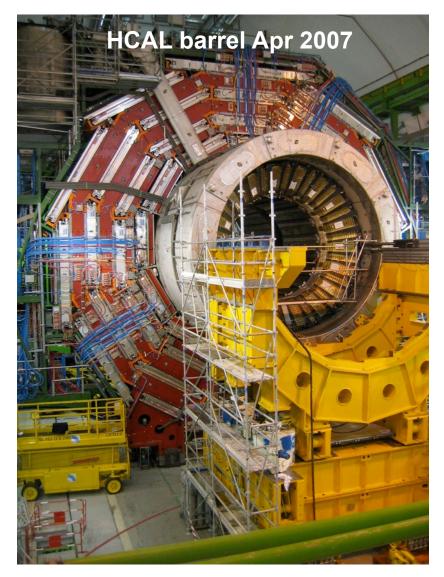


Lowering the Detector

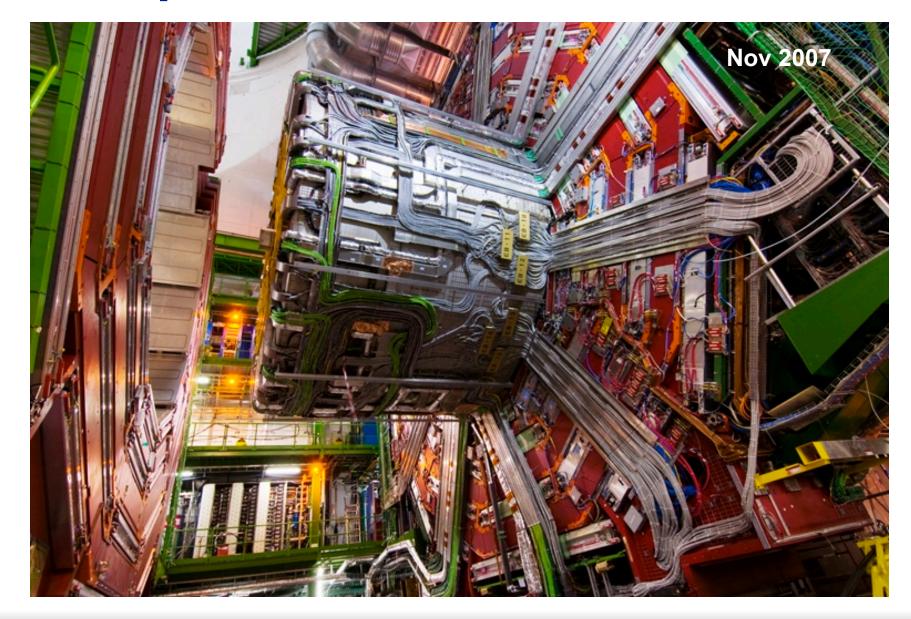


Putting the Detector in Place

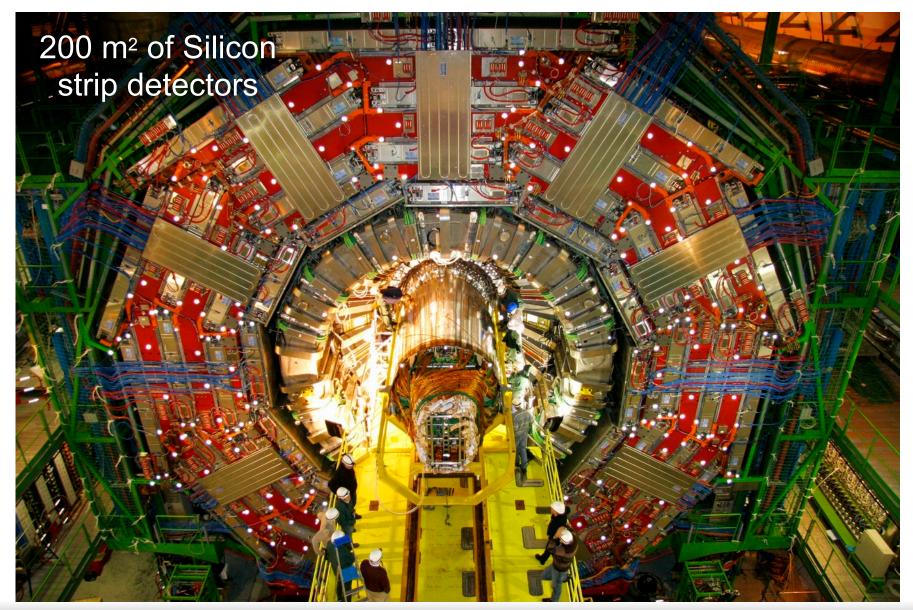




Completion of Services on YB0



Silicon Strip Tracker, Dec 2007



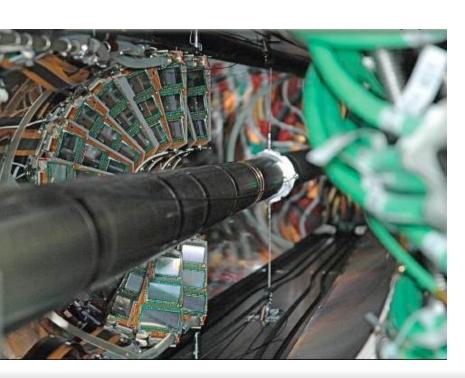
Beam-pipe Installed, May 2008



Installation of the Pixel System, Aug 2008

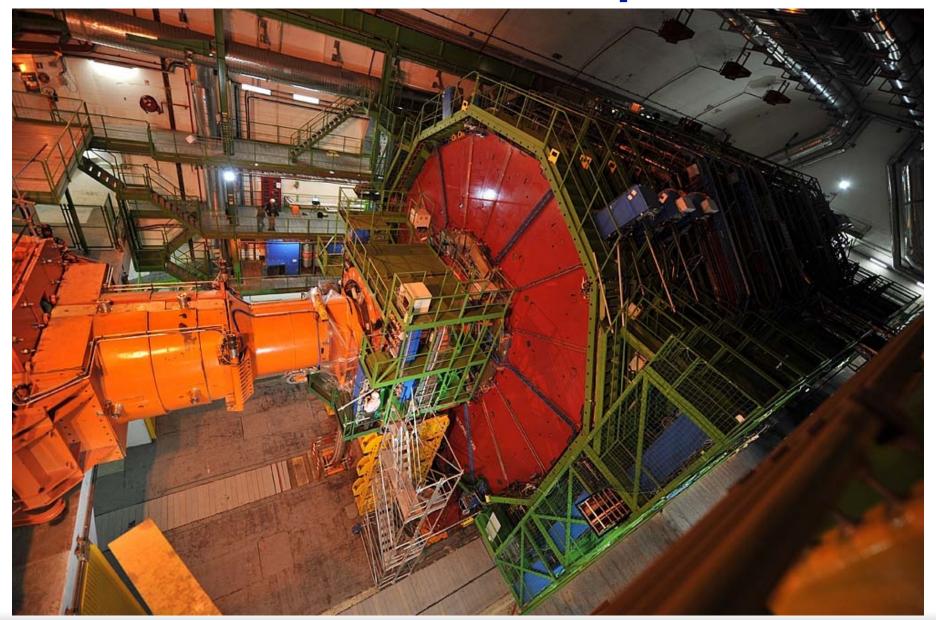
A 66 megapixel "camera"

Makes precise measurements of charged particle impact parameters to tag particles with a small but finite lifetime





Final Closure – Sep 2008



Cosmic Ray Muons



Cosmic ray muons used as probe of detector performance during no beam!

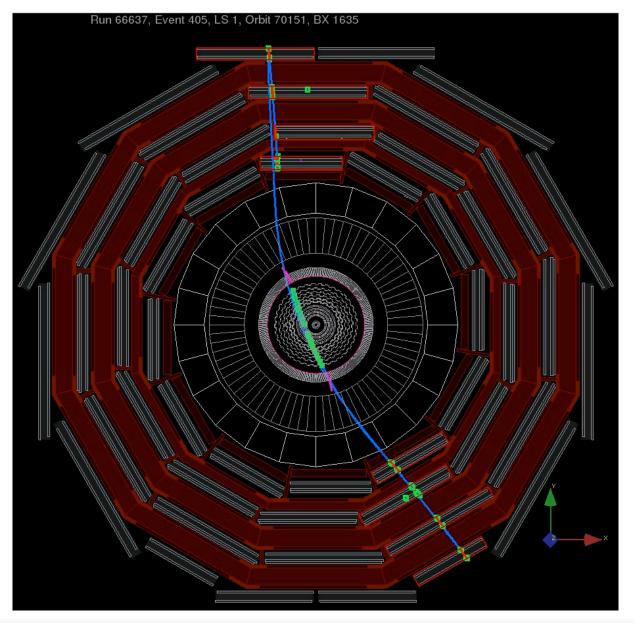
Total rate is about 350 Hz at 100 m depth (about 1% of rate on surface of Earth)

Cosmic rays are continuously bombarding Earth's atmosphere with far more energy than protons will have at the LHC, so cosmic rays would produce everything LHC can produce.

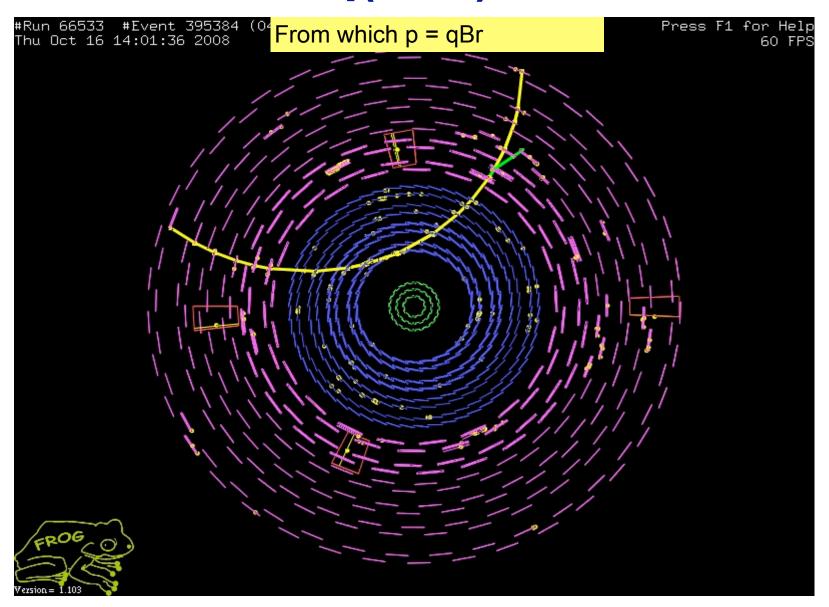
They have done so throughout the 4.5 billion years of the Earth's existence, and the Earth is still here!

The LHC just lets us see these processes in the lab (though at a much lower energy than some cosmic rays).

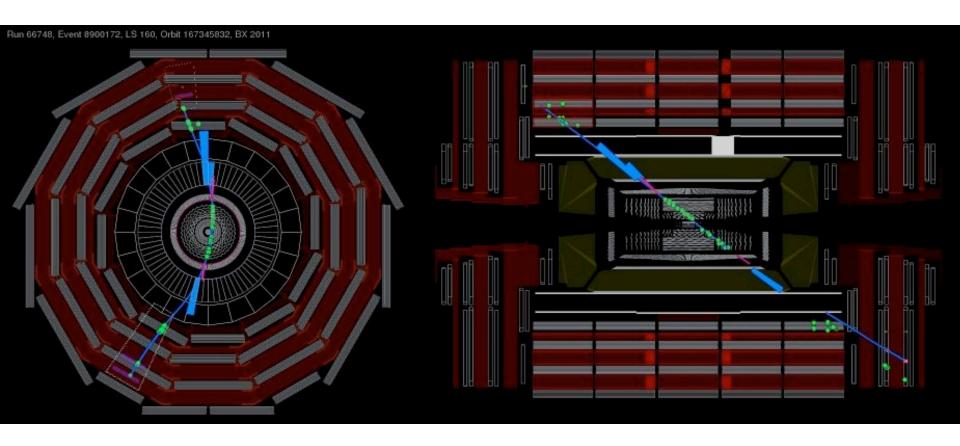
Cosmic Muons



$F=q(v\times B)$



Cosmic Muons



LHC Start-up: Sep. 10, 2008



Sep.10 – Big Media Day

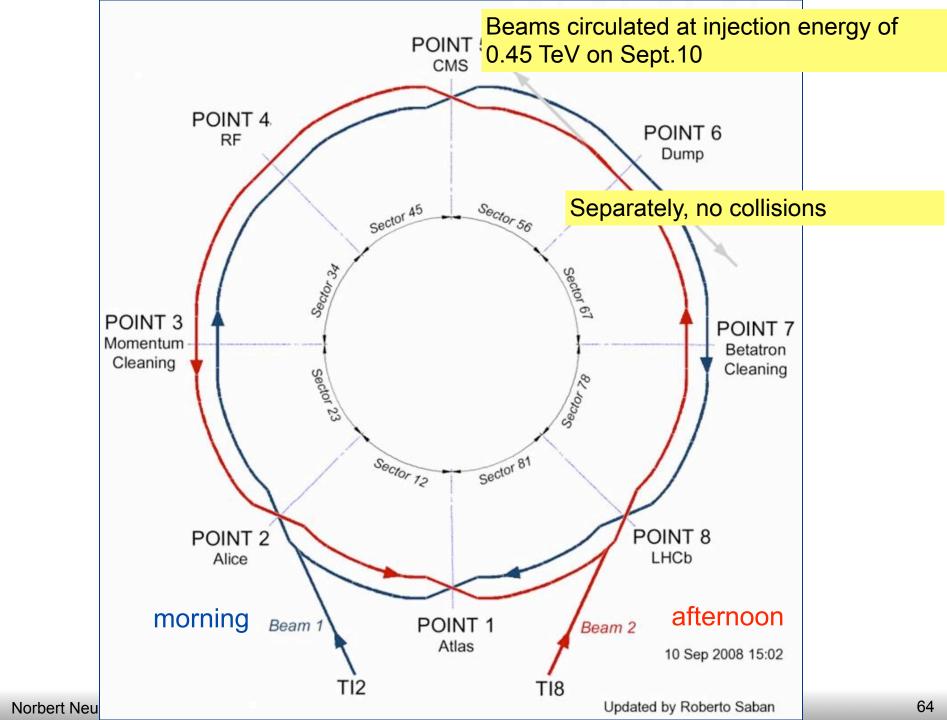






A Brief Diary of LHC Events

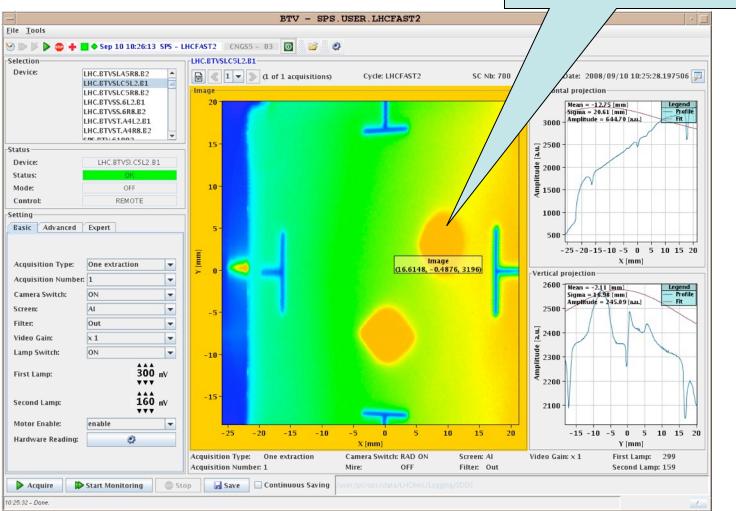
- Aug. 8 and Aug. 22 weekends
 - "Synchronization tests" sent protons through the first arcs of the LHC in both directions, past ALICE and then LHCb experiments
 - By Aug.22, alternate injections of beam 1 and 2: "...pretty blooming amazing..."
- Sun/Mon, 7-8 Sep.
 - Single shots of beam 1 onto a collimator 150m upstream of CMS.
- Tues, 9. Sep.
 - Additional single shots of beam 1 onto a collimator at CMS
- Wed., 10 Sep. (Media Day!)
 - Beam 1 circulated in the morning, 3 turns by 10:40am (1 hour)
 - Beam 2 circulated by 3:00pm
 - 300 turns of beam 2 by 11:15pm
- Thurs., 11 Sep.
 - RF system captures beam at 10:30pm (millions of orbits)



Sep. 10 Orbits

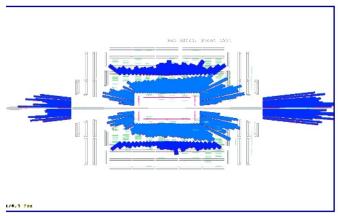
~2x109 protons

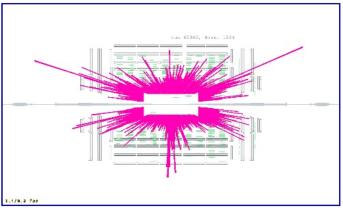
• 2 turns of clockwise beam:

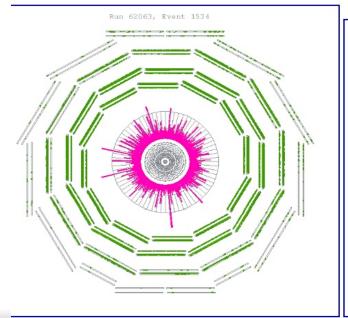


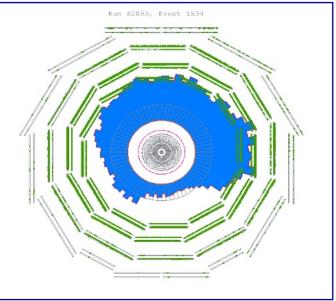
First Events: Collimators Closed

~2×10⁹ protons on collimator ~150 m upstream of CMS ECAL- pink; HB,HE - light blue; HO,HF - dark blue; Muon DT - green; Tracker Off



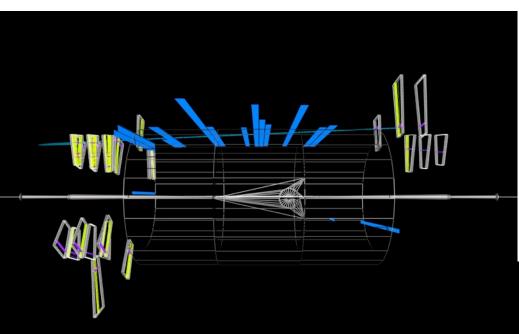


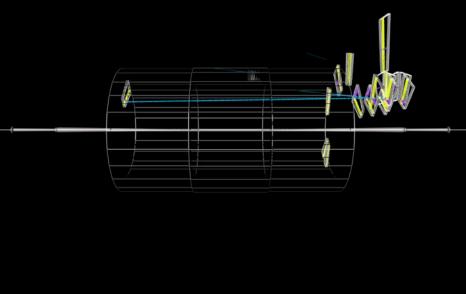




SURF Research Seminar 2009

Beam Halo Events





The Excitement of the First LHC Beam Measurements at CMS (September 2008)



Setback: Friday, Sep. 19

- An incident occurred during a powering test of one LHC sector for commissioning beam operation to 5 TeV
 - Massive helium loss in one arc of the tunnel (2 tons initially), cryogenics and vacuum lost
- The cause of the incident was determined to be a faulty electrical connection ("bus bar") between a dipole and a quadrupole
 - Mechanical damage occurred
 - Need to extract and repair dipole and quadrupole magnets in the region

Not enough time to make repairs before winter shutdown

→ Aim to restart LHC operations
Fall 2009



Physics Selection

Interactions/s:

- Lum = 10^{34} cm⁻²s⁻¹ = 10^7 mb⁻¹ Hz
- $\sigma_{\text{inel}}(pp) = 70 \text{ mb}$
- Interaction Rate, $R = 7 \times 10^8 \text{ Hz}$

Events/ beam crossing:

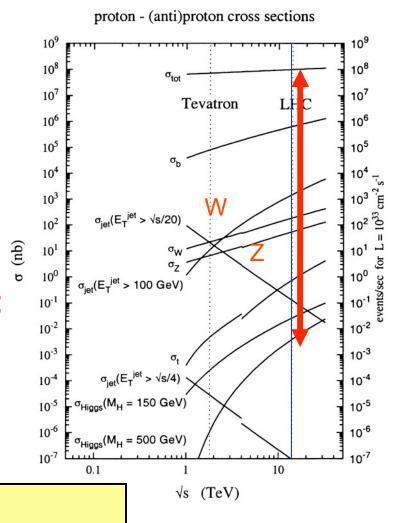
- $\Delta t = 25 \text{ ns} = 2.5 \times 10^{-8} \text{ s}$
- Interactions/crossing = 17.5

Cross-section of physics processes:

- inelastic: 109 Hz
- Higgs (600 GeV) : 10⁻² Hz

Selection needed: 1:10¹⁰⁻¹¹

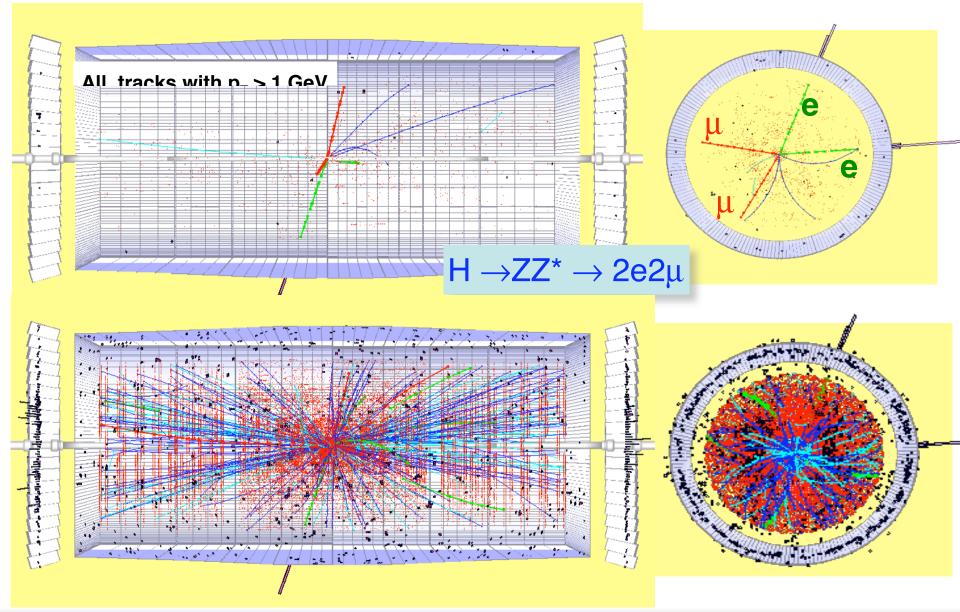
before branching fractions



Operating conditions:

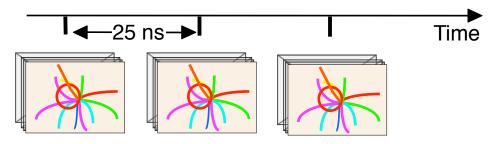
- 1) A "good" event containing a Higgs decay +
- 2) ~20 extra "bad" (minimum bias) interactions

pp Collisions at 14 TeV at 10³⁴ cm⁻²s⁻¹



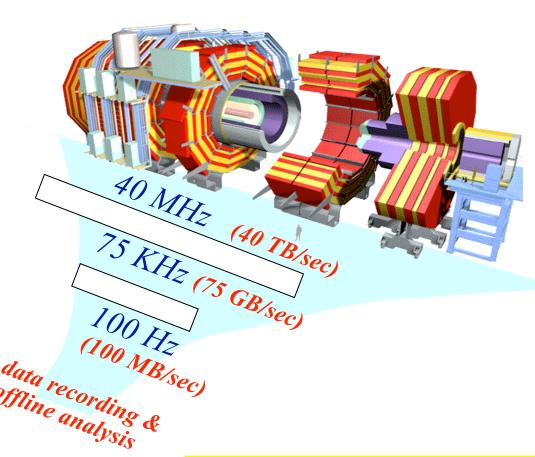
Event Selection

~20 collisions/25ns (109 event/sec) 107 channels (1016 bit/sec)



- The challenge is the identification of the most interesting (and potentially entirely new) physics processes amidst the much more copious occurrence of well-understood and studied processes.
- On-line event selection
 - Costume electronics (FPGAs, ASICs)
 - Online computing farm
 - A single processor analyzes one event at a time
- Out of a billion interactions/sec select one hundred for further analysis
 - Need to reject most interactions
 - Can store data at ≈100 Hz (= 100 events/s)

Data Recording



- Collision rate: 40 MHz
- Event size: ≈ 1 MByte
- Data size: 1 MByte/event
 100 events/s → 100 MByte/s
- •10⁷ s data taking per year (30%)
- Data size: 1 PetaByte =
 10³ TByte = 10⁶ GByte per year

∼ PetaBytes/year
 ∼10⁹ events/year
 ∼10³ batch and interactive users

LHC Data Challenge

Balloon (30 Km)

- The LHC generates 40×10⁶ collisions / s
- Combined the 4 experiments record:
 - 100 interesting collision per second
 - $-1 \div 12 \text{ MB} / \text{collision} \Rightarrow 0.1 \div 1.2 \text{ GB} / \text{s}$
 - $\sim 10 \text{ PB} (10^{16} \text{ B}) \text{ per year} (10^{10} \text{ collisions / y})$
 - LHC data correspond to 20×10⁶ DVD's / year!
 - Space equivalent to 400,000 large PC disks
 - Computing power ~ 10⁵ of today's PC

Using parallelism is the only way to analyze this amount of data in a reasonable amount of time

LHC data: DVD stack after 1 year! (~ 20 Km)

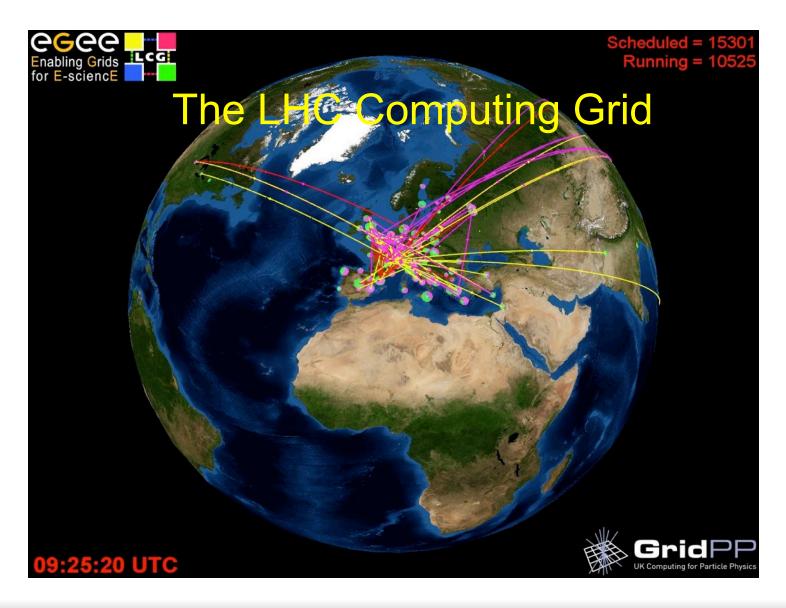




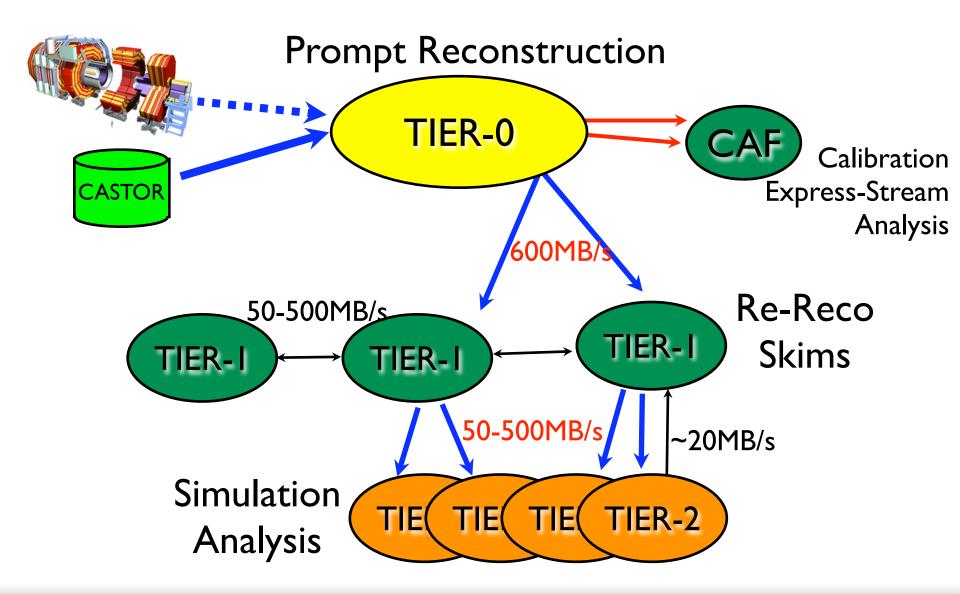
Mt. Blanc (4.8 Km)



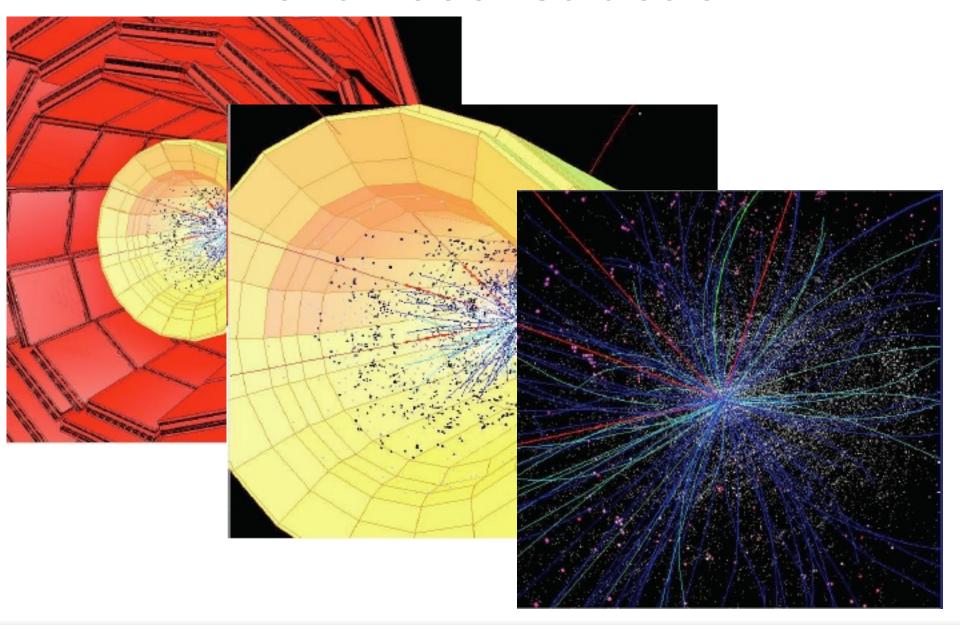
Computing and Software



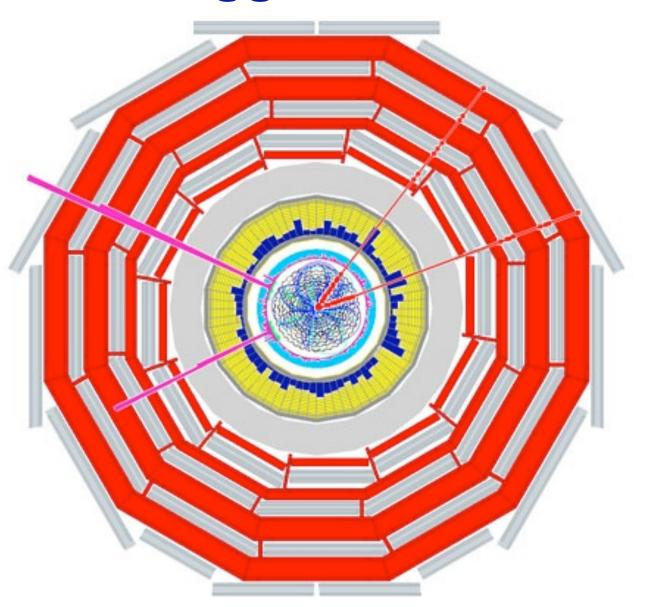
CMS Data Flow



Event Reconstruction



A Higgs Boson



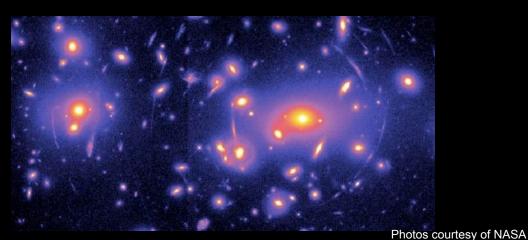
Dark Matter

In galaxies and galaxy clusters

There is not enough visible mass in rotating spiral galaxies to hold them together









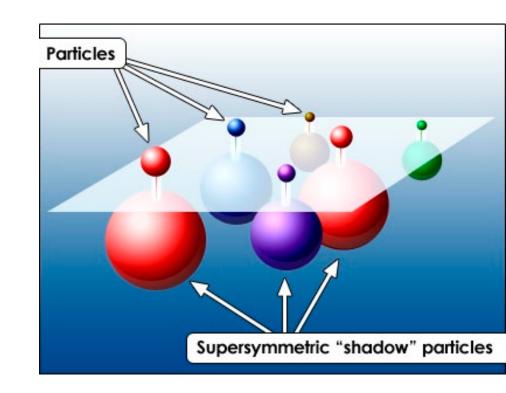
Supersymmetry

For fundamental particles, Supersymmetry says:

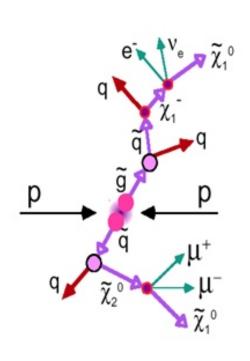
Every matter particle (fermion) should be associated with a massive "shadow" force carrier particle (boson).

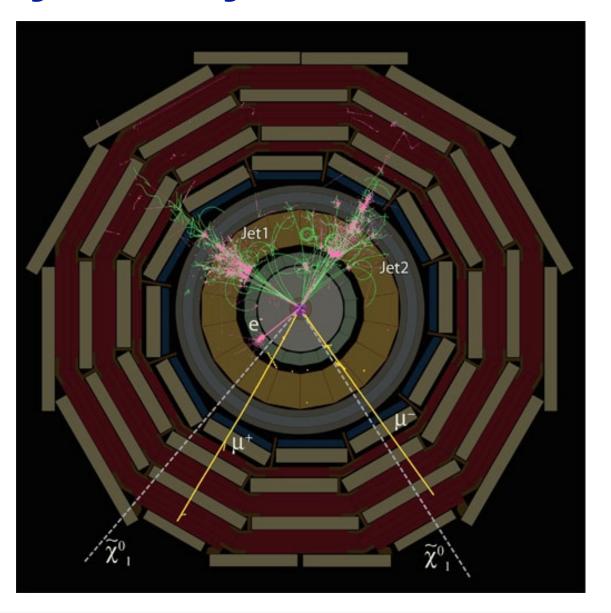
Every force carrier particle should have a massive "shadow" matter particle.

This has possible implications for <u>Dark Matter</u>



Supersymmetry in CMS





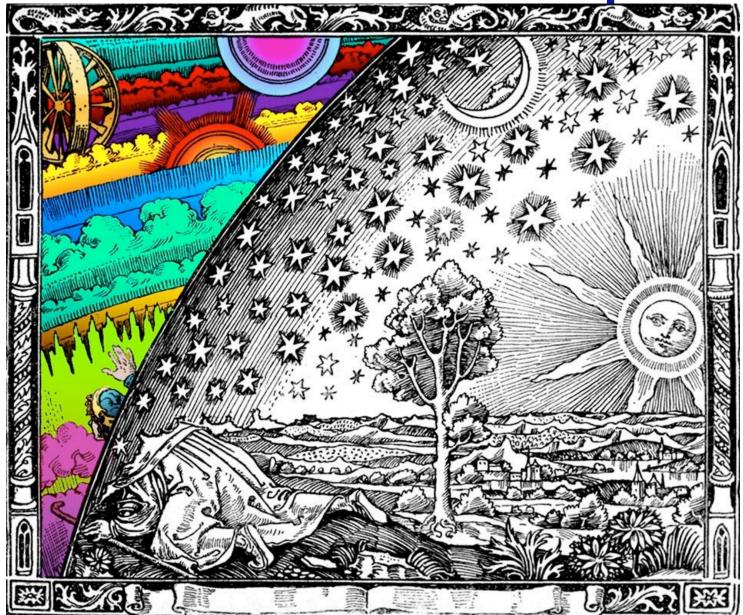
Is Matter also in Other Dimensions?

Are there extra dimensions of space that we cannot see?



(Dali, The Disintegration of the Persistence of Memory, 1954)

Extra Dimensions of Space



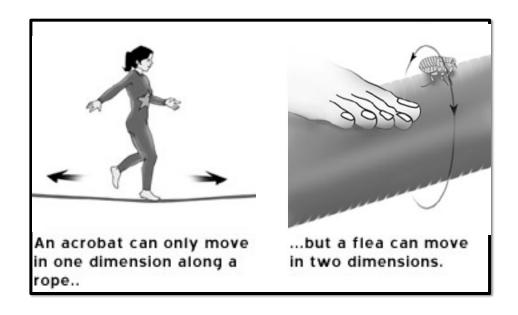
Extra Dimensions of Space

Think about an acrobat and a bug on a tight rope.

The acrobat can move forward and backward along the rope.

But the bug can also move sideways around the rope.

If the flea keeps walking to one side, it goes around the rope and winds up where it started.

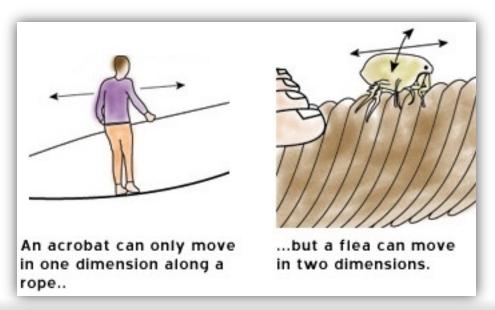


Extra Dimensions of Space

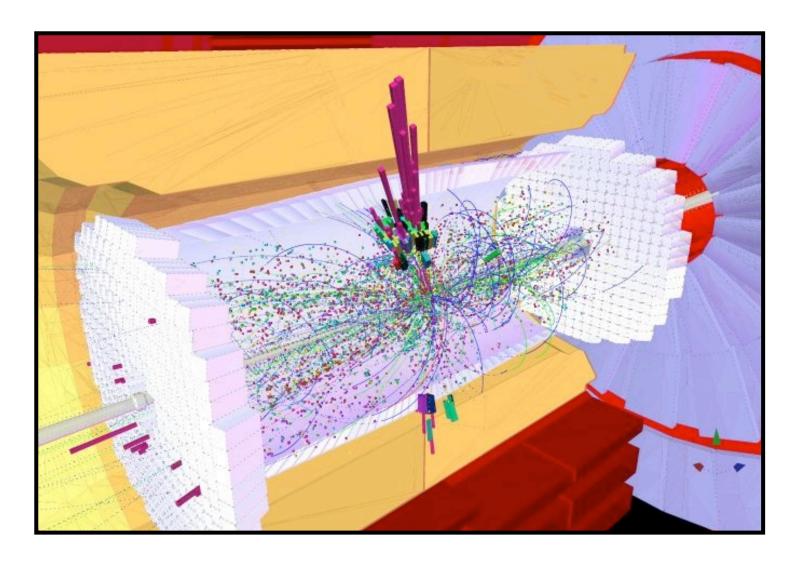
So the acrobat has one dimension, and the flea has two dimensions, but one of these dimensions is a small closed loop.

The acrobat can only detect the one dimension of the rope, just as we can only see the world in three dimensions, even though it might well have more.

This is impossible to visualize, precisely because we can only visualize things in three dimensions!



A hypothesized new particle Z'

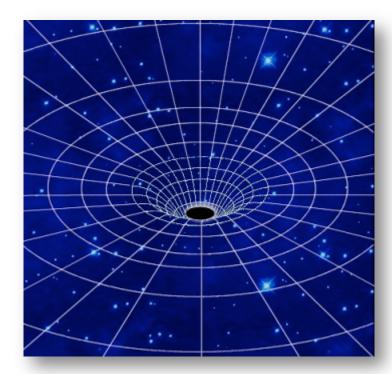


Microscopic Black Holes

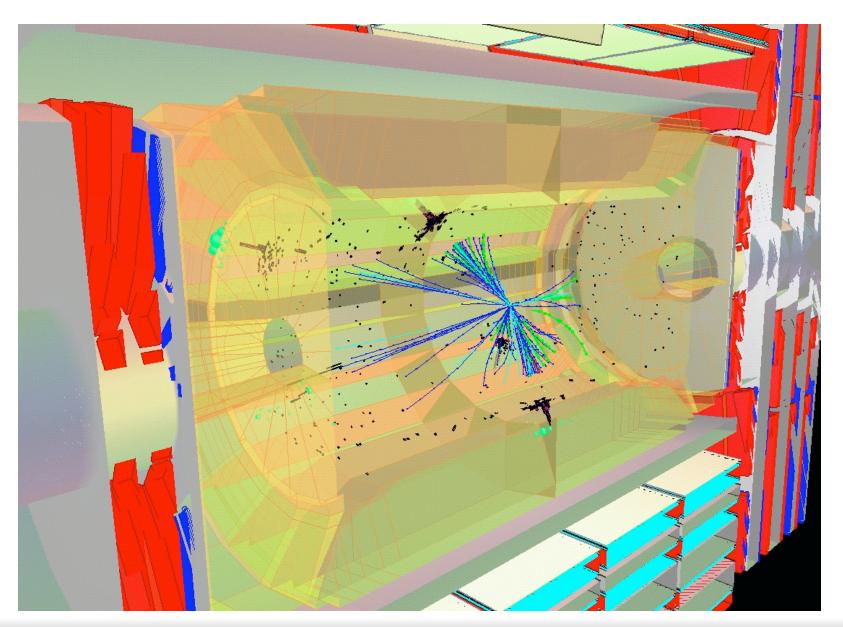
According to some theoretical models, tiny black holes could be produced in collisions at the LHC.

They would then very quickly decay and be detected by experiments (the tinier the black hole, the faster it evaporates).





Black Hole



Summary

- LHC will provide proton-proton collisions at 10 TeV soon
- LHC will provide access to conditions not seen since the early Universe
 - Analysis of LHC data has potential to change how we view the world
 - But LHC analysis will require finesse and care
- LHC is likely to provide answers to many key questions in Physics
- All expectations are based on a detailed understanding of particles and forces within the SM that we have garnered over the past several decades
- With the expected clarifications, we will address some very major issues pertaining to the universe, for example:
 - What is cosmic dark matter? Is it made of SUSY neutralinos?
 - Is the universe filled with a Higgs field? Is it related to dark energy?
 - What is the structure of space-time? Are there extra space dimensions?
- CMS is eagerly awaiting data

Exciting times ahead – stay tuned!

The Unknowns!

