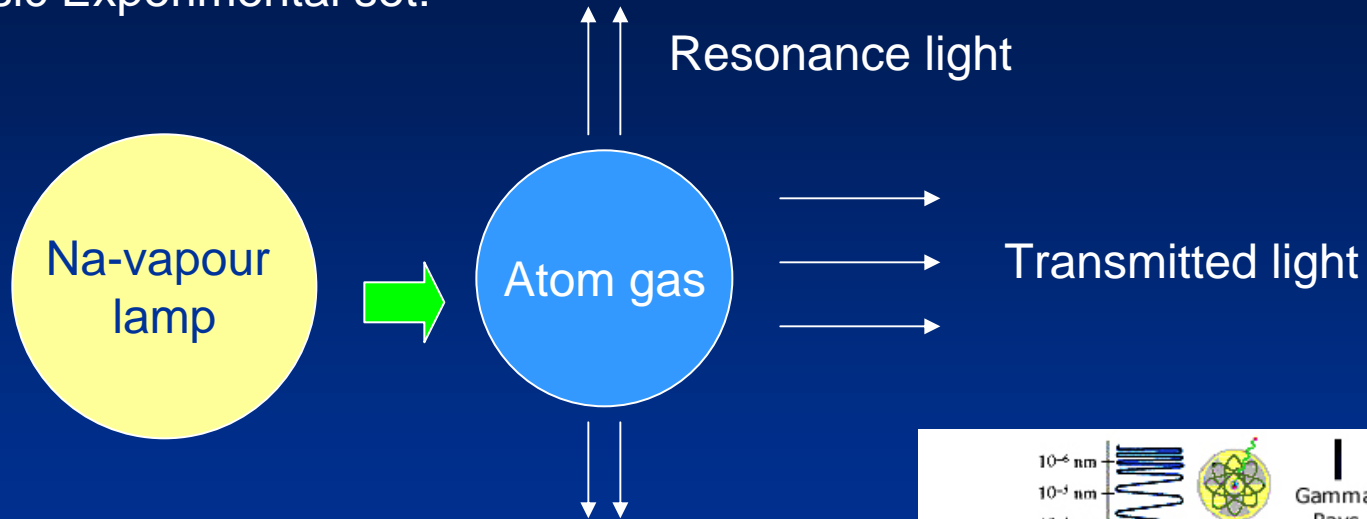


# Atom and Quanta: Bohr's Theory

Phys 460, Fall 2009, JP Hu

# Optical Spectrum of Atom

- Basic Experimental set:

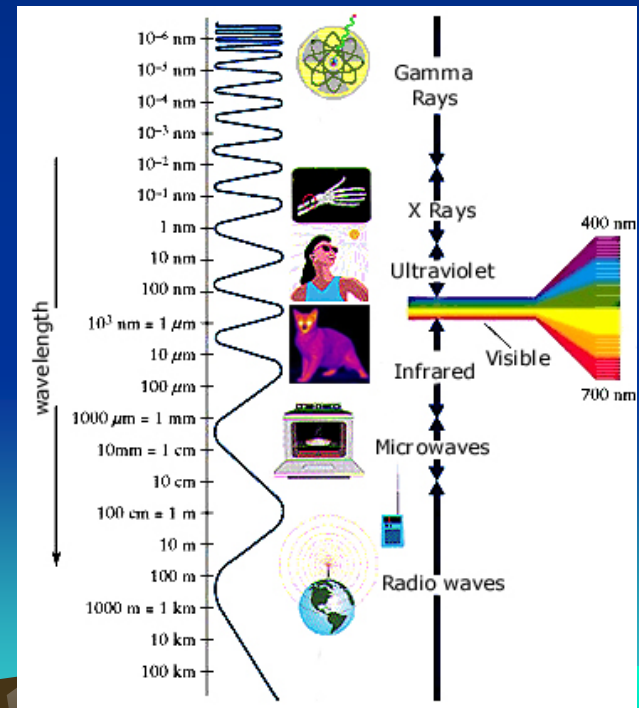


- Basic knowledge of electromagnetic spectrum

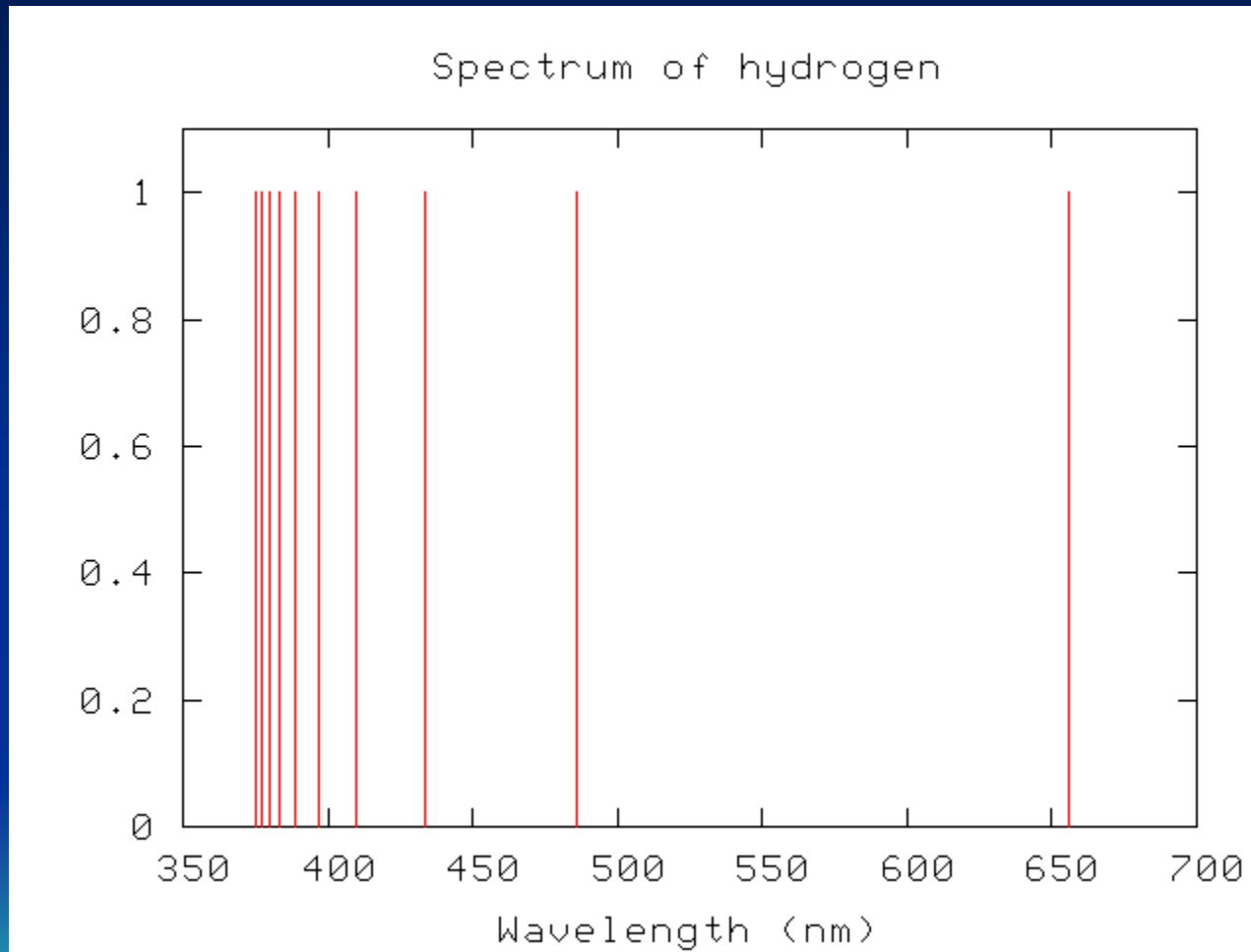
Wavelength:  $\lambda$

Wavenumber:  $\nu=1/\lambda$

Frequency:  $\omega=c/\lambda$



# Optical Spectra of Hydrogen atom



- Discrete numbers: 656.3, 486.1, 434(nm),...

# Balmer's formula (1885)

$$\lambda = \left( \frac{n^2}{n^2 - 4} \right) G$$

$$\nu = 1/\lambda = R_H \left( \frac{1}{2^2} - \frac{1}{n^2} \right)$$

$$R_H = 109677.5810 \text{ cm}^{-1}, n = 3, 4, \dots$$

$R_H$ : Rydberg constant

- How good is the formula?

Example: 15233.21(exp) vs 15233.00 (th) for  $n=3$

20564.77(exp) vs 20564.55 (th) for  $n=4$

# More spectra

$$\lambda = \left( \frac{n^2}{n^2 - n'^2} \right) G$$

$$\nu = 1/\lambda = R_H \left( \frac{1}{n'^2} - \frac{1}{n^2} \right)$$

$$R_H = 109677.5810 \text{ cm}^{-1}$$

$$n' = 1, 2, \dots, n = n' + 1, \dots$$

$n'$ ,  $n$ : principal quantum numbers

1906 Lyman, 1908, Paschen, Brackett 1922

# Bohr's Solution

- Classical orbits:

$$mr\omega^2 = \frac{e^2}{4\pi\epsilon_0} \frac{1}{r^2}$$

- Energy of electron:

$$E = \frac{1}{2}mr\omega^2 - \frac{e^2}{4\pi\epsilon_0} \frac{1}{r} = -\frac{e^{4/3}}{2(4\pi\epsilon_0)^{2/3}} (m\omega^2)^{1/3}$$

- Bohr's bold postulates:

- The classical motion of electron in atom is still valid. However, only discrete orbits with certain energy  $E_n$  is allowed.
- The motion of the electrons in these quantized orbits is radiationless.
- The light is emitted or absorbed when the electron transfers from one orbit to the other.

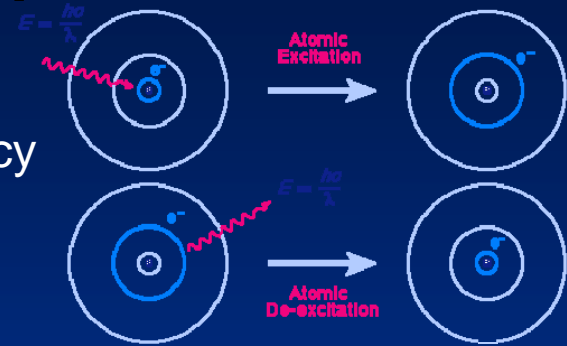
$$E_n = -Rhc / n^2$$

$$E_n - E_{n'} = h\Omega$$

- With increasing orbital radius  $r$ , the law should become identical to classical physics ---- Bohr's Correspondence Principle

# Bohr solution

- Classical:  
light frequency = classical electron orbiting frequency
- From Bohr correspondence principle, we have



$$E_n - E_{n-1} \rightarrow 2Rhc / n^3 = h\Omega = h\omega$$

$$E_n = -Rhc / n^2 = -\frac{e^{4/3}}{2(4\pi\epsilon_0)^{2/3}} (m(2Rc / n^3)^2)^{1/3}$$

$$R = me^4 / (8\epsilon_0^2 h^3 c) = 109737.318 \text{ cm}^{-1}$$

- Quantization of angular momentum:

$$2Rhc / n^3 = h\omega$$

$$l = m\omega^2 r = n\hbar$$

# Beyond Bohr's Theory

- Sommerfeld's extension of the Bohr Model  
( Relativistic mass change)

$$E_{n,k} = \frac{-Rhc}{n^2} \left[ 1 + \frac{\alpha^2 z^2}{n^2} (n/k - 3/4) + \dots \right]$$
$$\alpha = 1/137 = \frac{e^2}{2\epsilon_0 hc}$$

- Semiclassical quantization rule:

$$\oint pdq = n\hbar$$

# De Broglie's Particle-wave

- Classical waves have certain modes with fixed boundary

- From Plank-Einstein: a light with frequency  $\omega$

$$E = h\omega$$

$$p = h\omega / c = h / \lambda$$

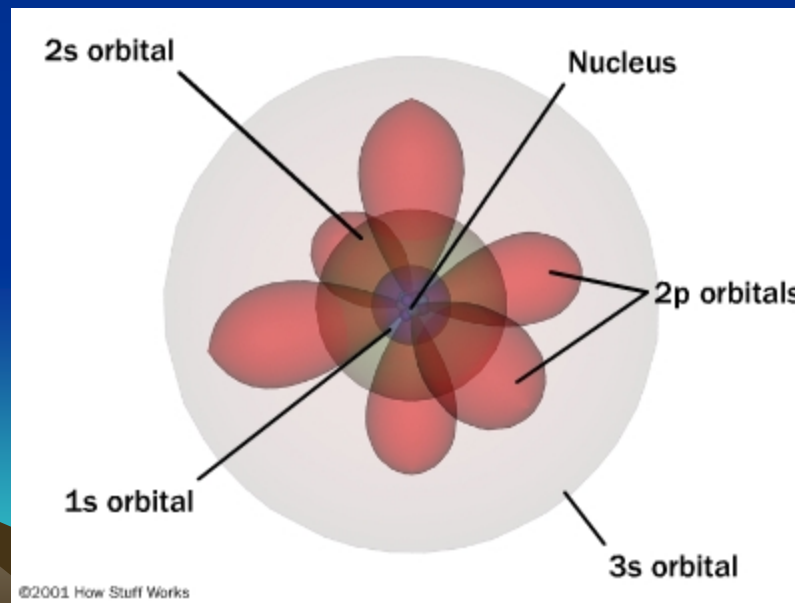
- Consider a particle as wave with wavelength defined as

$$\lambda = h / p$$



# De Broglie's interpretation of Bohr's orbit

$$2\pi r = n\lambda = nh / p$$
$$pr = l = \hbar n$$



# Limitation of Bohr's Theory

- It is a theory with conjectures
- Lack of real calculation power
- Limit to hydrogen-type atoms
- Do not know how to extend it to more complicated system

Anyway, Bohr's theory fundamentally changes our view of world.

- System is characterized by states.
- The physics is determined by final and initial states.
- Energy is quantized.

