

# Chap 1 Early Atomic Physics

This chapter we will go through some basic concepts and get rough sizes that are important for different atomic processes.

## Atomic Units Natural units for atoms

Charge  $e = 1.602 \times 10^{-19} \text{ C}$

Mass  $m_e = 9.109 \times 10^{-31} \text{ kg}$

Ang. mom.  $\hbar = 1.055 \times 10^{-34} \text{ kg m}^2/\text{s}$

Length  $a_0 = \hbar^2 / [ \frac{e^2}{4\pi\epsilon_0} m_e ] = 0.5292 \times 10^{-10} \text{ m}$

Speed  $v_0 = \hbar / (m_e a_0) = 2.188 \times 10^6 \text{ m/s}$

Time  $t_0 = a_0 / v_0 = 24.19 \times 10^{-18} \text{ s}$

Energy  $E_0 = e^2 / 4\pi\epsilon_0 a_0 = 4.360 \times 10^{-18} \text{ J} = 27.211 \text{ eV} \quad ( = \frac{\hbar^2}{m_e a_0^2} )$

Elec. Dip. Mom  $p_0 = e a_0 = 8.478 \times 10^{-30} \text{ Cm} \quad ( 1 \text{ Debye} = 3.336 \times 10^{-30} \text{ Cm} = 0.39 p_0 )$

Mag. Dip Mom  $\mu_0 = \frac{e}{t_0} a_0^2 = 1.855 \times 10^{-23} \text{ J/T} = 2 \mu_B$

Fine structure constant  $\alpha = \frac{v_0}{c} = \frac{1}{137.04}$

Bohr H atom - Electron moves in a circle around proton with  $m_e v r = n \hbar$  Gives quantized energies, radii, speeds, ...

$$E_n = -\frac{1}{2} \frac{E_0}{n^2}$$

$$r_n = a_0 n^2$$

$$v_n = v_0 / n$$

Note bunching of energy levels at high  $n$

Note the increasing size at high  $n$

Note the slower speed at high  $n$

Typical atoms have almost no interaction at distances larger than few nm. Highly excited atoms have noticeable interaction

Proposal to entangle atomic states using highly excited atoms through electric dipole-dipole interaction.



In terms of frequency  $\frac{E}{\hbar} \sim \frac{p^2}{4\pi\epsilon_0 R^3} = \frac{E_0}{\hbar} \left(\frac{a_0}{R}\right)^3 n^4 \approx 1 \text{ kHz} \left(\frac{n}{R/\text{nm}}\right)^3$

$R = 3 \mu\text{m}, n = 50 \quad E/\hbar \sim 200 \text{ MHz}$

More complicated because not permanent dipole moment

You can manipulate highly excited states with microwaves

$$\Delta E/\hbar = \frac{E_0}{\hbar} \left(\frac{1}{50^2} - \frac{1}{51^2}\right) = 102 \text{ GHz}$$

The energies can be shifted with small E-fields

$$F \sim \frac{E_0}{a_0 e} \frac{1}{n^4} \sim 5.14 \times 10^{11} \text{ V/m} \frac{1}{n^4} = 5.14 \times 10^9 \text{ V/cm} / n^4$$

At  $n = 100 \quad F \sim 51 \text{ V/cm} \sim \text{destroys atom}$

If there is a higher charge  $E_n \propto Z^2/n^2, r_n \propto n^2/Z, v_n \propto Z/n$   
X-Rays!

If the electron  $\rightarrow$  muon  $r \propto n^2 \frac{m_e}{m_\mu}, v \propto \frac{1}{n}, E_n \propto \frac{1}{n^2} \frac{m_\mu}{m_e}$   
muon catalyzed fusion, exotic atoms

For a charge interacting with electric and magnetic fields, the Hamiltonian is

$$H = \frac{(\vec{p} - q\vec{A})^2}{2m} + q\phi \quad \text{where} \quad \vec{E} = -\vec{\nabla}\phi - \frac{\partial\vec{A}}{\partial t} \quad \vec{B} = \vec{\nabla} \times \vec{A}$$

Show that Hamilton's equations give the correct equations of motion

$$\dot{x} = \frac{\partial H}{\partial p_x} = \frac{p_x - qA_x}{m} \Rightarrow p_x = m\dot{x} + qA_x = m v_x + qA_x(x, y, z, t)$$

$$\begin{aligned} \dot{p}_x &= -\frac{\partial H}{\partial x} = -q \frac{\partial \phi}{\partial x} + \frac{q}{m} \left[ \frac{\partial A_x}{\partial x} (p_x - qA_x) + \frac{\partial A_y}{\partial x} (p_y - qA_y) + \frac{\partial A_z}{\partial x} (p_z - qA_z) \right] \\ &= -q \frac{\partial \phi}{\partial x} + q \left( \frac{\partial A_x}{\partial x} v_x + \frac{\partial A_y}{\partial x} v_y + \frac{\partial A_z}{\partial x} v_z \right) \\ &= m\dot{v}_x + q \frac{\partial A_x}{\partial t} + q v_x \frac{\partial A_x}{\partial x} + q v_y \frac{\partial A_x}{\partial y} + q v_z \frac{\partial A_x}{\partial z} \end{aligned}$$

Shuffle

$$\begin{aligned} m\dot{v}_x &= -q \left( \frac{\partial \phi}{\partial x} + \frac{\partial A_x}{\partial t} \right) + q v_y \left( \frac{\partial A_y}{\partial x} - \frac{\partial A_x}{\partial y} \right) - q v_z \left( \frac{\partial A_x}{\partial z} - \frac{\partial A_z}{\partial x} \right) \\ &= q \left[ E_x + v_y B_z - v_z B_y \right] \quad \checkmark \end{aligned}$$

To understand the implications, let's look at two examples  
 1) a hydrogen atom in a constant, uniform B-field  
 2) a hydrogen atom in a plane light wave

Example 1)  $\phi(\vec{r}) = \frac{e}{4\pi\epsilon_0 r}$ ,  $\vec{A} = \frac{B}{2}(-y, x, 0) \Rightarrow \vec{B} = (0, 0, B)$

$$\begin{aligned} H &= \frac{1}{2m_e} \left[ \left( p_x - \frac{eB}{2} y \right)^2 + \left( p_y + \frac{eB}{2} x \right)^2 + p_z^2 \right] - \frac{e^2}{4\pi\epsilon_0 r} \\ &= \underbrace{\frac{p^2}{2m_e} - \frac{e^2}{4\pi\epsilon_0 r}}_{H_{\text{hyd}} (B=0)} + \underbrace{\frac{eB}{2m_e} (x p_y - y p_x)}_{\text{paramagnetic}} + \underbrace{\frac{e^2 B^2}{8m_e} (x^2 + y^2)}_{\text{diamagnetic}} \end{aligned}$$

Figure out what each term is doing

$$\frac{eB}{2m_e} (x p_y - y p_x) = \frac{e\hbar}{2m_e} \frac{L_z}{\hbar} B = \mu_B \overset{\rightarrow \mu_0/2}{\text{Cazimithal}} \overset{\text{quan. \#}}{m} B \quad \text{Zeeman}$$

Typical size  $\frac{\mu_B}{\hbar} \approx 2\pi \cdot 14 \text{ GHz/T} = 2\pi \cdot 14 \text{ MHz} / (10 \text{ Gauss})$

Typical size  $\frac{e^2 B^2}{8m_e} (x^2 + y^2) \sim \frac{e^2}{8m_e} B^2 a_0^2 n^4 \sim 10^{-29} \text{ J} \left( \frac{B}{1 \text{ T}} \right)^2 n^4$

Convert to frequency by  $1/\hbar \sim 15 \text{ kHz} \left( \frac{B}{1 \text{ T}} \right)^2 n^4 \Rightarrow$  (high n are easily perturbed)

In more complicated atoms, the Zeeman splittings can have substantial contribution from nucleus giving smaller result.

Example 2)  $\varphi = \frac{e}{4\pi\epsilon_0 r}$      $\vec{A} = -\frac{E_m}{\omega} (0, 0, \cos(kx - \omega t))$

Propagation direction of plane wave? Polarization?

Find  $\vec{E}$  &  $\vec{B}$  plane wave

$$\vec{E} = -\frac{\partial \vec{A}}{\partial t} = E_m (0, 0, \sin(kx - \omega t))$$

$$\vec{B} = -\frac{E_m k}{\omega} (0, \sin(kx - \omega t), 0)$$

$$\vec{E} \times \vec{B} = \frac{E_m^2}{c} (\sin^2(kx - \omega t), 0, 0)$$

If  $E_m \ll E$ -field in atom

$$H \approx \frac{p^2}{2m} - \frac{e^2}{4\pi\epsilon_0 r} - \frac{E_m e}{m\omega} p_z \cos(kx - \omega t)$$

Velocity gauge

In many situations, the wavelength of the light is much larger than the size of the atom

$$\cos(kx - \omega t) = \underbrace{\cos(\omega t)}_{\text{dipole}} - \underbrace{kx \sin(\omega t)}_{\text{quadrupole}} \dots$$

This is why dipole transition (example  $1s \rightarrow 3p$ ) is typically much stronger than quadrupole transition (example  $1s \rightarrow 3d$ ).

Gauge transformation can't change physics  $\vec{A} \rightarrow \vec{A} + \vec{\nabla} f$

$$\varphi \rightarrow \varphi - \frac{\partial f}{\partial t}$$

$$\vec{B} = \vec{\nabla} \times (\vec{A} + \vec{\nabla} f) = \vec{\nabla} \times \vec{A} \quad \checkmark$$

$$\vec{E} = -\vec{\nabla} \varphi - \frac{\partial \vec{A}}{\partial t} = -\vec{\nabla} \varphi + \vec{\nabla} \frac{\partial f}{\partial t} - \frac{\partial \vec{A}}{\partial t} - \frac{\partial}{\partial t} \vec{\nabla} f = -\vec{\nabla} \varphi - \frac{\partial \vec{A}}{\partial t} \quad \checkmark$$

Choose  $f = \frac{E_m z}{\omega} \cos(\omega t) \rightarrow \vec{A} = -\frac{E_m}{\omega} (0, 0, \cos(\omega t - kx) - \cos(\omega t))$

$$\varphi = E_m z \sin(\omega t)$$

For  $\lambda \gg$  size of atom  $\vec{A} \approx 0$

$$H = \frac{p^2}{2m} - \frac{e^2}{4\pi\epsilon_0 r} - e E_m z \sin(\omega t)$$

Length gauge

Remember the approximation that goes into it!

How to know when the laser is strong? The intensity of light (units  $W/m^2$ ) and frequency play role

$$I = \frac{1}{2} \epsilon_0 E_m^2 c$$

$$E_m = \sqrt{\frac{2I}{\epsilon_0 c}}$$

S Ghimire et al Nat Phys 7 138 (2011) had  $E_m = 0.6 \text{ V/\AA}$   
 $I = \frac{1}{2} 8.85 \times 10^{-12} (0.6 \times 10^{10})^2 3 \times 10^8 \text{ W/m}^2 = 4.8 \times 10^{16} \text{ W/m}^2 = 4.8 \times 10^{12} \text{ W/cm}^2$   
(nowhere near best I)

$$\lambda = 3.25 \text{ \mu m} \Rightarrow f = \frac{c}{\lambda} = 9.2 \times 10^{13} \text{ Hz} \Rightarrow \omega = 2\pi f = 5.8 \times 10^{14} \text{ s}^{-1}$$

Important concept is quiver energy and excursion amplitude

$$\ddot{x} = \frac{e E_m}{m_e} \cos(\omega t) \Rightarrow \dot{x} = \frac{e E_m}{m_e \omega} \sin(\omega t) \Rightarrow x = -\frac{e E_m}{m_e \omega^2} \cos(\omega t)$$

$$\text{Excursion amplitude} = \frac{e E_m}{m_e \omega^2} = 3.1 \times 10^{-9} \text{ m} = 3.1 \text{ nm}$$

$$\text{Quiver energy} = \frac{m_e}{2} \left( \frac{e E_m}{m_e \omega} \right)^2 \overset{\text{avg of } \sin^2}{=} \frac{1}{2} \left( \frac{e E_m}{m_e \omega} \right)^2 = 7.5 \times 10^{-19} \text{ J} = 4.7 \text{ eV}$$

Ponderomotive energy, Ponderomotive potential

$$\text{The energy of 1 photon } E = \frac{hc}{\lambda} = 6.12 \times 10^{-20} \text{ J} = 0.38 \text{ eV}$$

At this photon energy, 10+ photons are needed to ionize depending on the atom type.

Example 1: Laser wakefield acceleration

Short, high intensity laser pulse pushes electrons out of region moving at  $v \approx c$ . Electrons pulled into wake can accelerate to GeV+ energies. Why electrons pushed out?

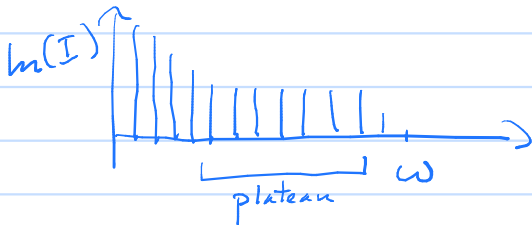
## Example 2: High harmonic generation

At high intensities, the material will emit photons at frequencies  $1\omega, 3\omega, 5\omega, \dots$  (for symmetric potentials)

Why only odd for symmetric potentials?

$$\ddot{x} = -\omega^2 x - c x^3 + \alpha \sin(\omega t)$$

$$x = x_1 \sin(\omega t) + x_3 \sin(3\omega t) + x_5 \sin(5\omega t) \dots$$



$$\text{Cutoff} \sim I_p + 3.17 U_p$$

↳ ponderomotive energy

3 step model: 1) Tunnel out, 2) Free motion, 3) return to nucleus + emit photon.

After tunnel 
$$V(t) = U_0 \sin(\omega t) + U_0 = U_0 \left[ \sin(\omega t) + \frac{U_0}{U_0} \right]$$

$$X(t) = -\frac{U_0}{\omega} \cos(\omega t) + \frac{U_0}{\omega} + U_0 t = \frac{U_0}{\omega} \left[ 1 - \cos(\omega t) + \frac{U_0}{U_0} \omega t \right]$$

Find  $V(t)$  when  $X(t) = 0$ . Maximum is when  $U_0/U_0 \approx -0.28$

$$V(t) \approx 1.265 U_0 \Rightarrow KE \approx 3.2 U_p \quad \text{Why add } I_p \text{ to this?}$$

## Example 3: Above threshold ionization (ATI)

When laser intensity is high, atoms/molecules/solids can absorb many more photons than needed to ionize the atom.

Hydrogen atom + 1 eV photons  $\Rightarrow$  ionization with 14, 15, 16... photons

$$\text{Electron energy} = n h f - I_p > 0 \quad n > I_p / h f$$